

# A method to determine the flux limiter via the motion of the M-band emission region in Au hohlraum

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## Abstract

In this paper, the influence of the electron flux limiter  $f_e$  on hohlraum plasmas is studied by using the two-dimensional code LARED-H, and a method to experimentally determine  $f_e$  via the motion of the M-band emission region in Au hohlraum is proposed. From our simulations, the limited free streaming flux may dominate the heat conduction in the regions with steep temperature gradient, including the laser deposition region, the flux-heated overcritical region, and the laser channel boundary between the hot laser plasmas and the surrounding radiation ablated plasmas, while these are important X-ray emission regions. Hence, the choosing of  $f_e$  may influence the wall plasma expansion and the laser spot motion, and further influence the motion of the emission regions. From our study, the motion of the M-band ( $>1.5$  keV) emission region is sensitive to  $f_e$  when the limited free streaming flux dominates the heat conduction of the wall plasma expansion region, and so it is possible to determine  $f_e$  via the motion of the M-band emission region. In this work, the model used in our simulations is taken from the wall and laser spot motion experiments done by Huser *et al.* (2009).

**Keywords:** Flux limiter; Hohlräume; Plasma expansion

## INTRODUCTION

Electron heat conduction plays a major role in inertial fusion physics, because it carries the deposited laser energy from the critical surface to the ablation front (Lindl, 1995; Lindl *et al.*, 2003; Atzeni & Meyer-ter-Vehn, 2004; Gupta & Godwal, 2001). The calculation of the electron transport influences the laser deposition and plasma status inside the hohlraum, it further influences the plasma emission, the laser plasma interaction and the radiation uniformity of the capsule placed in the center of the hohlraum, and finally, it significantly influences the ignition. Therefore, worldwide experimental and theoretical efforts on electron transport in plasmas are in progress (Sunahara *et al.*, 2003; Colombant *et al.*, 2005; Shurtz *et al.*, 2000), especially in recent initial ignition campaign at the National Ignition Facility (Lindl & Moses, 2011; Haan *et al.*, 2011; Rosen *et al.*, 2011). Up to now, various theoretical models have been developed for electron transport, such as the classical electron transport model of Spitzer-Härm (1953; Cohen *et al.*, 1950), the flux

limit model (FL model) (Malone *et al.*, 1975), the nonlocal electron transport model (Luciani *et al.*, 1983), and the non-Maxwellian model (Huo *et al.*, 2012). Among all these models, the FL model is the most widely used one in calculating the electron heat flux in hohlraum simulation codes, like LASNEX (Zimmerman & Kruer, 1975), HYDRA (Marinak *et al.*, 1996), FCI2 (Buresi *et al.*, 1986), and LARED-H (Pei, 2007). In the last decades, numerical experiments are designed with FL model (flux limiter) and good results were obtained (Li *et al.*, 2010; Lan *et al.*, 2010, 2004; Cook *et al.*, 2008; Seifter *et al.*, 2009; Orlov *et al.*, 2011). In the FL model, a flux limiter  $f_e$  is taken to make sure that the flux does not exceed the limit of the free streaming heat flux  $Q_f$ , and the heat flux is often computed either as  $\min(Q_{SH}, f_e Q_f)$  or in a “parallel connection” formulation  $[Q_{SH}^{-1} + (f_e Q_f)^{-1}]^{-1}$ , where  $f_e$  is often chosen from 0.05 to 0.15 (Atzeni & Meyer-ter-Vehn, 2004; Rosen & Nuckolls, 1979; Kline *et al.*, 2011). In recent theoretical work done for the National Ignition Campaign (Rosen *et al.*, 2011), a flux limiter of 0.15 was finally taken in the ignition design after comparing with the nonlocal model. However, the choice of  $f_e$  depends on the hohlraum and the laser parameters used in the study. Hence, it is necessary to find a way to determine  $f_e$  experimentally for different designs.

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In fact, the heat flux may be dominated by the limited free-stream flux in the regions with a steep temperature gradient, including the laser deposition region, the flux-heated overcritical region, and the laser channel boundary between the hot laser plasmas and the surrounding radiation ablated plasmas, while these are important X-ray emission regions. As a result, the choice of  $f_e$  may influence the laser spot motion and the wall plasma expansion, and therefore influence the motion of the emission regions. Therefore, it is possible to determine  $f_e$  via the diagnostics of the X-ray emissions. In this paper, we will study the influence of the electron flux limiter  $f_e$  on hohlraum plasmas by using a two-dimensional (2D) radiation hydrodynamics code LARED-H, and we will propose a method to experimentally determine  $f_e$  via the motion of the M-band ( $>1.5$  keV) emission region in Au hohlraum.

The hohlraum and laser parameters used in this study are taken from the experiments done by Huser *et al.* (2009). In their experiments, the authors measured the wall and laser spot motion in the empty, propane filled, and CH-lined Au hohlraums on the Omega laser facility, and compared the experimental results with their simulations with the 2D radiation hydrodynamics code FCI2. Huser *et al.* (2009) observed the wall motion by using axial 2D X-ray imaging and observed the laser spot motion perpendicularly through a thinned wall using streaked hard X-ray imaging. Their experiments gave valuable quantitative results, which are worthy to be compared with the 2D simulations. We will mainly simulate the empty hohlraum used in the experiments (Huser *et al.*, 2009) with different  $f_e$  and choose a suitable  $f_e$  by comparing our simulations with the observed motion of the X-ray emission regions. Using the chosen  $f_e$ , we will also simulate the gas-filled hohlraum used in the experiments of Huser *et al.* (2009) and study the effect of gas filling on the wall plasma expansions.

The paper is organized as follows. In Section 2, we will introduce the code and model used in our study. In Section 3, we will use the LARED-H to simulate the empty hohlraum used in Huser's *et al.* (2009) experiments with different  $f_e$  and study the influence of  $f_e$  on plasma status. In Section 4, we will study the influence of  $f_e$  on the X-ray emission region in the overcritical plasmas. In Section 5, we will study the influence of  $f_e$  on the X-ray emission from the plasma corona. In above simulations and discussions, we mainly focus on the empty hohlraum. In Section 6, we study the effect of gas filling on the wall plasma expansion in the gas-filled hohlraum used in Huser's *et al.* (2009) experiments. Finally, we will present a summary in Section 7.

