

# Petrogenesis of plagiogranites in the Muslim Bagh Ophiolite, Pakistan: implications for the generation of Archaean continental crust

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**Abstract** – High-SiO<sub>2</sub> rocks referred to as oceanic plagiogranites are common within the crustal sequences of ophiolites; however, their mode of petrogenesis is controversial with both late-stage fractional crystallization and partial melting models being proposed. Here, we present new whole-rock data from plagiogranitic dyke-like bodies and lenses from the lower and middle sections of the sheeted dyke complex of the Cretaceous Muslim Bagh Ophiolite, northwestern Pakistan. The plagiogranites have similar geochemical signatures that are inconsistent with them being the fractionation products of the mafic units of the Muslim Bagh Ophiolite. However, the plagiogranites all display very low TiO<sub>2</sub> contents (<0.4 wt%), implying that they formed by partial melting of mafic rocks. Melt modelling of a crustal gabbro from the Muslim Bagh Ophiolite shows that the trace-element signature of the plagiogranites can be replicated by 5–10% melting of a crustal hornblende gabbro with amphibole as a residual phase, resulting in a concave-up middle rare Earth element pattern. Compositional similarities between the Muslim Bagh Ophiolite plagiogranites and Archaean TTG (trondhjemite–tonalite–granodiorite) has implications for the generation of juvenile Archaean continental crust. As the Muslim Bagh Ophiolite was derived in a supra-subduction zone, it is suggested that some Archaean TTG may have been derived from melting of mafic upper crust in early subduction-like settings. However, due to the small volume of Muslim Bagh Ophiolite plagiogranites, it is inferred that they can be instructive on the petrogenesis of some, but not all, Archaean TTG.

Keywords: Pakistan, Muslim Bagh, ophiolite, oceanic plagiogranite, partial melting

## 1. Introduction

Within obducted Phanerozoic ophiolite sequences, suites of felsic rocks termed ‘oceanic plagiogranites’ (Coleman & Peterman, 1975; Le Maitre *et al.* 2002, p. 118) occur as small-volume (<10%) components (Coleman & Peterman, 1975; Koepke *et al.* 2007). The petrogenesis of these plagiogranites is controversial, having been variously proposed to have formed by the late-stage crystallization of mafic melts (Coleman & Peterman, 1975), hydrous partial melting (and assimilation) of mafic rocks (Gerlach, Leeman & Ave Lallemand, 1981; Amri, Benoit & Ceuleneer, 1996; Gillis & Coogan, 2002; France, Ildefonse & Koepke, 2009; France *et al.* 2010; Erdmann *et al.* 2015) or silicate–liquid immiscibility (Dixon & Rutherford, 1979).

Significantly, plagiogranites have compositional similarities to trondhjemite, tonalite and granodiorite (TTG) rocks that are common in Archaean terranes from the period 4.0–2.5 Ga (e.g. Drummond, Defant & Kepezhinskas, 1996; Kerrich & Polat, 2006; Moya & Martin, 2012; Kusky *et al.* 2013). Although themselves controversial, Archaean TTG are considered, by many, to be generated by the partial melting of mafic igneous

source regions (e.g. Drummond, Defant & Kepezhinskas, 1996; Foley, Tiepolo & Vannucci, 2002; Rapp, Shimizu & Norman, 2003; Martin *et al.* 2005; Moya & Stevens, 2006; Nutman *et al.* 2009; Hastie *et al.* 2015, 2016). Significantly, the compositional similarity of Phanerozoic oceanic plagiogranites to Archaean TTG suggests that, if we can better understand how plagiogranites are formed, it may further our understanding of how primitive continents were formed on the early Earth (Rollinson, 2008, 2009, 2014).

In this paper, we present major- and trace-element data for oceanic plagiogranites sampled from a sheeted dyke complex within the Late Cretaceous (Neo-Tethyan) Muslim Bagh Ophiolite in northwestern Pakistan (Kakar *et al.* 2012). We investigate the composition of these plagiogranitic lenses and dykes in the sheeted dyke complex to determine their petrogenesis. We then discuss the implications of these results for the generation of Archaean continental crust.

## 2. Ophiolites and plagiogranites

Oceanic plagiogranites are found throughout geological time, in strata of both Precambrian (e.g. Samson *et al.* 2004; Kaur & Mehta, 2005) and

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Phanerozoic (e.g. Tilton, Hopson & Wright, 1981; Flagler & Spray, 1991; Rollinson, 2009) age, and are common in the crustal sections of ophiolitic sequences (e.g. Flagler & Spray, 1991; Amri, Benoit & Ceuleneer, 1996; Twining, 1996; Yaliniz, Floyd & Goncuoglu, 2000; Samson *et al.* 2004). Plagiogranites have also been recovered from recent oceanic ridge systems around the world, for example the Southwest Indian (e.g. Dick *et al.* 2000), Central Indian (e.g. Nakamura *et al.* 2007) and Mid-Atlantic ridges (e.g. Aranovich *et al.* 2010; Grimes *et al.* 2011). The morphology of oceanic plagiogranites is complex and they have been documented in a range of intrusive forms from small veins (millimetre- to centimetre-scale; e.g. Dick *et al.* 2000; Nakamura *et al.* 2007), to dykes and inclusions (millimetre- to metre-scale; e.g. Flagler & Spray, 1991; Jafri, Charan & Govil, 1995), to large kilometre-scale plutonic bodies (e.g. Rollinson, 2009).

Oceanic plagiogranites are predominantly composed of sodic plagioclase and quartz, with mafic (usually hornblende and pyroxene) minerals being minor constituents (<10%) and K-feldspar being a rare phase. In addition to the major modal mineralogy, several accessory minerals including zircon, magnetite and ilmenite are also commonly found in oceanic plagiogranites (Coleman & Peterman, 1975; Coleman & Donato, 1979).

In the mid-1970s, plagiogranites were considered to represent the likely silicic end-products of crystallizing basaltic magmas (Coleman & Peterman, 1975; Coleman & Donato, 1979). Although such a crystallization model is still advocated by some authors, who have shown that oceanic plagiogranites fall along the liquid lines of descent of evolving magmas in other ophiolite units (e.g. Jafri, Charan & Govil, 1995; Rao, Rai & Kumar, 2004; Freund *et al.* 2014), the genesis of oceanic plagiogranites is more commonly attributed to the partial melting of mafic igneous source regions (Gerlach, Leeman & Ave Lallemand, 1981; Flagler & Spray, 1991; see Koepke *et al.* 2007 for a review of oceanic plagiogranite petrogenesis models). Melting models propose that oceanic plagiogranites are derived through partial melting of mafic protoliths, either by hydrous partial melting of crustal gabbros (e.g. Gerlach, Leeman & Ave Lallemand, 1981; Flagler & Spray, 1991; Amri, Benoit & Ceuleneer, 1996) or the assimilation and partial melting of hydrothermally altered sheeted dykes (e.g. Gillis & Coogan, 2002; France, Ildfonse & Koepke, 2009; France *et al.* 2010; Erdmann *et al.* 2015).

A partial melting origin is supported by the experimental work of Koepke *et al.* (2004), who undertook hydrous melting experiments on oceanic cumulate gabbros at temperatures from 900 – 1060 °C and a relatively shallow pressure of 0.2 GPa. Koepke *et al.* (2004) showed that lower temperature runs (900 – 940 °C) generated partial melts with similar major element compositions to natural oceanic plagiogranites. One important finding from the P-T experiments was that the melts replicate the low TiO<sub>2</sub> concentrations

that can be found in oceanic plagiogranites (<1 wt%; Koepke *et al.* 2004). Low TiO<sub>2</sub> is now considered a key characteristic of oceanic plagiogranites that have been derived by partial melting, as opposed to oceanic plagiogranites derived through fractional crystallization that display higher TiO<sub>2</sub> contents (>1 wt%; Koepke *et al.* 2004, 2007). Further experimental work conducted by France *et al.* (2010) has also shown that oceanic plagiogranites derived by partial melting have low TiO<sub>2</sub> contents, supporting the experimental work of Koepke *et al.* (2004).