## 2. CODE AND MODEL

The LARED-H is a 2D radiation hydrodynamics code for laser target coupling and hohlraum physics study, which is one of the main simulation codes in designing and analyzing the hohlraum experiments on SGII and SGIII Prototype laser facilities (Li *et al.*, 2011; Huo *et al.*, 2010). The three-

temperature hydrodynamic equations in LARED-H are:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{u}) = 0, \quad (1)$$

$$\rho \frac{d\mathbf{u}}{dt} + \nabla(P_e + P_i + P_r + q) = 0, \quad (2)$$

$$C_{ve} \frac{dT_e}{dt} = W_{pe} + W_{fe} + W_l + W_{er} + W_{ei}, \quad (3)$$

$$C_{vi} \frac{dT_i}{dt} = W_{pi} + W_{fi} - W_{ei}, \quad (4)$$

$$\frac{d}{dt} \left( \frac{E_r}{\rho} \right) = W_{pr} + W_{fr} - W_{er}. \quad (5)$$

Here, the subscripts  $e$ ,  $i$ , and  $r$  indicate electron, ion, and radiation, respectively; the symbol  $\rho$  denotes the mass density,  $\mathbf{u}$  is the flow velocity;  $T$  is the temperature;  $P$  is the pressure, and  $q$  is the artificial viscosity;  $C_{ve}$  and  $C_{vi}$  are the specific heats of free electron and ion, respectively;  $W_p$  is the work term due to pressure;  $W_f$  is the thermal conduction term;  $W_{er}$  is the electron-radiation energy exchange term;  $W_{ei}$  is the electron-ion energy exchange term; and  $E_r$  is the radiation energy density;  $W_l$  is the laser energy deposition via inverse bremsstrahlung and is calculated with a three-dimensional ray-tracing package. An average-atom atomic physics model is included in the code, and the following atomic processes are involved: (1) collisional excitation and de-excitation, (2) spontaneous emission, (3) collisional ionization and three-body recombination, (4) photon ionization and radiative recombination, (5) dielectron capture and auto-ionization.

The expressions of the above energy terms are:

$$W_{fe} = -\frac{1}{\rho} \nabla \cdot \mathbf{F}_e, \quad (6)$$

$$W_{fi} = -\frac{1}{\rho} \nabla \cdot \mathbf{F}_i, \quad (7)$$

$$W_{fr} = -\frac{1}{\rho} \nabla \cdot \mathbf{F}_r, \quad (8)$$

$$W_{pe} = -P_e \frac{d}{dt} \left( \frac{1}{\rho} \right), \quad (9)$$

$$W_{pi} = -(P_i + q) \frac{d}{dt} \left( \frac{1}{\rho} \right), \quad (10)$$

$$W_{pr} = -P_r \frac{d}{dt} \left( \frac{1}{\rho} \right), \quad (11)$$

$$W_{ei} = 4 \sqrt{\frac{2\pi m_e}{k_B}} \frac{N_e N_i}{\rho} \times \frac{e^4 z^2}{m_i} \times \frac{T_e - T_i}{T_e^{3/2}} \times \ln \Lambda_{ei} \quad (12)$$

$$W_{er} = W_{fb} + W_{bb} + W_{ff}, \quad (13)$$

Here,  $m_e$  is the electron mass;  $m_i$  is the ion mass;  $e$  is the electron charge;  $\ln \Lambda_{ei}$  is the Coulomb logarithm for collisions between electrons and ions;  $k_B$  is the Boltzmann constant;  $z$  is the average ionization degree;  $N_i$  is the ion density, and

$N_e$  is the electron density;  $W_{bf}$ ,  $W_{bb}$ , and  $W_{ff}$  represent the energy exchanges due to free-bound, bound-bound and free-free transitions of atomic processes, respectively. The electron heat flux  $F_e$  taken in our code is the minimum of the S-H flux<sup>11</sup>  $F_{SH}$  and the limited free-stream heat flux  $f_e F_l$ , here  $F_l$  is the free-stream heat flux.

In LARED-H, the thermodynamic quantities are derived either from the ideal gas model or from data of realistic equation of state. The mean opacity is calculated with relativistic  $F_{SH}$  self-consistent average-atom-model OPINCH (Serduke *et al.*, 2000) where the contributions from free-free, free-bound, and bound-bound transitions are taken into account.

We had used LARED-H to simulate the laser spot moving experiment on NOVA<sup>2</sup>. In the experiment, a 1 ns flat-top laser pulse was used to drive a hohlraum of 2500  $\mu\text{m}$  in length and 1600  $\mu\text{m}$  in diameter with one laser entrance hole of 800  $\mu\text{m}$  diameter at each end. From LARED-H, the relationship between the laser power  $P_L$  and the angular velocity  $d\theta/dt$  of the emission center seen by capsule is  $d\theta/dt = 5.7^\circ/\text{ns} + 3.2^\circ/\text{ns}(\log_{10} P_L)$ , which is obtained by using the least square method.

In this work, the hohlraum and laser parameters used are taken from the experiments done by Huser *et al.* (2009). Both empty and propane-filled hohlraums are considered. The cylindrical Au hohlraum is 1.5 mm long with 1.6 mm diameter, with a full-open laser entrance hole at each end. The hohlraum wall thickness is taken as 25  $\mu\text{m}$ . The laser pulse of about 9.5 kJ in total has the shape of Omega PS26N01A with 1 ns pre-pulse followed by 1.5 ns main-pulse with 5:1 contrast. The thirty laser beams simultaneously irradiate the hohlraum from two ends, with respectively, 10 beams at an incidence cone of  $42^\circ$  angle and 20 beams at  $59^\circ$  angle. In our 2D simulations, the widths of both laser rings at  $42^\circ$  and  $59^\circ$  are taken as 400  $\mu\text{m}$  on the inner wall of hohlraum. The laser rings at  $42^\circ$  from two sides not only overlap each other at the hohlraum center, but also overlap the rings at  $59^\circ$  from the same side at about  $z = 200 \mu\text{m}$ . Here,  $z$  is the axial position from hohlraum center. Thus, the middle part of the hohlraum wall, around 600  $\mu\text{m}$  in length, is fully under the irradiation of the laser. In addition, the initial width of the first Lagrangian cell on the inner surface of wall is taken as 10 nm and the initial temperature of the hohlraum is taken as 300 K. For the empty hohlraum, a pseudo-gas of Au with  $8 \times 10^{-4} \text{g/cm}^3$  initial density is used, and it is assumed that there is no laser deposited in the pseudo-gas. For the X-ray emissions, same as in the experiments (Huser *et al.*, 2009), we also focus on O-band (about 450 eV), N-band (about 800 eV) and M-band ( $> 1.5 \text{keV}$ ) and use the same spectral responses for these three bands.

### 3. HOHLRAUM PLASMA STATUS UNDER DIFFERENT FLUX LIMITERS

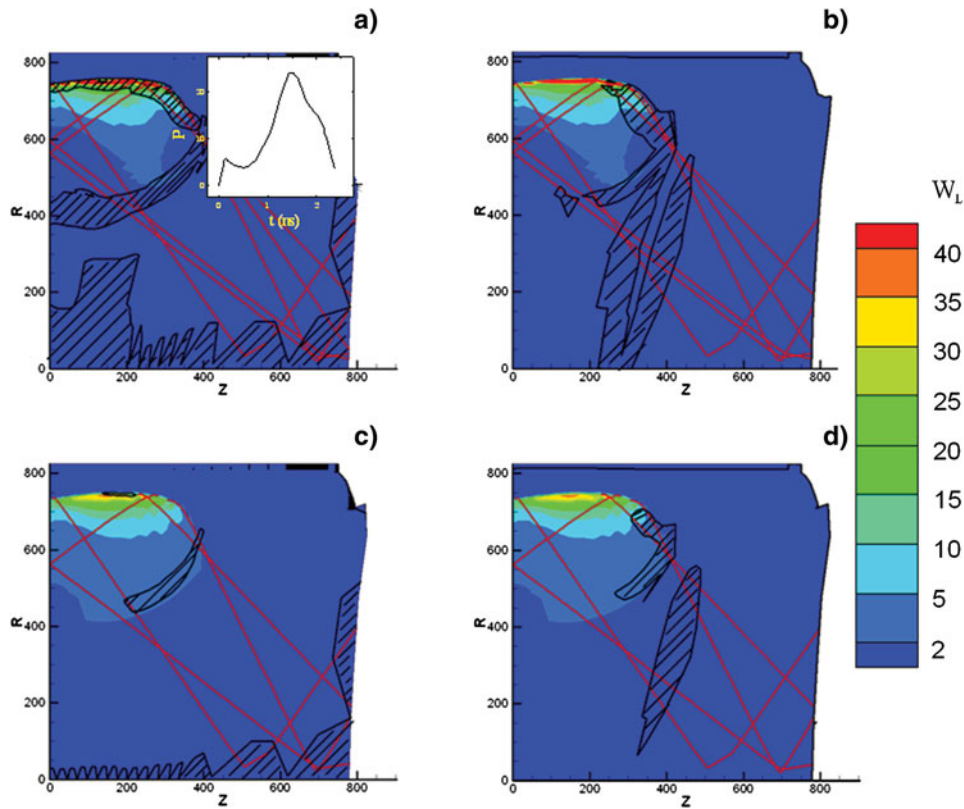
The electron conduction carries the deposited laser energy from the critical surface to the over-critical plasma and the

plasma corona, so the calculation of the heat conduction influences the plasma status and the X-ray emissions inside hohlraum. In this section, we study the 2D simulation results of the empty hohlraum model and compare the hohlraum plasma status under different  $f_e$ . From our simulations, the results are sensitive to  $f_e$  when  $f_e$  is smaller than 0.1, but insensitive when  $f_e > 0.1$ . Therefore, we mainly present below the results with  $f_e = 0.05$  and 0.1.

First, one should know in which regions the heat conduction is limited by the flux limiters inside hohlraum. In Figure 1, we present the 2D maps of the laser energy deposition term  $W_l$  and the areas in which the heat flux is dominated by the limited free-stream flux  $f_e F_l$  with respect to  $f_e$  of 0.05 and 0.1, at 1 ns when the pre-pulse ends while the main pulse starts. The pulse shape of drive laser is also presented in Figure 1a. As shown, the heat flux is dominated by  $f_e F_l$  in the regions with steep temperature gradient, including the laser deposition region, over-critical plasma region and the laser channel boundary between the hot laser plasmas and the surrounding radiation ablated plasmas, which are important regions for the X-ray emissions. However, the limited regions are narrower at a higher  $f_e$ . When  $f_e$  is taken as 0.2, from our simulations, the heat conduction is mainly dominated by the S-H flux almost in the whole hohlraum. Notice that  $f_e F_l$  dominates the heat conduction also in the pseudo-gas and at the boundary between the pseudo-gas and the laser channel, but we do not concern these regions that are related to the pseudo-gas in this paper. From our simulation, the plasma-filling is not serious even at the peak laser time. Thus, the laser is mainly deposited near wall for this model.

Now, we study and compare the plasma status along the laser channel under different  $f_e$ . Because the laser energy at the  $59^\circ$  angle is two times as large as that at the  $42^\circ$ , so we mainly focus on the laser channel at  $59^\circ$ . In Figures 2 and 3, we present the distributions of  $T_e$ ,  $n_e$ , and ionization degree  $Z_{eff}$  along the laser channel at  $59^\circ$  at 1.7 ns, just after the laser begins to fall and the radiation temperature  $T_r$  in the hohlraum reaches its peak. As shown, the maximum  $T_e$  at  $f_e = 0.05$  is about 400 eV higher than at  $f_e = 0.1$ ,  $Z_{eff}$  at  $R_c$  is about 1.5 higher, and both  $R_c$  and the ionization front are around 10  $\mu\text{m}$  closer to wall. Here,  $R_c$  is defined as the critical position at  $n_e/n_c = 1$ . The reason for a higher maximum  $T_e$  at a smaller  $f_e$  is that less energy is conducted to the over-critical region and less mass ablated from wall when the heat conduction is more limited. As a result, at a smaller  $f_e$ , the laser ablated plasmas are hotter and expand faster along the radial direction, resulting in a lower  $n_e$  in the plasma corona. Obviously, the critical position and the ionization front move slower towards the hohlraum axis at a smaller  $f_e$ .

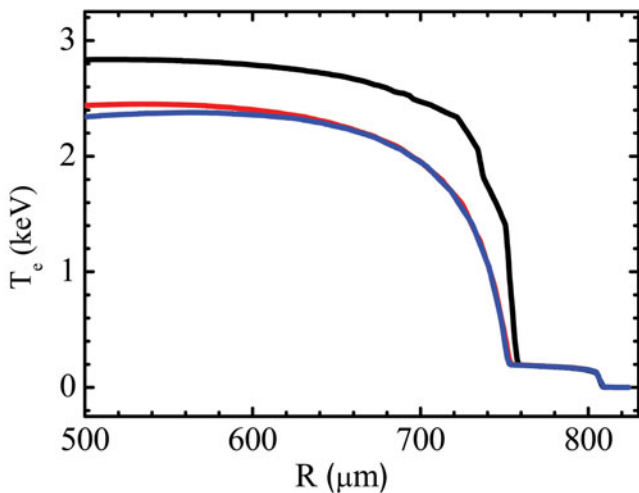
In Figure 4, we present the temporal  $n_e$  of the Lagrangian cells that produce respective maximum emissions of the three bands at  $f_e = 0.05$  and 0.1. As shown, for both flux limiters, during the whole laser driving, the maximum emissions of N-band and O-band are generated at the same positions in