### 3. Geological setting

#### 3.a. Regional setting

The Muslim Bagh Ophiolite (MBO) is one of a number of ophiolites (i.e. Bela, Waziristan, Khost, Zhob) of Neo-Tethyan origin (Kakar *et al.* 2014) that comprise the Western Ophiolite Belt of the Zhob Valley, northwestern Pakistan (Ahmad & Abbas, 1979; Mahmood *et al.* 1995; Gnos, Immenhauser & Peters, 1997) (Fig. 1). These ophiolites represent fragments of Neo-Tethyan Ocean crust that were obducted onto the margin of the Indian continent prior to its final collision with Asia (e.g. Gnos, Immenhauser & Peters, 1997; Khan *et al.* 2009); they therefore mark the boundary between the Indian and Eurasian plates (Asrarullah, Ahmad & Abbas, 1979; Mengal *et al.* 1994; Gnos, Immenhauser & Peters, 1997).

The Muslim Bagh area comprises four main geological units (Fig. 1). These units are (south to north) the Indian Passive Margin, the Bagh Complex, the MBO and the Flysch Belt (Mengal *et al.* 1994; Kakar *et al.* 2014). Triassic–Paleocene sediments of the Indian Passive Margin (Kakar *et al.* 2014) are overthrust by the Mesozoic Bagh Complex along the Gawal Bagh thrust (Mengal *et al.* 1994). The Bagh Complex comprises a series of thrust-bounded units including a melange unit, two volcanic units (basalt-chert unit (Bbc), hyaloclastite-mudstone unit (Bhm)) and a sedimentary unit (Bs); see Mengal *et al.* (1994) for detailed descriptions of each unit. Thrusted over the Bagh Complex is the MBO (Kakar *et al.* 2014), described in more detail in the following section. The uppermost unit is the Eocene–Holocene Flysch Belt that rests unconformably on top of the MBO and Bagh Complex in the Katawaz Basin (Mengal *et al.* 1994; Qayyum, Niem & Lawrence, 1996; Kasi *et al.* 2012). The Flysch Belt can be broadly divided into four thrust-bounded formations (Nisai, Khujak, Multana and Bostan formations) comprising fluvial and deltaic successions (Qayyum, Niem & Lawrence, 1996; Kasi *et al.* 2012).

#### 3.b. Muslim Bagh Ophiolite

The MBO is exposed as two massifs: the Saplai Tor Ghar and Jang Tor Ghar massifs (Ahmad & Abbas, 1979; Mahmood *et al.* 1995; Gnos, Immenhauser &

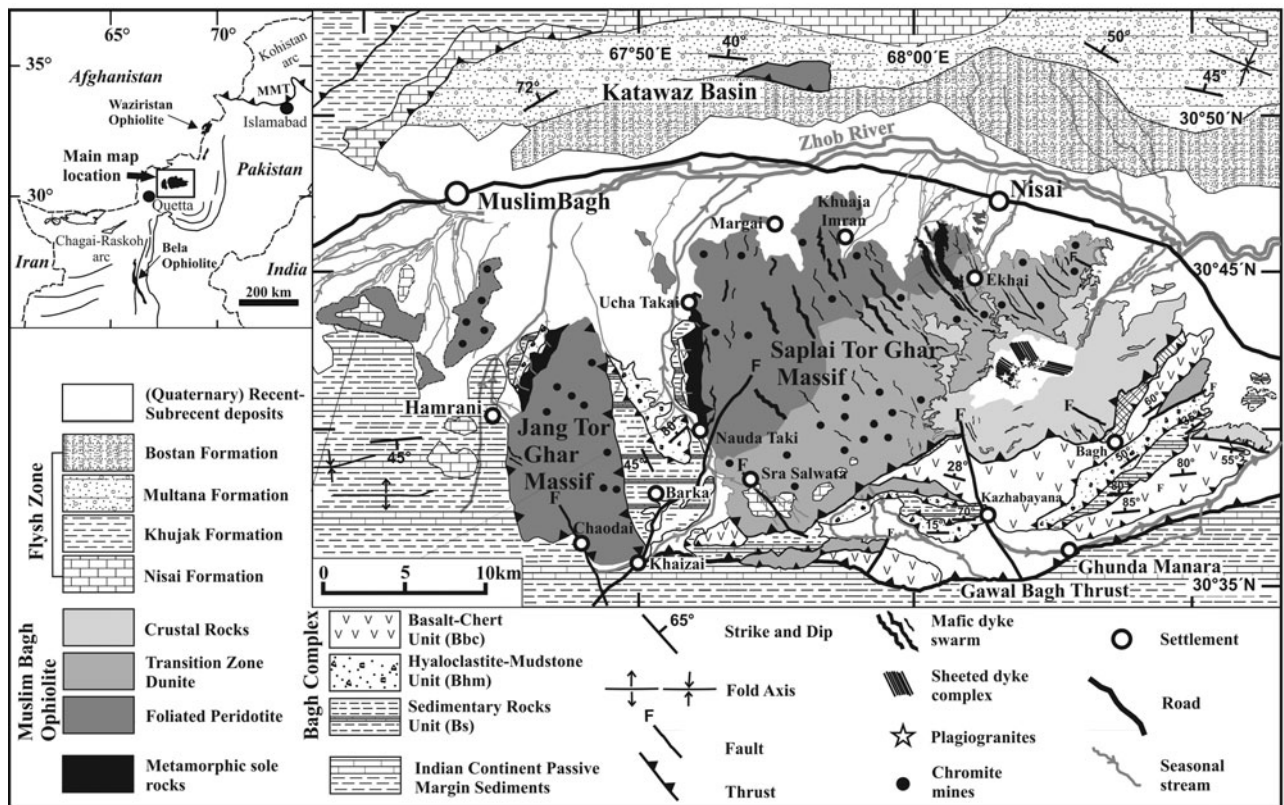


Figure 1. Geological map of the Muslim Bagh area. Inset highlights the location of the Muslim Bagh Ophiolite in northwestern Pakistan (modified from Kakar *et al.* 2014).

Peters, 1997) (Fig. 1). The tectonic setting of formation of the MBO has been variously interpreted as a mid-ocean ridge (Mahmood *et al.* 1995), a back-arc basin (Siddiqui *et al.* 1996) or an island arc (Khan, Kerr & Mahmood, 2007). Most recently however, Kakar *et al.* (2014) have presented evidence that the MBO formed above a slow-spreading supra-subduction zone, based on both the structure of the ophiolite and its arc-like geochemistry. Recent U–Pb dating of zircons in MBO plagiogranites by Kakar *et al.* (2012) gave a crystallization age of  $80.2 \pm 1.5$  Ma, similar to the *c.* 82–81 Ma K–Ar ages obtained by Sawada *et al.* (1995). Dating of amphiboles from the sub-ophiolitic metamorphic sole have yielded K–Ar and plateau Ar/Ar ages of  $80.5 \pm 5.3$  Ma (Sawada *et al.* 1995) and  $70.7 \pm 5$  Ma (Mahmood *et al.* 1995), respectively. The younger age of  $70.7 \pm 5$  Ma (Mahmood *et al.* 1995) is interpreted to date the age of emplacement of the MBO which, when taken in conjunction with the crystallization age of the ophiolite, suggests that the ophiolite was obducted soon after formation (e.g. Kakar *et al.* 2014).

The Saplai Tor Ghar Massif displays a near-complete ophiolite sequence (Kakar *et al.* 2014), with only the extrusive basalts absent (Mahmood *et al.* 1995). The Jang Tor Ghar Massif only preserves mantle sequence rocks (i.e. foliated peridotite) of the oceanic lithosphere, however (Mahmood *et al.* 1995; Kakar *et al.* 2014). The mantle sequence of the MBO has been divided into a foliated peridotite section and

mantle–crust transition zone (Kakar *et al.* 2014). The foliated peridotite is located in both massifs, and comprises serpentinized harzburgite with minor dunite and chromite deposits (Mahmood *et al.* 1995; Khan, Kerr & Mahmood, 2007; Kakar *et al.* 2014). Lherzolite is also found in the lower part of the mantle sequence (Kakar *et al.* 2014). The mantle–crust transition zone of the MBO is a dunite-rich zone with minor gabbro, wherlite, pyroxenite and chromite only exposed in the Saplai Tor Ghar Massif (Mahmood *et al.* 1995; Khan, Kerr & Mahmood, 2007; Khan, Mahmood & Casey, 2007; Kakar *et al.* 2014). Chromite bodies of the transition zone are larger than those in the foliated peridotite section of the mantle sequence (Kakar *et al.* 2014).