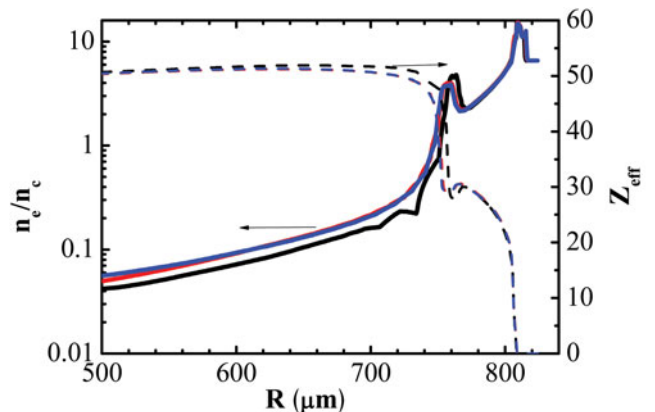
**Fig. 1.** (Color online) 2D maps of the laser energy deposition term  $W_L$  ( $10^{15}$  W/m<sup>3</sup>) and the areas in which the heat flux is dominated by  $f_L F_l$  along, respectively, the  $\hat{l}$  direction (**left images**) and the  $\hat{k}$  direction (**right images**), at 1 ns with  $f_e = 0.05$  (**upper images**) and 0.1 (**lower images**), in which the red lines represent the laser pass. The directions of  $\hat{l}$  and  $\hat{k}$  are defined for 2D simulation, which are time-dependent and related to cells. The initial directions of  $\hat{l}$  and  $\hat{k}$  are in  $\hat{r}$  and  $\hat{z}$ , respectively. The inner plot in (a) is the pulse shape of drive laser.

the overcritical region,  $n_e \approx 2-4$  during the main laser pulse, while the maximum emissions of M-band are generated in the laser deposited region with  $n_e < 1$ . From our simulations, the value of  $f_e$  influences not only the peaks of the O-band, N-band, and M-band emissions, but also the positions of

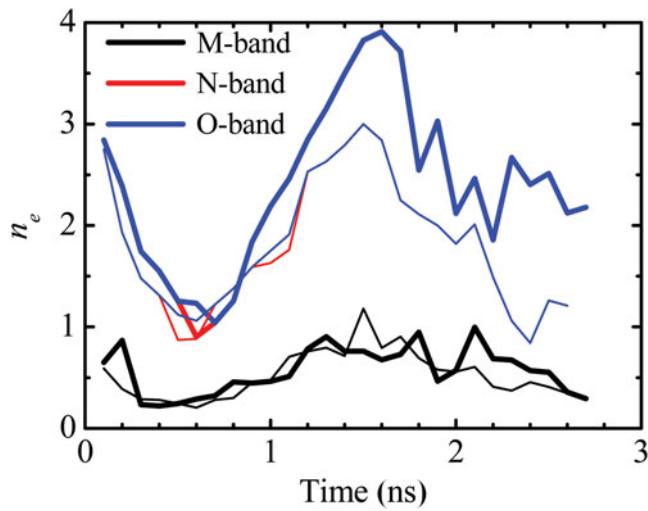
the emission regions. As shown in Figure 5, compared to the results at  $f_e = 0.1$ , the peaks of O-band, N-band, and M-band are about, respectively, 50%, 100%, and 200% higher at  $f_e = 0.05$  and the peak positions for all three bands are around 10  $\mu\text{m}$  closer to wall. Therefore, we can



**Fig. 2.** (Color online) The distributions of  $T_e$  along the laser pass of the 59° beams at 1.7 ns with  $f_e$  of 0.05 (black), 0.1 (red) and 0.2 (blue).



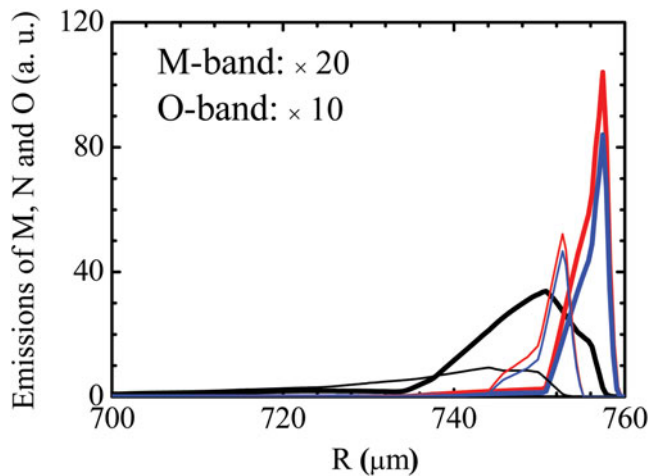
**Fig. 3.** (Color online) The distributions of  $n_e$  (solid) and  $Z_{\text{eff}}$  (dashed) along the laser pass of the 59° beams at 1.7 ns with  $f_e$  of 0.05 (black), 0.1 (red) and 0.2 (blue).



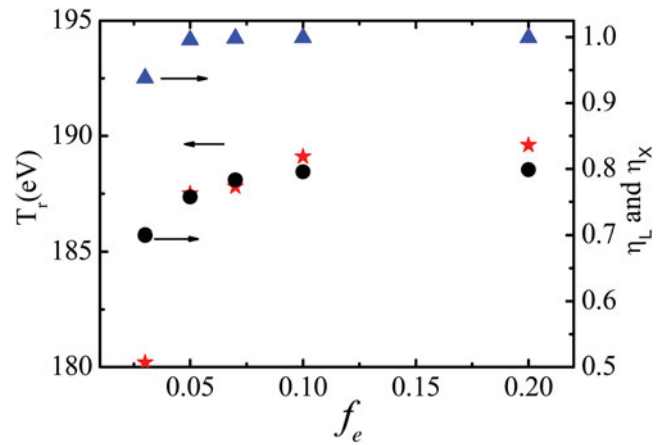
**Fig. 4.** (Color online) The temporal  $n_e$  of the Lagrangian cells which produce, respectively, the maximum M-band (black), N-band (red) and O-band (blue) emissions inside hohlraum, with  $f_e = 0.05$  (thick lines) and  $f_e = 0.1$  (thin lines). The curve of N-band almost overlaps with O-band during whole time.

determine  $f_e$  for a hohlraum design via the X-ray emissions from experiments.

We also study the influence of  $f_e$  on the hohlraum  $T_r$ , the laser absorption efficiency  $\eta_L$  and the laser X-ray conversion efficiency  $\eta_X$ . Shown in Figure 6 are variations of  $\eta_L$ ,  $\eta_X$ , and the post-processed  $T_r$  along the line  $40^\circ$  from hohlraum axis through LEH as  $f_e$  at the peak laser time 1.7 ns. For  $\eta_L$ , it is about 1 at  $f_e \geq 0.05$ , while it decreases to 94% at  $f_e = 0.03$ . The reason is that the laser deposition is lower in a hotter hohlraum generated with a smaller  $f_e$ , thus more laser energy escapes from the two LEHs. For  $\eta_X$ , it increases as  $f_e$ , because more mass is ablated and emits at a larger  $f_e$ . As shown,  $\eta_X$  is 70% at  $f_e = 0.03$  and increases to 80% after  $f_e \geq 0.07$ . For  $T_r$ ,



**Fig. 5.** (Color online) The distributions of M-band (black), N-band (red) and O-band (blue) emissions along the laser pass of the  $59^\circ$  beams at 1.7 ns with  $f_e$  of 0.05 (thick) and 0.1 (thin).



**Fig. 6.** (Color online) Variations of  $T_r$  (red star),  $\eta_L$  (blue triangle) and  $\eta_X$  (dark circle) as  $f_e$  at 1.7 ns.

it increases as  $f_e$  because of higher  $\eta_L$  and  $\eta_X$  at a larger  $f_e$ . Nevertheless, the variation of  $T_r$  is small, just within 3 eV when  $f_e \geq 0.05$ . As shown,  $T_r$  is about 188 eV from LARED-H with  $f_e = 0.07$ , which agrees with the experimental observations (Huser *et al.*, 2009).

#### 4. INFLUENCE OF $F_E$ ON THE X-RAY EMISSION REGION IN OVERCRITICAL PLASMAS

In this section, we study the influence of  $f_e$  on the motion of the X-ray emission region in the overcritical plasmas, which is often observed through a thinned wall hohlraum by a camera in experiments. Same as Huser *et al.* (2009) we also define the axial centroid position of emission intensity as:

$$z_I = \frac{\int I(z')z'dz'}{\int I(z')dz'}$$

where  $I(z)$  is the post-processed emission intensity for whole spectrum observed from the hohlraum axis along the radial direction. Moreover, we define the axial centroid velocity as  $v_I = dz_I/dt$ . In Figure 7, we present the simulated temporal relative displacements of  $z_I$  under different  $f_e$  for the empty hohlraum and compare them with the experimental observations. As shown, the emission region moves slower with a smaller  $f_e$ , which reason will be discussed below. From 1 ns to 2 ns, the experimental results are between the simulations with flux limiter of 0.05 and 0.07, with  $v_I$  being about  $55 \mu\text{m/ns}$ . However, after 2.0 ns when the main laser pulse falls rapidly,  $v_I$  is about  $205 \mu\text{m/ns}$  from our simulations, much higher than the experimental observations of  $55 \mu\text{m/ns}$ . We still cannot understand this big difference between the simulations and the observations.

In order to understand why the emission region moves slower with a smaller  $f_e$ , we present in Figure 8 the electron pressure  $P_e$  along the center line of the laser channel. As shown, there are two peaks of  $P_e$  at about  $760 \mu\text{m}$  and at

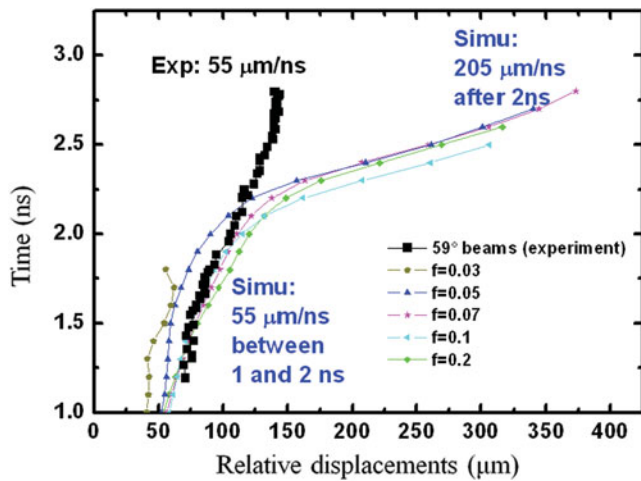


Fig. 7. (Color online) Comparisons of the simulated temporal relative displacements of  $z_l$  under different  $f_e$  with the experimental observations (Huser et al., 2009) for the empty hohlraum. Here,  $v_l$  in blue is obtained from simulations, in black is observed from experiments for the laser beams injected at  $59^\circ$ . Notice that the comparisons start from 1 ns in this figure and the following figures, because there is no experimental data before 1 ns.

about  $810 \mu\text{m}$ . The first peak at  $760 \mu\text{m}$  is caused by the laser ablation. As is known from Figure 2,  $T_e$  in the laser ablation region is higher at a smaller  $f_e$ , which therefore drives a higher laser ablation  $P_e$ . The second peak at about  $810 \mu\text{m}$  is caused by the radiation ablation. The radiation ablation  $P_e$  is mainly decided by  $T_r$ , while  $T_r$  changes little as  $f_e$  when  $f_e \geq 0.05$ , as discussed above. Hence, the radiation ablation  $P_e$  is also changes little as  $f_e$  when  $f_e \geq 0.05$ . The motion of the X-ray emission region in over-critical region is decided by the difference between the radiation ablation pressure and the laser ablation pressure. Because the pressure difference is smaller at a smaller  $f_e$ , so the motion of the X-ray emission region is slower. Nevertheless, we can see that the

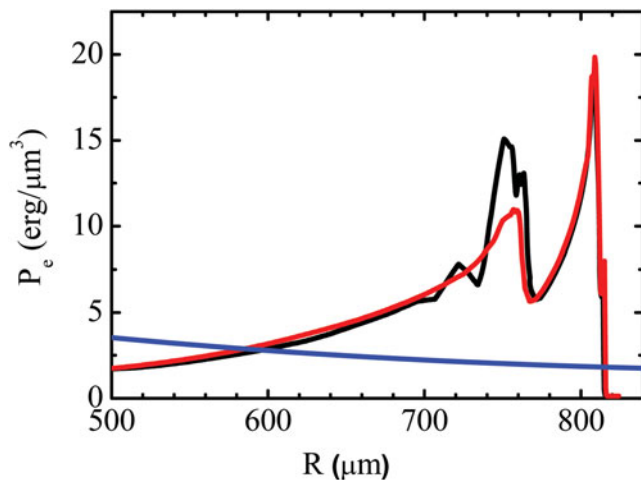


Fig. 8. (Color online) The distributions of  $P_e$  at 1.7 ns along the laser channel with  $f_e$  of 0.05 (black), 0.1 (red) and 0.2 (blue).

motion of the X-ray emission region in overcritical region is somewhat insensitive to  $f_e$  for this model.