The oceanic crustal sequence, as exposed in the Saplai Tor Ghar Massif, comprises a 200–1500 m thick ultramafic–mafic cumulate zone (Ahmad & Abbas, 1979; Siddiqui *et al.* 1996) and a 1 km thick, poorly developed sheeted dyke complex (Siddiqui *et al.* 1996; Khan, Kerr & Mahmood, 2007). The ultramafic–mafic cumulate zone displays both single and cyclic sequences grading from basal dunite through pyroxenite to gabbro, with infrequent anorthosite at the top of the cumulate zone (Ahmad & Abbas, 1979; Siddiqui *et al.* 1996; Khan, Kerr & Mahmood, 2007; Kakar *et al.* 2014). Above the cumulate zone, the sheeted dykes are doleritic, dioritic and plagiogranitic in composition, and all display greenschist to amphibolite grade metamorphism (Sawada *et al.* 1995; Kakar *et al.* 2014).



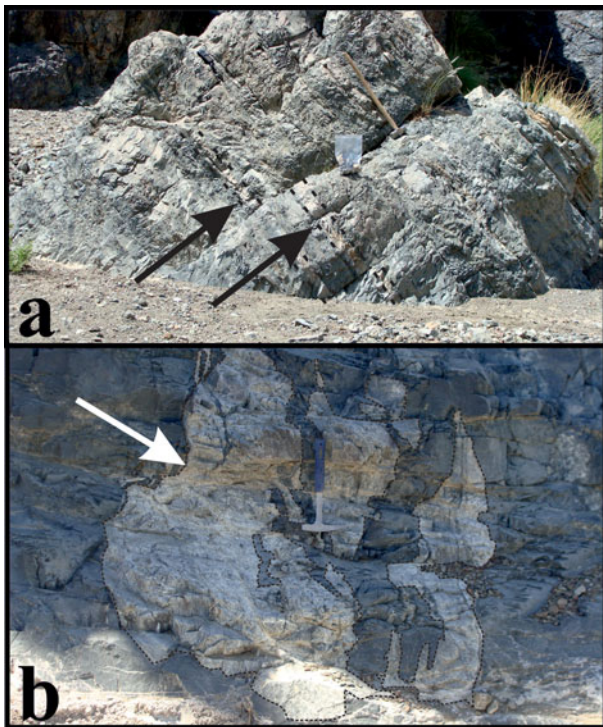


Figure 2. (Colour online) Field photographs of the Muslim Bagh Ophiolite plagiogranites. Plagiogranites are exclusively located within the sheeted dyke complex of the ophiolite crustal sequence, where they take the form of (a) dyke-like bodies and (b) lenses.

Plagiogranites of the MBO are exclusively located at the base and middle portions of the sheeted dyke complex (Mahmood *et al.* 1995; Siddiqui *et al.* 1996). The plagiogranites are rare, comprising <5% by volume of the sheeted dyke complex, and take the form of dykes and small lenses (Fig. 2). They are discontinuous, intrusive bodies, sometimes tapering, displaying a range of sizes. Lenses range from 0.1×0.3 to 1.0×3.0 m, with more dyke-like bodies ranging from 0.3×1.0 to 1.5×3.0 m. The plagiogranites have sharp contacts with the enclosing sheeted dykes, and have also undergone greenschist-amphibolite facies metamorphism with foliated to mylonitized textures (Sawada *et al.* 1995; Siddiqui *et al.* 1996; Kakar *et al.* 2014). Samples for the current study were collected from a range of separate plagiogranite dykes and lenses from across the region. The general sampling locality is shown on Figure 1; more detailed localities and sample information are given in Supplementary Table S1 (available at <http://journals.cambridge.org/geo>).

#### 4. Petrography

The plagiogranites sampled from the MBO for the current study are predominately composed of quartz (*c.* 40 vol.%) and plagioclase (*c.* 50 vol.%), with hornblende and pyroxene comprising minor amounts (<<10 vol.%; hornblende > pyroxene) and zircon and Fe-Ti oxides common as accessory phases. Phenocryst phases of plagioclase, quartz, hornblende and

pyroxene are surrounded by a fine groundmass composed of plagioclase, quartz, hornblende, pyroxene, potassium feldspar (rare) and accessory phases. All phenocryst phases have subhedral to anhedral crystal shapes, with plagioclase displaying simple and albite twinning; hornblende twinning is rare. Throughout the sections, quartz is composed of sub-grains. However, unlike Coleman & Peterman's (1975) original definition of oceanic plagiogranites, the MBO plagiogranites do not display vermicular intergrowths of quartz and plagioclase. Evidence for hydrothermal alteration and low-grade metamorphism includes moderate sericitization of plagioclase crystals (concentrated in the core of crystals; see online Supplementary Figure S1, available at <http://journals.cambridge.org/geo>).

### 5. Geochemical results

#### 5.a. Analytical techniques

Plagiogranite samples were prepared and analysed for major, minor and trace elements at the School of Earth and Ocean Sciences, Cardiff University, Wales, UK. Loss on ignition (LOI) was measured using *c.* 1.5 ± 0.0001 g of sample powder baked at 900 °C in a Vecstar Furnace for 2 hours. Major and minor elements and Sc were measured using a JY-Horiba Ultima 2 inductively coupled plasma optical emission spectrometry (ICP-OES). Minor, trace and the rare Earth elements (REE) were measured using a thermochemical X series (X7) inductively coupled plasma mass spectrometer (ICP-MS) following methods described by McDonald & Viljoen (2006). The accuracy and precision of the data were assessed using the international standard reference materials JB1a, JA2 and JG-3 (obtained analysis, certified values and detection limits for JB1a are shown in Supplementary Table S2, available at <http://journals.cambridge.org/geo>). The full plagiogranite sample dataset is shown in Table 1.

#### 5.b. Element mobility

The altered nature of the plagiogranite samples means that some of the major elements and large-ion lithophile elements (LILE) may have been mobilized relative to the high-field-strength elements (HFSE) and REE (e.g. Hastie *et al.* 2007). Although low LOI values (0.59–2.45 wt%) suggest that the plagiogranites have suffered little alteration, the high proportion of quartz (*c.* 40%) means that the effective LOI of the non-quartz components may double the whole-rock values. However, major element (*v.* LOI) variation plots of the plagiogranite samples show no correlation with LOI, all displaying very low  $R^2$  values (see Supplementary Fig. S2, available at <http://journals.cambridge.org/geo>). With the exception of MgO (<0.52), all major elements display  $R^2$  values of <0.32. These data suggest that the major-element concentrations are not primarily controlled by alteration, and can confidently be used to compare with literature Archaean TTG

Table 1. Major- and trace-element analyses of the Muslim Bagh Ophiolite plagiogranites.