## 5. INFLUENCE OF $F_E$ ON THE X-RAY EMISSION REGION IN PLASMA CORONA

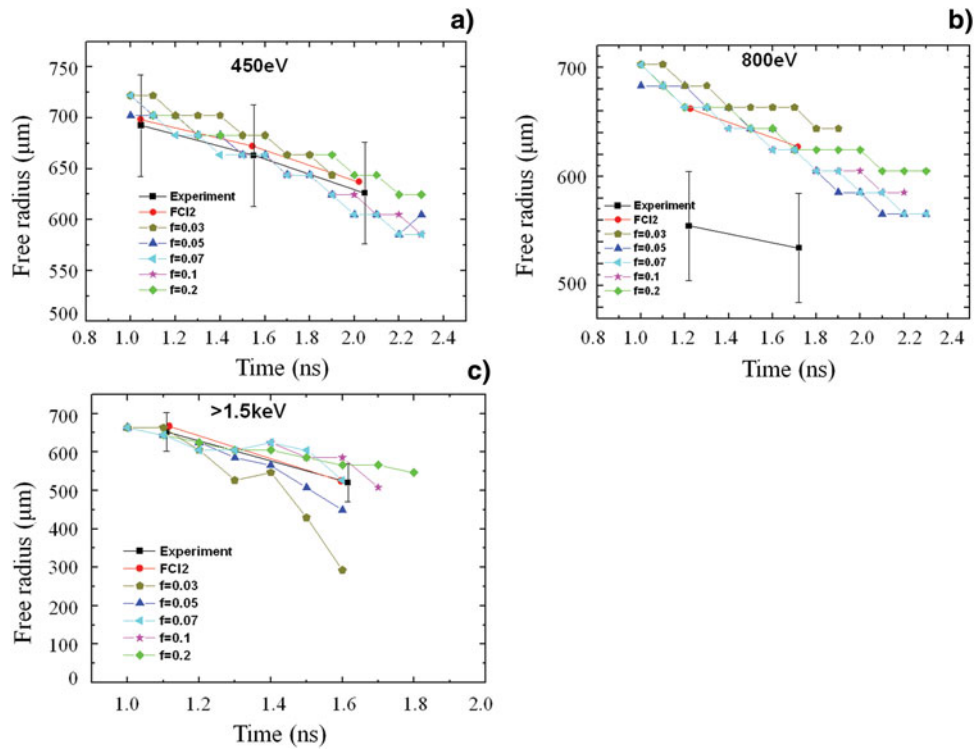
As the laser ablating, the plasma corona inside hohlraum expands toward to the axis. The on-axis imaging of the different X-ray bands were diagnosed by Huser et al. (2009). In their work, a free radius  $r_{free} = r(I = I_{center} + (I_{max} - I_{center})/2)$  is defined to describe the plasma expansion, where  $I_{center}$  is the residual intensity at the axis of the hohlraum and  $I_{max}$  is the maximum intensity along the radial direction. In order to compare with their observations, we also define such free radii, denoted as  $r_{free}^{O\text{-band}}$ ,  $r_{free}^{N\text{-band}}$  and  $r_{free}^{M\text{-band}}$  for O-band, N-band and M-band, respectively. The corresponding velocities are denoted as  $v_{free}^{O\text{-band}} = dr_{free}^{O\text{-band}}/dt$ ,  $v_{free}^{N\text{-band}} = dr_{free}^{N\text{-band}}/dt$ , and  $v_{free}^{M\text{-band}} = dr_{free}^{M\text{-band}}/dt$ .

We compare our simulations of the temporal  $r_{free}^{O\text{-band}}$ ,  $r_{free}^{N\text{-band}}$ , and  $r_{free}^{M\text{-band}}$  at different  $f_e$  with those from the experimental observations in Figure 9. The simulation results from FC12 (Huser et al., 2009) are also presented for comparison.

For O-band, as shown, all simulations from LARED-H with different  $f_e$  are located within the experimental error bars. For the N-band, however, the simulated expansion velocity from both LARED-H and FC12 are obviously faster than the experimental results. We cannot explain this disagreement, but we will check our simulations through a laser spot motion experiment to be done on SGIII-prototype in near future. As shown, the influence of  $f_e$  on  $r_{free}^{O\text{-band}}$  and  $r_{free}^{N\text{-band}}$  are not remarkable, because they are very near to  $R_c$  and  $R_c$  changes little as  $f_e$  from Figure 3. In other words, it is hard to decide  $f_e$  from  $r_{free}^{O\text{-band}}$  and  $r_{free}^{N\text{-band}}$  experimentally.

For M-band, however, the influence of  $f_e$  on  $r_{free}^{M\text{-band}}$  is more remarkable than for other two bands, and  $v_{free}^{M\text{-band}}$  is faster at a smaller  $f_e$ , as shown in Figure 9c. The reason is that the M-band is mainly emitted from the hot plasmas ablated by laser, while  $T_e$  of the hot plasmas is higher at a smaller  $f_e$ , as discussed in the last section. As a result, the expansion velocity of the M-band emission region is faster at a smaller  $f_e$ . As shown, the simulation results with  $f_e = 0.03$  and  $0.05$  are beyond the experimental error bars, and the best agreement generates at  $f_e = 0.07$ .

The variation of  $v_{free}$  as  $f_e$  is presented in Figure 10 for the three bands. As discussed above,  $v_{free}^{O\text{-band}}$  and  $v_{free}^{N\text{-band}}$  changes little with  $f_e$  while  $v_{free}^{M\text{-band}}$  obviously decreases as  $f_e$ . As presented,  $v_{free}^{M\text{-band}}$  drops from 390 to  $150 \mu\text{m/ns}$  when  $f_e$  is changed from 0.03 to 0.2, and  $v_{free}^{M\text{-band}}$  is  $230 \mu\text{m/ns}$  at  $f_e = 0.07$ . Moreover,  $v_{free}^{O\text{-band}}$  and  $v_{free}^{N\text{-band}}$  are much smaller than  $v_{free}^{M\text{-band}}$ , and the best agreement between the LARED-H simulations and the experimental observation is at  $f_e = 0.07$ . Due to the sensitivity of the motion of the M-band emission region to  $f_e$ , we therefore propose a method to experimentally determine  $f_e$  via the diagnostics of the temporal evolutions of the M-band emission region in plasma

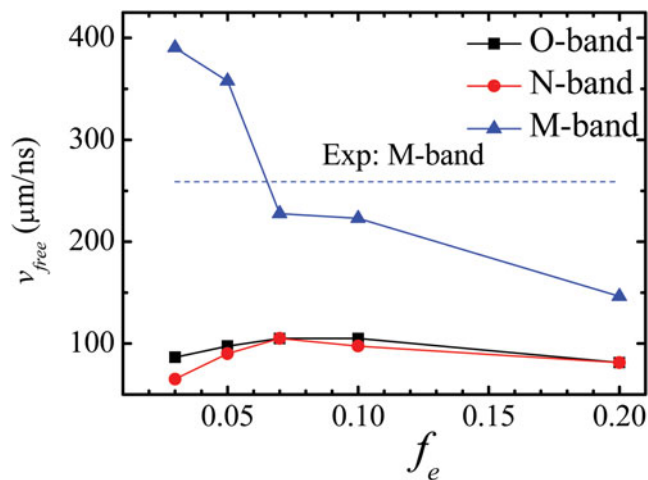


**Fig. 9.** (Color online) Comparisons of the temporal  $r_{free}^{O\text{-band}}$  (a),  $r_{free}^{N\text{-band}}$  (b) and  $r_{free}^{M\text{-band}}$  (c) in the empty hohlraum at different  $f_e$  with those from the experimental observations (with black square) and the simulations from FCI2 (with red circle) (Huser *et al.*, 2009). The free radii are postprocessed results from LARED-H simulations along the hohlraum axis.