Sample	PI-01	PI-02	PI-03	PI-06	PI-07	PI-13	PI-15	PI-17	PI-19	PI-21	PI-22	PI-23	PI-25
SiO <sub>2</sub> (wt%)	73.08	70.71	70.26	70.00	71.66	72.06	72.40	74.06	74.85	74.69	72.56	79.73	76.85
TiO <sub>2</sub>	0.30	0.31	0.30	0.34	0.30	0.31	0.34	0.30	0.29	0.29	0.39	0.10	0.28
Al <sub>2</sub> O <sub>3</sub>	12.98	14.83	14.87	15.50	15.63	15.03	13.37	10.50	13.42	13.79	14.41	12.27	11.68
Fe <sub>2</sub> O <sub>3(T)</sub>	2.81	3.48	3.00	2.47	2.40	2.60	2.84	4.03	3.02	2.79	4.47	1.25	3.87
MnO	0.04	0.05	0.04	0.04	0.04	0.04	0.04	0.03	0.03	0.02	0.04	0.01	0.04
MgO	1.81	1.01	0.78	0.66	0.63	0.42	0.92	1.22	0.50	0.40	0.37	0.10	0.19
CaO	3.68	2.87	3.02	3.94	3.38	4.01	5.67	4.72	4.01	3.93	3.10	1.91	1.86
Na <sub>2</sub> O	3.12	4.42	4.35	4.19	4.34	4.24	3.13	1.70	3.40	3.65	3.98	4.20	4.14
K <sub>2</sub> O	0.52	0.98	0.64	0.62	0.71	0.55	0.43	0.33	0.18	0.17	0.29	1.06	0.12
P <sub>2</sub> O <sub>5</sub>	0.03	0.03	0.04	0.05	0.05	0.05	0.06	0.06	0.07	0.07	0.14	0.02	0.06
LOI	1.83	1.71	1.75	1.16	1.32	0.81	1.09	2.45	0.82	0.76	1.72	0.59	0.81
Total	100.19	100.38	99.05	98.97	100.46	100.12	100.29	99.40	100.05	99.99	100.92	100.53	99.22
Sc (ppm)	6.7	7.8	9.0	10.7	11.8	12.5	13.4	13.0	13.4	9.4	7.3	2.9	10.5
V	59	64	55	45	38	55	45	48	62	44	49	10	8
Cr	13	26	2	13	8	7	80	174	12	21	116	18	10
Co	8.6	14.8	13.5	7.9	5.6	7.1	9.3	13.7	6.8	7.3	6.6	2.0	14.3
Ni	15.7	50.2	22.4	78.5	10.4	8.9	8.6	10.5	25.3	5.0	89.1	10.1	14.2
Ga	10.9	11.3	10.7	11.0	10.4	11.2	11.2	8.8	11.6	12.1	15.8	10.0	12.1
Rb	3.8	6.1	3.4	2.4	2.9	2.8	4.1	8.1	2.0	1.5	2.9	20.7	1.6
Sr	160	163	167	170	144	163	159	196	155	155	214	105	91
Y	13.0	13.6	11.3	17.3	18.0	14.2	15.8	17.7	19.4	16.3	13.0	17.9	22.0
Zr	40.3	198.8	46.6	94.9	42.7	51.2	20.7	35.2	80.0	57.8	283.0	61.1	129.8
Nb	1.28	1.05	1.17	0.79	0.69	0.68	0.80	0.80	0.89	1.42	4.13	5.19	0.70
Cs	0.05	0.07	0.10	0.07	0.09	0.06	0.05	0.12	0.04	0.03	0.07	0.13	0.02
Ba	68	69	71	79	74	71	78	67	60	80	198	502	59
La	5.14	2.87	2.97	3.70	2.12	3.28	2.14	3.16	1.50	1.69	21.83	15.45	4.77
Ce	9.58	6.28	6.08	7.89	5.06	6.47	5.00	6.50	4.26	4.56	37.17	27.28	11.27
Pr	1.15	0.84	0.84	1.09	0.77	0.84	0.75	0.94	0.70	0.71	4.23	3.03	1.70
Nd	4.68	3.71	3.91	5.22	3.99	3.89	3.72	4.52	3.51	3.43	14.97	10.13	7.69
Sm	1.37	1.28	1.21	1.59	1.44	1.35	1.50	1.50	1.43	1.33	2.96	2.15	2.31
Eu	0.53	0.53	0.63	0.68	0.60	0.58	0.63	0.65	0.52	0.54	0.82	0.54	0.88
Gd	1.54	1.38	1.33	1.96	1.96	1.69	1.78	1.91	2.00	1.73	2.86	2.27	2.76
Tb	0.28	0.27	0.24	0.38	0.35	0.30	0.33	0.35	0.42	0.35	0.40	0.37	0.51
Dy	1.88	1.88	1.71	2.53	2.67	2.01	2.35	2.45	2.90	2.34	2.12	2.30	3.09
Ho	0.43	0.45	0.38	0.56	0.60	0.46	0.51	0.53	0.60	0.49	0.40	0.50	0.64
Er	1.41	1.38	1.10	1.74	1.81	1.42	1.61	1.68	1.85	1.46	1.18	1.57	1.84
Tm	0.24	0.26	0.19	0.30	0.31	0.24	0.28	0.27	0.31	0.27	0.19	0.30	0.33
Yb	1.58	1.75	1.22	2.04	2.03	1.46	1.76	1.80	2.07	1.68	1.24	2.15	2.13
Lu	0.24	0.30	0.19	0.32	0.32	0.22	0.26	0.27	0.31	0.27	0.21	0.34	0.33
Hf	1.42	5.81	1.48	2.90	1.37	1.65	0.72	1.15	2.08	1.57	6.72	1.43	3.04
Ta	0.11	0.07	0.14	0.10	0.06	0.06	0.07	0.07	0.06	0.10	0.29	0.47	0.04
Pb	1.34	1.39	2.48	1.88	1.41	1.80	1.27	1.17	1.11	0.64	1.82	3.45	1.62
Th	2.17	0.51	0.25	0.26	0.20	0.20	0.23	0.47	0.45	0.33	11.92	4.95	0.61
U	0.30	0.10	0.11	0.09	0.07	0.06	0.05	0.12	0.09	0.09	0.93	0.95	0.13

Fe<sub>2</sub>O<sub>3(T)</sub>: total iron

data. Additionally, Sr (v. LOI; Supplementary Fig. S3, available at <http://journals.cambridge.org/geo>) also displays a very low  $R^2$  value of <0.45. Consequently, the following discussion focuses on the major elements and HFSE and REE, generally regarded as relatively immobile up to greenschist facies (e.g. Floyd & Winchester, 1975; Pearce & Peate, 1995; Hastie *et al.* 2007, 2009).

### 5.c. Major elements

The plagiogranites display a relatively narrow, high-SiO<sub>2</sub> range (70.8–80.2 wt%, anhydrous values), with most also having relatively high Al<sub>2</sub>O<sub>3</sub> (10.7–15.8 wt%) and Na<sub>2</sub>O (1.7–4.5 wt%) contents (Fig. 3). Samples have low TiO<sub>2</sub> (<0.4 wt%), MgO (0.1–1.8 wt%) and K<sub>2</sub>O (<1.1 wt%). Al<sub>2</sub>O<sub>3</sub>, MnO (not shown), MgO and K<sub>2</sub>O decrease with increasing SiO<sub>2</sub>, while other oxides such as TiO<sub>2</sub>, Na<sub>2</sub>O, Fe<sub>2</sub>O<sub>3(T)</sub> and CaO show little to no correlation (Fig. 3). Further, the

plagiogranites do not fall on clear liquid lines of descent along with the gabbros and sheeted dykes of the MBO. On a normative ternary An–Ab–Or plot, the plagiogranites classify as tonalites and trondhjemites (Fig. 4).

The major-element abundances of the plagiogranites are very similar to those of Archaean TTG (Condie, 2005; Martin *et al.* 2005; Moyen & Martin, 2012), with TTG compositions consistently plotting at the lower SiO<sub>2</sub> end of the plagiogranite compositions (Fig. 3). However, this similarity is not observed in K<sub>2</sub>O contents, with TTG generally having much higher K<sub>2</sub>O contents (1.65–2.22 wt%) compared with the MBO plagiogranites (<1.1 wt%).

### 5.d. Trace elements

The plagiogranites show no convincing intra-formation fractionation trends on trace-element variation plots (Fig. 5). This is not surprising, considering

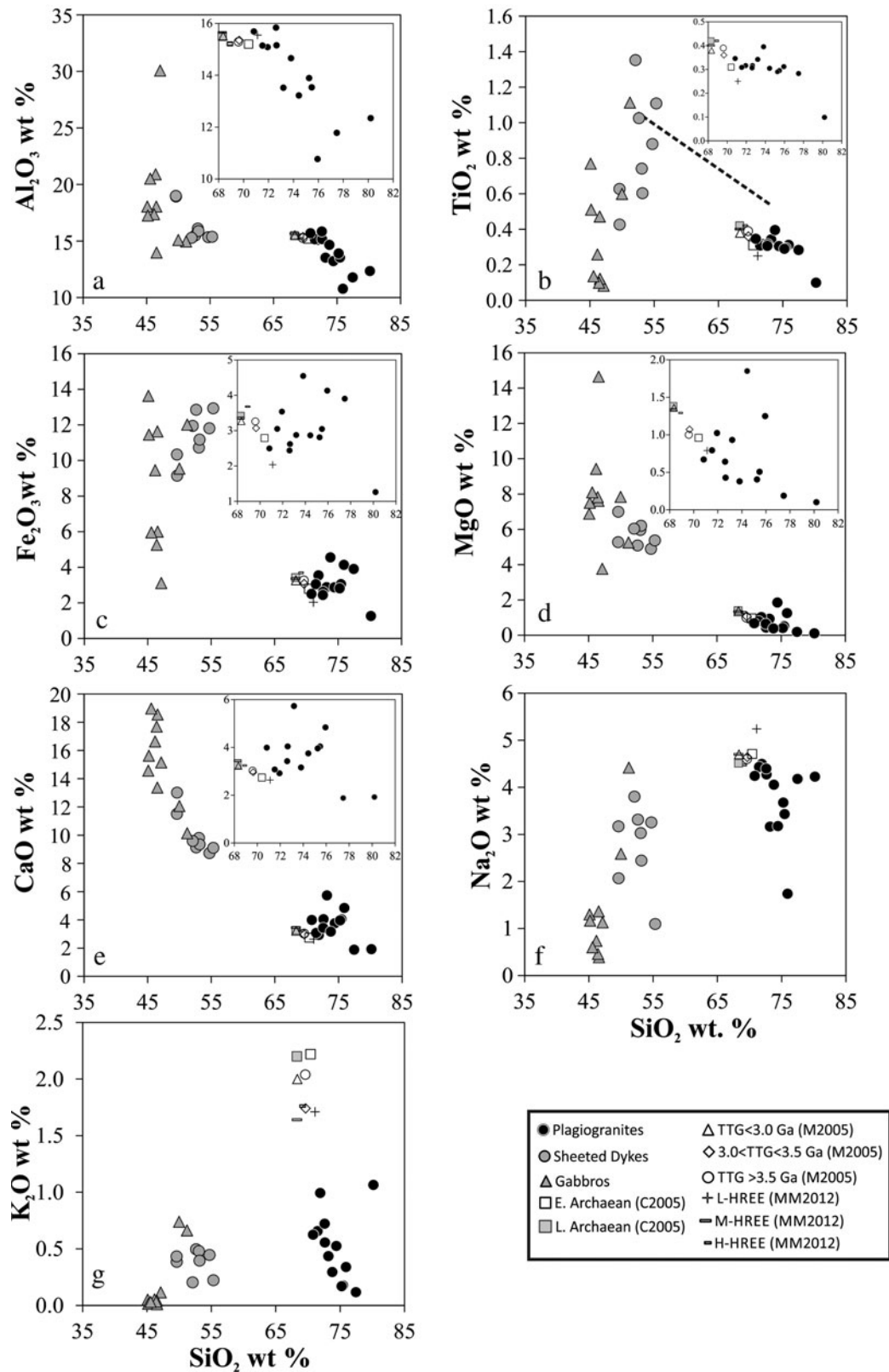


Figure 3. (a–g) Major-element variation plots (v. SiO<sub>2</sub>) of the Muslim Bagh Ophiolite plagiogranites. Also plotted are the sheeted dykes and gabbros of the crustal section of the Muslim Bagh Ophiolite (data from Kakar *et al.* 2014) and Archaean TTG average compositions; C2005 (Condie, 2005), M2005 (Martin *et al.* 2005) and MM2012 (Moyen & Martin, 2012). The black dashed line in (b) separates plagiogranites derived by hydrous partial melting (below the line) and those plagiogranites derived through differentiation or liquid immiscibility (above the line) (after Koepke *et al.* 2007).



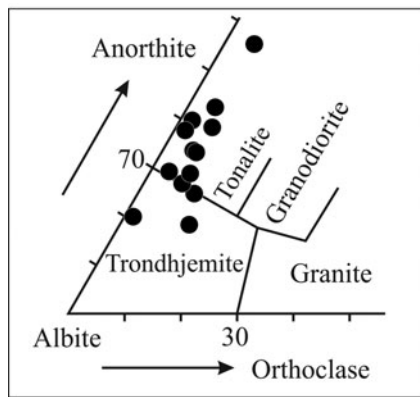


Figure 4. Normative An–Ab–Or ternary plot. Muslim Bagh Ophiolite plagiogranites classify as either tonalites or trondhjemites. Fields from Barker (1979).

that the samples are collected from a diverse range of geographically distinct dykes and lenses. The plagiogranites span a wide range in Zr concentrations (*c.* 20–280 ppm), although the majority of samples fall in the range 20–90 ppm; only three have higher concentrations (130, 199, 283 ppm), suggestive of zircon accumulation (e.g. Rollinson, 2009). In general, the plagiogranites have lower trace-element concentrations than the sheeted dyke complex of the MBO and, with the exception of Sr, have trace-element contents similar to, or slightly greater than, the majority of the gabbros of the crustal section of the ophiolite (Fig. 5). As is the case with the major elements (Fig. 3), the plagiogranites do not fall on clear liquid lines of descent along with the gabbros and sheeted dykes of the MBO (Fig. 5). As seen above, the major-element compositions of the MBO plagiogranites are very similar to those of TTG compositions (Fig. 3); however, this similarity is not as evident for the trace elements (Fig. 5).

The plagiogranites show broadly coherent trends in the middle- to heavy-REE (M/HREE) on chondrite-normalized REE plots, but have variable light-REE (LREE) contents, from markedly enriched to relatively depleted patterns (e.g. 4.8–0.7 (La/Sm)<sub>N</sub>; Fig. 6a, c). The LREE-enriched patterns shown by the majority of the plagiogranite samples are inconsistent with the original definition of plagiogranites (Coleman & Peterman, 1975), and are shown to be enriched relative to the well-studied crustal plagiogranites from the Oman and Troodos Ophiolites (Fig. 6a). However, plagiogranites from the Sjenica (Milovanovic *et al.* 2012) and Tasriwine (Samson *et al.* 2004) ophiolites with LREE-enriched patterns have recently been reported (Fig. 6a). When compared with Archaean TTG compositions, plagiogranite samples are not as enriched in LREE (Fig. 6a). Most samples also show a slight chondrite-normalized enrichment in the heaviest REE relative to MREE, and display small U-shaped (concave upwards) patterns. The U-shaped patterns can be quantified using the Dy/Dy\* ratio of Davidson, Turner & Plank (2012), which ranges over

0.96–0.43 (Fig. 6b). Most plagiogranites have weak positive Eu anomalies (1.06–1.51 (Eu/Eu)\*), with only three samples having negative Eu anomalies (0.74–0.94; Fig. 6c). Interestingly, two of the three samples with negative Eu anomalies are also significantly enriched in LREE.

On normal mid-ocean-ridge basalt (N-MORB) normalized multi-element plots most plagiogranites display relatively flat patterns at concentrations just below N-MORB, with positive Th anomalies and negative Nb–Ta–Ti anomalies (Fig. 7a). Zr and Hf contents vary from enriched to depleted, relative to N-MORB. Most samples also have positive Sr anomalies; however, three samples have negative Sr anomalies, two of which display corresponding negative Eu anomalies (Fig. 6c).

## 6. Discussion

The modal abundance of quartz and plagioclase in combination with the low K<sub>2</sub>O contents (<1.1 wt%) of the MBO plagiogranites is similar to oceanic plagiogranites found elsewhere (e.g. Gerlach, Leeman & Ave Lallemand, 1981; Amri, Benoit & Ceuleneer, 1996; Rollinson, 2009). Additionally, the trace-element compositions of plagiogranites from the Muslim Bagh, Oman and Troodos ophiolites all show a high degree of compositional overlap (Fig. 7a, plagiogranite field) (Rollinson, 2009; Freund *et al.* 2014). Nevertheless, the LREE-enriched and slightly concave-upwards MREE patterns of the majority of MBO samples are distinct relative to the original oceanic plagiogranite definition (Coleman & Peterman, 1975; Coleman & Donato, 1979).

High SiO<sub>2</sub> (>70 wt%) and Na<sub>2</sub>O (3 < Na<sub>2</sub>O < 4.5 wt%) concentrations and low modal K-feldspar contents, low K<sub>2</sub>O/Na<sub>2</sub>O ratios and low Fe<sub>2</sub>O<sub>3</sub> + MgO + MnO + TiO<sub>2</sub> (most < 5 wt%) of the MBO plagiogranites make them compositionally similar to Archaean TTG as defined by Martin *et al.* (2005) and Moyen & Martin (2012). Additionally, when compared with Archaean TTG compositions on an N-MORB normalized multi-element plot, the plagiogranites display broadly similar concentrations, overlapping the TTG field at the lower LREE and higher HREE concentrations (Fig. 7b).