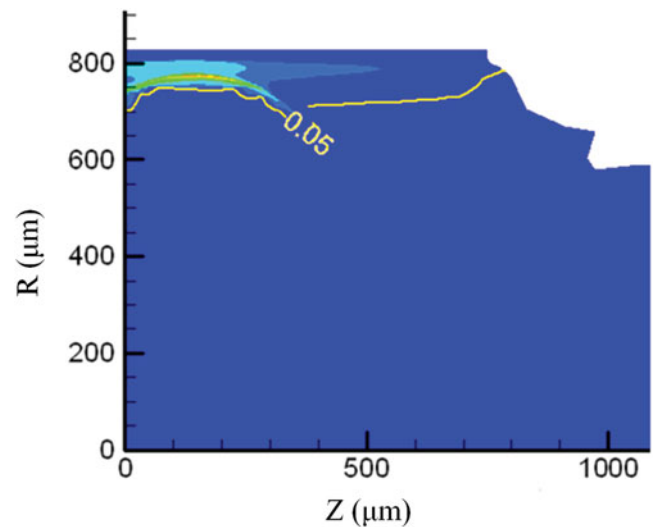
corona inside a Au hohlraum which has strong M-band emissions under drive conditions concerned in indirect-drive inertial fusion study.

It is also possible to measure the X-ray emission region in plasma corona from LEH, but a suitable angle should be chosen for observation. Presented in Figure 11 is the X-ray emission intensity map and the contour line of  $\rho = 0.05 \text{ g/cm}^3$  at the laser power peak time 1.7 ns. In the region

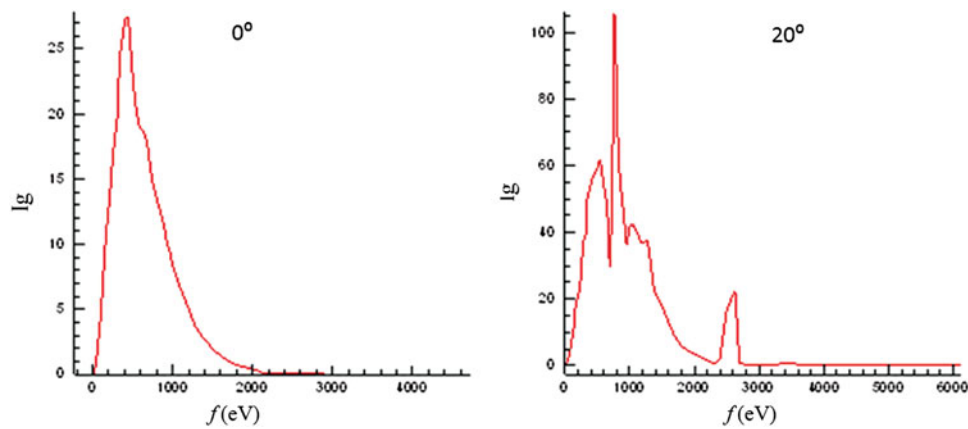
above the yellow line, there is  $\rho > 0.05 \text{ g/cm}^3$ . As shown, the maximum X-ray emission area is located at  $r \approx 770 \mu\text{m}$  and  $z \approx 200 \mu\text{m}$ . Between the X-ray emission region and the LEH, it is the radiation ablated plasmas, which are about  $400 \mu\text{m}$  in length with  $\rho = 0.05 \text{ g/cm}^3$  and  $T_r \approx 200 \text{ eV}$ . From our opacity model, the Rosseland mean free paths at  $T_r \approx 200 \text{ eV}$  and  $\rho_e = 0.05$  are  $56.55 \mu\text{m} \times 6.23 \mu\text{m}$  and  $46.50 \mu\text{m}$ , respectively,  $h\nu = 400 \text{ eV}$ ,  $800 \text{ eV}$ , and  $1.5$



**Fig. 10.** (Color online) Variations of  $v_{free}$  as  $f_e$  for O-band (black), N-band (red) and M-band (blue).



**Fig. 11.** (Color online) The X-ray emission intensity map and the contour line of  $\rho = 0.05 \text{ g/cm}^3$  at 1.7 ns, with  $f_e = 0.05$ .



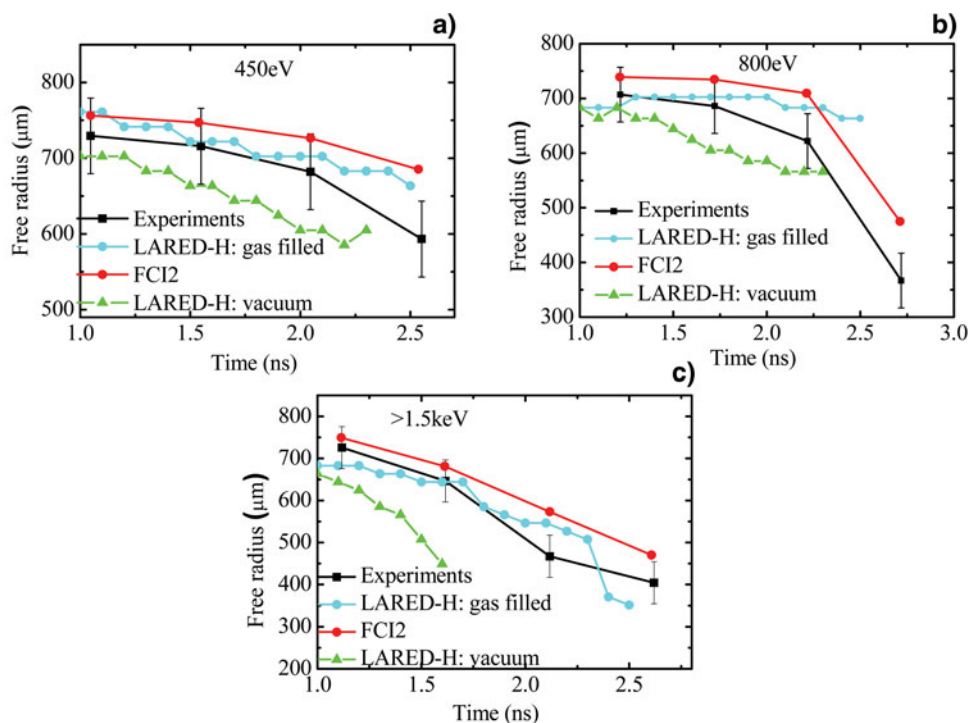
**Fig. 12.** (Color online) The postprocessed x-ray flux spectrum along the lines at 0° and 20° from the Hohlraum axis through LEH at 1.7 ns, with  $f_e = 0.05$ .