### 6.a. Plagiogranite petrogenesis

The majority of plagiogranites display enrichment in LREE relative to HREE (Fig. 6c), and all plagiogranites have negative Nb–Ta and positive Th anomalies (Fig. 7a). Additionally, the N-MORB-like concentrations of the other trace elements suggest that the plagiogranites (Fig. 6c) were generated at a MOR setting with a subduction input, likely a supra-subduction zone. This supports recent work by Kakar *et al.* (2014) who propose a supra-subduction model for the formation of the MBO. However, the petrogenesis of oceanic plagiogranites is controversial; fractional

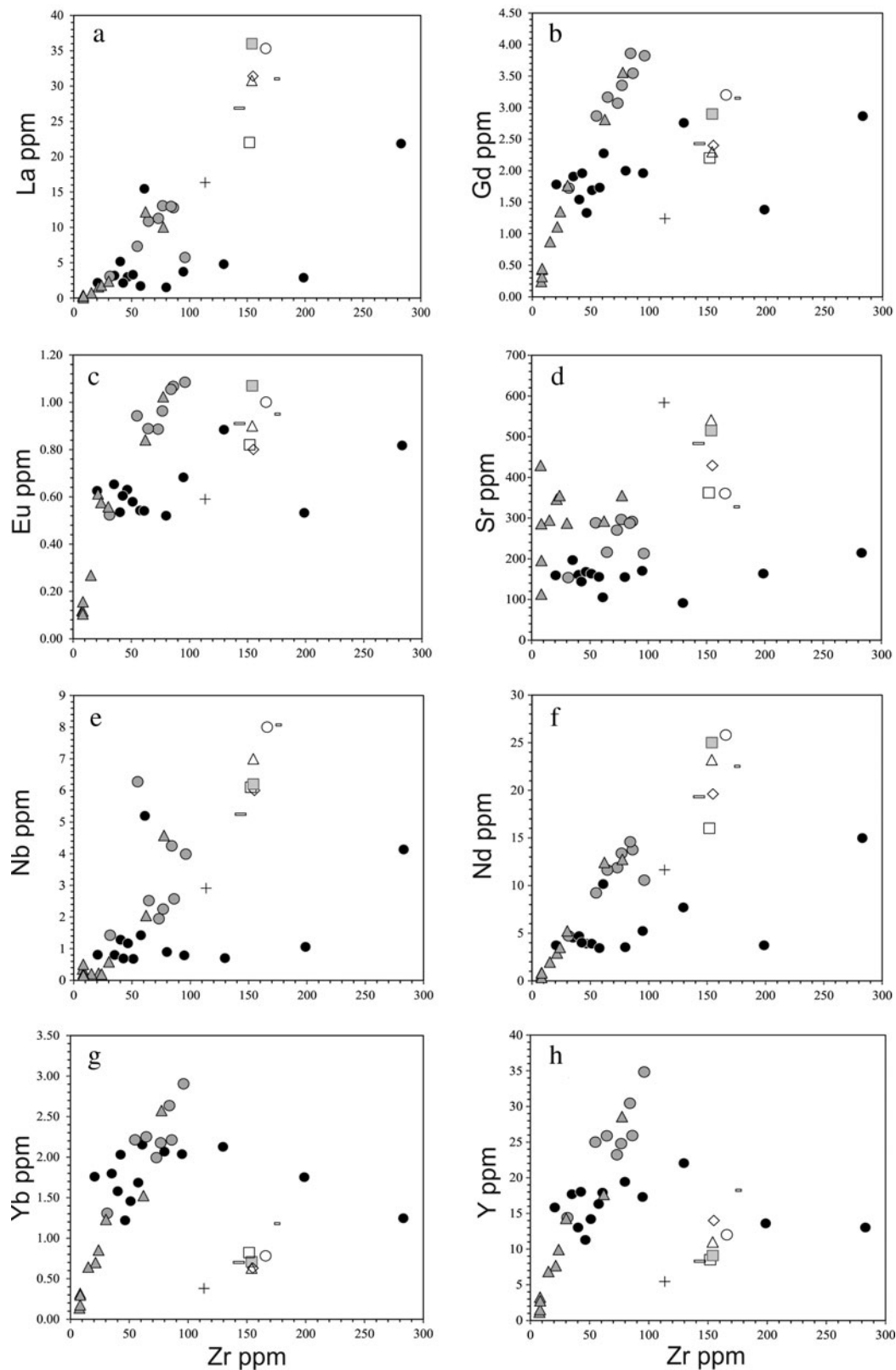


Figure 5. (a–h) Representative trace-element variation plots of the Muslim Bagh Ophiolite plagiogranites. The sheeted dykes and gabbros of the crustal section of the Muslim Bagh Ophiolite (data from Kakar *et al.* 2014) and Archaean TTG average compositions are also plotted; symbols and references as in Figure 3.



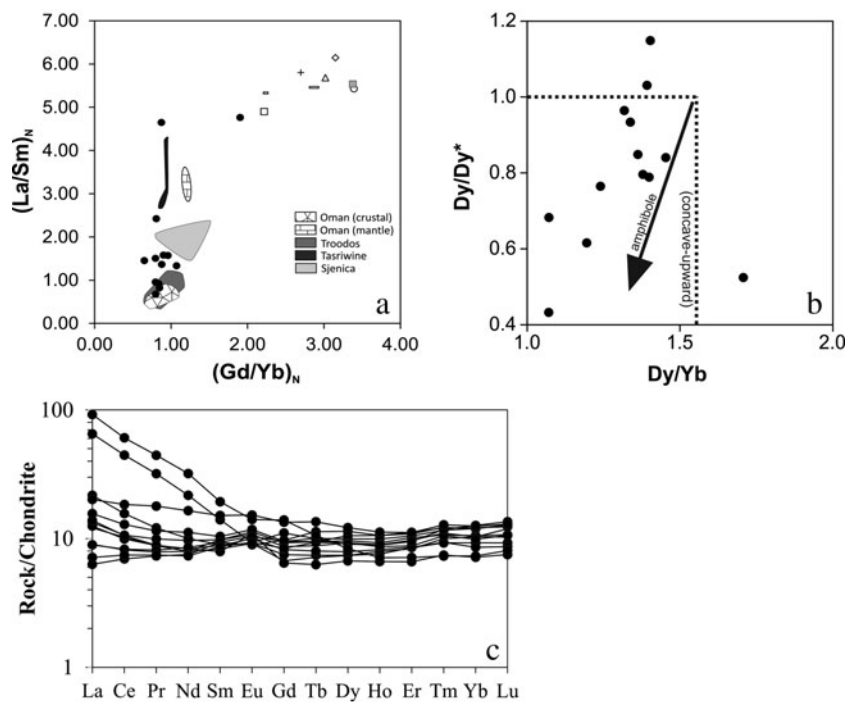


Figure 6. (a) Plot of  $(La/Sm)_N$  v.  $(Gd/Yb)_N$  highlighting the LREE-enriched nature of the majority of the Muslim Bagh Ophiolite plagiogranites relative to the depleted Oman (Rollinson, 2009) and Troodos Ophiolites (Freund *et al.* 2014). Also plotted are LREE-enriched plagiogranites from the Sjenica (Milovanovic *et al.* 2012) and Tasriwine Ophiolites (Samson *et al.* 2004) and Archaean TTG average compositions (symbols and references as in Fig. 3). (b) Plot of  $Dy/Dy^*$  v.  $Dy/Yb$  showing the majority of the Muslim Bagh Ophiolite plagiogranites to plot in the concave-upwards quadrant (black dotted lines) and follow the amphibole vector (arrow) in figure 4 of Davidson, Turner & Plank (2012). The plot quantifies the degree of concavity, and supports a role for amphibole in the petrogenesis of the plagiogranites either as a residual or crystallizing phase. (c) Chondrite-normalized REE plot of the Muslim Bagh Ophiolite plagiogranites. Normalizing values after Sun & McDonough (1989).

crystallization, partial melting or silicate–liquid immiscibility are variously proposed as petrogenetic models (see Koepke *et al.* 2007 for a review). In the following sections, we discuss the implications the plagiogranite compositions have for each of the possible petrogenetic models.

#### 6.a.1. Fractional crystallization and liquid immiscibility

The layered gabbros and sheeted dykes of the MBO crustal section represent possible cumulates and parental melts, respectively, from which to derive the plagiogranites by crystallization. However, major- and trace-element variation diagrams (Figs 3, 5) show that the plagiogranites do not plot along the same liquid lines of descent as any of the other MBO units. The fact that the plagiogranites define their own distinct field clearly indicates that they are not related to the other units by simple fractional crystallization processes. The lack of intermediate units within the ophiolite sequence also argues against an origin for the plagiogranites by fractional crystallization from a basic parental melt. Additionally, the narrow  $SiO_2$  range of the plagiogranites would suggest that fractional crystallization did not play a primary role in their petrogenesis.

Concave-upwards patterns displayed by the plagiogranites (Fig. 6c) support a role for amphibole during

their petrogenesis, as a result of amphiboles preference for MREE over LREE and HREE (e.g. Davidson, Turner & Plank, 2012). However, the concave-upwards pattern on its own does not indicate whether amphibole was crystallizing from a parental magma or acting as a residual phase during the fusion of a mafic protolith.

An origin by silicate–liquid immiscibility (e.g. Dixon & Rutherford, 1979) is also unlikely for the MBO plagiogranites. This is evidenced by the absence of the associated immiscible Fe-rich liquid (as Fe-rich mafic units) from the MBO.

#### 6.a.2. Partial melting

Experimental work of Koepke *et al.* (2004) and France *et al.* (2010) has shown that low  $TiO_2$  contents (<1 wt%) are characteristic of oceanic plagiogranites derived through partial melting of a mafic protolith; this is a consequence of the gabbroic protoliths having initially low  $TiO_2$  contents, typical of cumulate gabbros of the oceanic crust (Koepke *et al.* 2004, 2007). Low  $TiO_2$  contents of the MBO plagiogranites (Fig. 3b) are similar to those in the experimentally derived high- $SiO_2$  melts of Koepke *et al.* (2004), suggesting they were derived by partial melting of a gabbroic protolith in the crustal sequence of the MBO. In addition,  $TiO_2$  contents of the MBO plagiogranites plot below the boundary line drawn by Koepke *et al.* (2007)

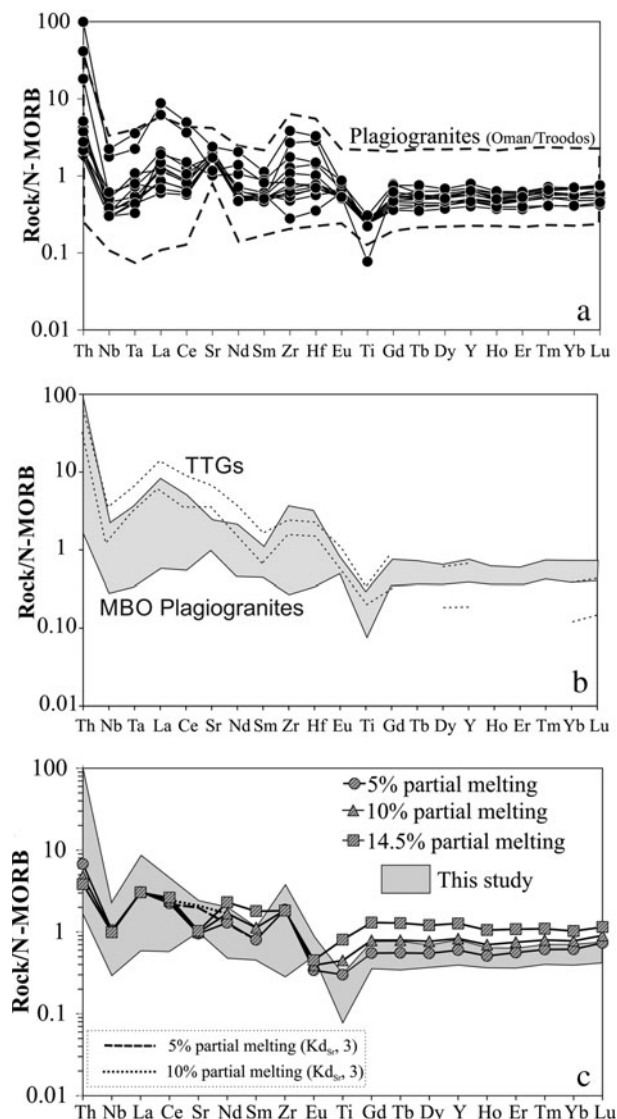


Figure 7. (a) Normal-MORB normalized multi-element plot of the Muslim Bagh Ophiolite plagiogranites. Dashed plagiogranite field represents analyses of plagiogranites from the Troodos (Freund *et al.* 2014) and Oman (Rollinson, 2009) Ophiolites. (b) Normal-MORB normalized multi-element plot comparing the Muslim Bagh Ophiolite plagiogranites with Archaean TTG average compositions (Condie, 2005; Martin *et al.* 2005; Moyen & Martin, 2012). (c) Trace-element modelling of batch melting. The primitive mantle-normalized multi-element plot compares the trace-element composition resulting from trace-element melt modelling of a crustal hornblende gabbro with the composition of the Muslim Bagh Ophiolite plagiogranites. Plagiogranite compositions can be replicated by 5–10% partial melting of a hornblende gabbro. Dashed black lines represent melts derived by partial melting when using a lower (i.e. 3; Laurent *et al.* 2013) partition coefficient for Sr in plagioclase. Normalizing values after Sun & McDonough (1989).

which separates plagiogranites derived by hydrous partial melting (plot below black dashed line, Fig. 3b) from those plagiogranites derived by crystallization or immiscibility processes (plot above black dashed line).

Additionally, as shown in Figure 3, major-element concentrations of the MBO plagiogranites are similar to Archaean TTG (Condie, 2005; Martin *et al.* 2005;

Moyen & Martin, 2012), generally regarded to have been generated through partial melting of a mafic igneous protolith (e.g. Drummond, Defant & Kepezhinskias, 1996; Foley, Tiepolo & Vannucci, 2002; Rapp, Shimizu & Norman, 2003; Martin *et al.* 2005; Moyen & Stevens, 2006; Nutman *et al.* 2009; Hastie *et al.* 2015, 2016). We suggest that the lower K<sub>2</sub>O contents displayed by the plagiogranites, compared with Archaean TTG, is the result of the TTG rocks being derived from a more primitive mantle prior to continental crust extraction, and therefore a less-depleted mantle than the present. Trace-element variation plots (Fig. 5) do not show as convincing a similarity between the MBO plagiogranites and Archaean TTG as the major-element variation plots (Fig. 3), however. Nevertheless, the MBO plagiogranites have broadly similar trace-element compositions to Archaean TTG (Fig. 7b).

Negative Eu and Sr anomalies (Fig. 6c, 7a) and decreasing Al<sub>2</sub>O<sub>3</sub> with increasing SiO<sub>2</sub> (Fig. 3) in some samples could potentially be explained by a small amount of late-stage plagioclase fractional crystallization. However, negative Eu and Sr anomalies can also be the result of plagioclase in the melting residue, while the decrease in Al<sub>2</sub>O<sub>3</sub> with SiO<sub>2</sub> can be reproduced through small degrees of partial melting as demonstrated by Beard & Lofgren (1991). In the following section we use trace-element modelling to test a partial melting model for the MBO plagiogranites.

### 6.b. Modelling of partial melting

To model the partial melting of a mafic protolith, the non-modal batch melting equation of Shaw (1970) was used for the calculations:

$$C_1 = \frac{C_0}{D_0 + F(1 - P)} \quad (1)$$

where  $C_1$  is the concentration of a particular trace element in a resultant melt,  $C_0$  is the concentration of an element in the source region prior to partial melting,  $F$  is the mass fraction of melt generated,  $D_0$  is the bulk partition coefficient of an element prior to partial melting and  $P$  is the partition coefficient of an element weighted by the proportion contributed by each mineral phase to the melt. Hornblende gabbro, C51 (from Kakar *et al.* 2014) was used as the protolith. This sample was collected from the cumulate sequence of the crustal section of the MBO and was chosen as the protolith since the concave-upwards pattern shown by the plagiogranites suggests that amphibole may have been left behind in the melting residue. The partition coefficients used are those for elements in equilibrium with TTG-like silicic melts from Bedard (2006). Mineral modes of the hornblende gabbro are those of Kakar *et al.* (2014) and Siddiqui *et al.* (1996). Melt modes were calculated using 1 kbar experimental runs from Beard & Lofgren (1991) as they provide enough petrological information to carry out the calculation. Melting was stopped at 14.5%, as this is the point

at which hornblende is exhausted from the protolith. Mineral and melt modes, partition coefficients, hornblende gabbro starting composition and model results can be found in Supplementary Table S3, available at <http://journals.cambridge.org/geo>.