keV, which are much shorter than 400  $\mu\text{m}$ . Hence, the radiation ablated plasmas are optically thick. Presented in Figure 12 is the postprocessed X-ray flux spectrum along the line 0° and 20° from the Hohlraum axis through LEH. As shown, the spectrum observed at 0° through the optically thick plasmas is very near to the Planckian distribution, while at 20° the structures of O-band, N-band and M-band from the X-ray emission region near the laser deposited region can be clearly observed from the spectrum. Therefore, an observation angle should be carefully chosen to measure the emissions near the laser spot region through LEH. Together with the measurement of the X-ray emissions along the axial

direction, it is possible to obtain the motion of the X-ray emission region along the radial direction.

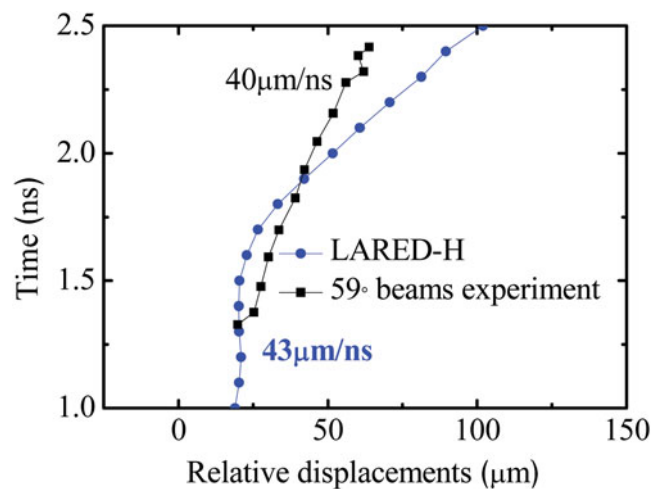
## 6. THE EFFECT OF GAS FILLING ON THE WALL PLASMA EXPANSION

To hold back the wall plasma expansion and to ensure that the laser beams propagate to positions near the hohlraum wall, a gas-filled hohlraum is usually used in inertial fusion study. In the experiments done by Huser *et al.* (2009) the authors also measured the wall and laser spot motion in a propane-filled hohlraum. In this section, we take  $f_e$  as 0.07



**Fig. 13.** (Color online) Comparisons of the temporal  $r_{free}^{\text{O-band}}$  (a),  $r_{free}^{\text{N-band}}$  (b), and  $r_{free}^{\text{M-band}}$  (c) in the gas-filled hohlraum at  $f_e = 0.07$  with those from the experimental observations (with black square) and the simulations from FCI2 (with red circle). The LARED-H results for the empty hohlraum are also presented for comparison.





**Fig. 14.** (Color online) Comparisons of the simulated temporal relative displacements  $z_l$  at  $f_e = 0.07$  with the experimental observations for the gas-filled hohlraum. Here,  $v_l$  in blue is obtained from simulations, in black are observed from experiments (Huser *et al.*, 2009) for the laser beams injected at  $59^\circ$ .

and compare our simulations with the observations (Huser *et al.*, 2009).

In Figure 13, we present the LARED-H simulations of the temporal  $r_{free}^{O\text{-band}}$ ,  $r_{free}^{N\text{-band}}$ , and  $r_{free}^{M\text{-band}}$  in the gas-filled hohlraum and compare them with the experimental results and the FCI2 simulations. For comparison, the simulations of the empty hohlraum are also presented. As shown, from the LARED-H simulations, the wall plasma expansion in the gas-filled hohlraum is much slower than that in the empty hohlraum. Obviously, the LARED-H simulations agree well with the experiments for all three bands when it is shorter than 2 ns, at which the main laser pulse falls around to its half maximum. However, there is some disagreement after 2.3 ns when the main pulse falls rapidly, especially for the N-band and the M-band. From Figure 13, compared to the experimental results and the FCI2 simulations, the N-band emission region after 2.3 ns moves much slower while the M-band moves much faster from the LARED-H.

In Figure 14, we compare the simulated temporal relative displacements  $z_l$  at  $f_e = 0.07$  with the experimental observations for the gas-filled hohlraum. As shown, the simulations from LARED-H agree with the observations in general, but there are some differences in detail. From the experiments,  $v_l$  is almost the same, around  $40 \mu\text{m/ns}$ , for whole observation time. However, from the LARED-H simulations,  $z_l$  has little change within the first 1.5 ns before the main laser pulse reaches its peak, then  $z_l$  moves with  $43 \mu\text{m/ns}$  between 1.5 ns and 1.8 ns, and with  $98 \mu\text{m/ns}$  after 1.8 ns when the laser pulse falls rapidly.

## SUMMARY

We have studied the influence of  $f_e$  on hohlraum plasmas by using the LARED-H simulations and proposed a method to

experimentally determine  $f_e$  via the motion of the M-band emission region in Au hohlraum. We have simulated the wall and laser spot motion experiments done by Huser *et al.* (2009) by taking different  $f_e$ . From our simulations, the limited free streaming flux may dominate the heat conduction in the regions with steep temperature gradient, which are important X-ray emission regions. As a result of our study, the motion of the M-band emission region is sensitive to the  $f_e$  when the limited free streaming flux dominates the heat conduction of the wall plasma expansion region, and one can determine  $f_e$  via the motion of the M-band emission region.

In general, the LARED-H simulations agree well with the experimental results when  $f_e$  is taken as 0.07 for both empty hohlraum and gas-filled hohlraum used in Huser's (2009) experiments. However, there are also some remarkable deviations from the observations, such as the motion of the N-band emissions in the overcritical region and the emission behaviors after 2.3 ns in both empty and gas-filled hohlraums. It is possible that we may have some problems in simulating the hohlraums when laser pulse falls rapidly and hohlraum becomes cold. To understand the disagreements between the simulations and the observations, we will have an experiment for measuring the motion of X-ray emissions on the SG-III prototype laser facility in near future.

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