Figure 7c shows that the incompatible trace-element patterns (including negative Nb and Ti anomalies and positive Th and Zr anomalies) of the plagiogranites can be replicated by 5–10% partial melting of the hornblende gabbro. Nonetheless, the modelling generates a larger negative Sr anomaly than seen in the MBO plagiogranites. This result is attributed to the use of a high Sr partition coefficient in plagioclase (6.65; Beard, 2006) and this discrepancy can be removed if a lower partition coefficient is used (i.e. 3, based on the range reported by Laurent *et al.* 2013).

Despite the evidence supporting a partial melting model for the MBO plagiogranites, the reason behind the negative K<sub>2</sub>O trend displayed by the plagiogranites when plotted against SiO<sub>2</sub> (Fig. 3) is uncertain. It is however possible that the negative trends displayed by both K<sub>2</sub>O and Al<sub>2</sub>O<sub>3</sub> are the result of an interplay between fractional crystallization (plagioclase and biotite?) and/or varying degrees of partial melting and source variation.

#### 6.c. Comparison with other Tethyan Ophiolite plagiogranites and implications for the tectonomagmatic setting of the MBO

As we have shown, some geochemical characteristics of the MBO plagiogranites (i.e. LREE enriched and concave-upwards MREE patterns) do not conform to the definition of oceanic plagiogranites as proposed by Coleman & Peterman (1975). The results from this study are similar to previous plagiogranite analyses from the MBO presented by Kakar *et al.* (2014), who also report MBO plagiogranites with LREE-enriched patterns (1–7, (La/Sm)<sub>N</sub>), as well as negative Nb–Ta–Ti anomalies and low TiO<sub>2</sub> contents ( $\leq 0.20$  wt %).

The MBO plagiogranites are significantly different from those of other Tethyan Ophiolites in terms of both field and geochemical characteristics. First, LREE contents of Troodos and Oman ophiolite crustal plagiogranites are relatively depleted compared with HREE (Fig. 6a) (Rollinson *et al.* 2009; Freund *et al.* 2014); a more depleted source is therefore required for these plagiogranites relative to the MBO plagiogranites. It is however beyond the scope of this study to investigate further the difference in source enrichment between the MBO plagiogranites and those plagiogranites situated in the Oman and Troodos ophiolites. Secondly, the plagiogranites of the MBO are solely located in the crustal section of the ophiolite, whereas geochemically distinct groups of plagiogranites have been identified in crust and mantle sections of the Troodos and Oman ophiolites (Rollinson, 2009, 2014; Freund *et al.* 2014). Thirdly, the MBO plagiogranites are generally smaller intrusive bodies (on a scale of no more than a few metres) than those found in both the

Troodos and Oman ophiolites, where plagiogranites range from several tens of metres to kilometre-sized plutons (Rollinson *et al.* 2009; Freund *et al.* 2014).

The poorly developed sheeted dyke complex (Khan, Kerr & Mahmood, 2007; Kakar *et al.* 2014) of the MBO crustal section is likely the result of the imbalance between spreading rate and magma supply in a supra-subduction zone tectonic setting (Robinson *et al.* 2008). Robinson *et al.* (2008) have proposed that both the fore-arc and back-arc of a supra-subduction zone generally experience lower magma supply rates, due to eruptions at the volcanic arc, and high extensional strain rates. The small size, restricted distribution and lack of geochemical variability (i.e. uniform composition) among the MBO plagiogranites could therefore be a result of this decreased magma supply in the supra-subduction zone where the MBO crystallized. Consequently, the decreased magma supply results in a small degree of partial melting of the plagiogranite source (i.e. crustal hornblende gabbros).

#### 6.d. Implications for Archaean TTG genesis

Most previous and current research into Archaean TTG petrogenesis favours models in which juvenile Archaean continental crust is generated by partial melting of mafic igneous protoliths (e.g. Sen & Dunn, 1994; Wolf & Wyllie, 1994; Foley, Tiepolo & Vannucci, 2002; Rapp, Shimizu & Norman, 2003; Moya & Stevens, 2006; Laurie & Stevens, 2012; Zhang *et al.* 2013; Ziaja *et al.* 2014; Hastie *et al.* 2016), the setting of which is still controversial; both subduction/flat slab subduction/underthrusting (e.g. Drummond, Defant & Kepezhinskas, 1996; Martin *et al.* 2005; Nutman *et al.* 2009; Hastie *et al.* 2015) and intracrustal (Hamilton, 1998; Hawkesworth, Cawood & Dhuime, 2016) settings have been proposed for the derivation of Archaean TTG of various ages.

Since the original definition of oceanic plagiogranites in the mid-1970s by Coleman & Peterman (1975), oceanic plagiogranites have been shown to differ compositionally from Archaean TTG by being less potassic and having MORB-like LREE and flat HREE patterns. Numerous studies on oceanic plagiogranites from the Oman Ophiolite (Rollinson, 2008, 2009, 2014) have suggested that although the Oman Ophiolite plagiogranites have compositions that are similar to oceanic plagiogranites (as defined by Coleman & Peterman, 1975) and differ compositionally from Archaean TTG, they can be instructive on Archaean TTG genesis. Rollinson (2009) noted that, in addition to the conditions of plagiogranite petrogenesis, a source region enriched in LREE is also required in order to generate the LREE-enriched nature of Archaean TTG. Additionally, Rollinson (2008) has suggested that trondhjemite (plagiogranite) petrogenesis in the Oman Ophiolite acts as a possible analogue for the generation of Earth's first felsic crust during Hadean time. Rollinson (2008) has argued that early (Hadean) felsic crust was of low



volume, and this corresponds to the low volume of plagiogranites we see in recent ophiolite sequences.

The MBO plagiogranites are compositionally different (LREE-enriched and concave-upwards MREE patterns) from the original oceanic plagiogranite definition, but are geochemically similar to Archaean TTG (e.g. Condie, 2005; Martin *et al.* 2005; Moyen & Martin, 2012) (Figs 3, 7b). Consequently, the MBO plagiogranites can be used as a recent (Late Cretaceous) analogue to investigate the formation of some Archaean TTG rocks.

The MBO plagiogranites are found within mafic crust that was formed at a convergent margin, specifically the upper plate above the subduction zone (e.g. Siddiqui *et al.* 1996, 2011; Kakar *et al.* 2014). The similarity in composition between the MBO plagiogranites and Archaean TTG suggests that some of the earliest silicic continental crust may have been derived from melting the overriding plates in primitive subduction-like zones. We acknowledge that there is a contrast in volume between the MBO plagiogranites and Archaean TTG; however, we infer that the genesis of these plagiogranites can be instructive on the generation of some, but not all, Archaean TTG. In addition, the overall greater enrichment in LREE relative to HREE of Archaean TTG compared with the MBO plagiogranites suggests that to source a larger portion of Archaean TTG requires a slightly more enriched source than that of the MBO plagiogranites (e.g. Rollinson, 2009). Again, this could possibly be due to the extraction of continental crust and depletion of the mantle over time.

## 7. Conclusions

1. Oceanic plagiogranites of the MBO are exclusively located at the base and middle portions of the sheeted dyke complex, where they form small, intrusive dyke-like bodies and lenses.

2. Low TiO<sub>2</sub> contents (<0.4 wt%) in the plagiogranites and a lack of intermediate rocks in the sheeted dyke complex suggest an origin by partial melting of mafic rocks. This is confirmed by batch melt trace-element modelling of a crustal hornblende gabbro from the crustal sequence of the MBO. This modelling shows that the plagiogranites can be replicated by 5–10% partial melting, possibly with a small degree of late-stage fractional crystallization of plagioclase(?) to account for negative Sr and Eu anomalies and a decrease in Al<sub>2</sub>O<sub>3</sub> with SiO<sub>2</sub>.

3. The similarity in composition of the MBO plagiogranites to Archaean TTG rocks supports the model that some Archaean TTG could be generated by partial melting of a mafic protolith, possibly in the overriding plate of a subduction-like zone.

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## Declaration of interest

None.

## Supplementary material

To view supplementary material for this article, please visit <https://doi.org/10.1017/S0016756818000250>

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