

Elastic-plated gravity currents with a temperature-dependent viscosity

Clement Thorey^{1,†} and Chloé Michaut¹

¹Planetary and Space Science, Institut de Physique du Globe de Paris, UMR 7154 CNRS, F-75013 Paris, France

(Received 11 February 2016; revised 8 July 2016; accepted 12 August 2016; first published online 16 September 2016)

We develop a set of equations to explore the behaviour of cooling elastic-plated gravity currents for constant influx conditions. In particular, we introduce a temperature-dependent viscosity to couple the flow thermal structure with the velocity field. We show that this coupling results in important deviations from the isoviscous case. In particular, the bending and gravity asymptotic regimes, characteristic of the isoviscous case, both split into three different thermal phases: a first ‘hot’ isoviscous phase, a second phase where the spreading rate drastically decreases and the flow thickens and a third ‘cold’ isoviscous phase. The viscosity that controls the spreading rate differs in both asymptotic regimes; it is the average viscosity of a small peeling region at the current tip in the bending regime and the average flow viscosity in the gravity regime. In both regimes, we characterize the evolution of the thermal anomaly and determine the time scale of the phase changes in terms of the Péclet number and of the viscosity contrast. Finally, we show that the evolution with bending and gravity can result in six different evolution scenarios depending on the combination of dimensionless numbers considered. We provide a phase diagram which summarizes them as a function of the flow Péclet number and viscosity contrast.

Key words: gravity currents, lubrication theory, magma and lava flow

1. Introduction

Elastic-plated gravity currents involve the spreading of viscous material beneath an elastic sheet. The applications range from the emplacement of magma in the shallow crust (Bunger & Cruden 2011; Michaut 2011) and gravity-driven lava flows under a solidified crust at the surface (Slim *et al.* 2009; Hewitt, Balmforth & De Bruyn 2015) in geosciences to the manufacture of flexible electronics and microelectromechanical systems in engineering (Hosoi & Mahadevan 2004).

When the thickness of the flow is small compared to its extent, the lubrication approximation applies and the study of elastic-plated gravity currents amounts to the study of a sixth-order nonlinear diffusion equation for the flow thickness h (Michaut 2011; Lister, Peng & Neufeld 2013; Hewitt *et al.* 2015). Nonetheless, this equation is degenerate at the contact line, i.e. where $h \rightarrow 0$, and there cannot be any

[†] Email address for correspondence: clement.thorey@gmail.com

advancing solutions (Flitton & King 2004; Lister *et al.* 2013; Hewitt *et al.* 2015). This problem is similar to the well-known problem for surface-tension-driven flow where full continuum theories of fluid mechanics are unable to describe the flow near the contact line without introducing molecular-scale physics (Bertozzi 1998).

For practical purposes, different forms of regularization at the tip have been proposed such as the introduction of a thin prewetting film of fluid of initially constant thickness or the use of a fluid lag filled with gas at constant pressure (Lister *et al.* 2013; Hewitt *et al.* 2015; Peng *et al.* 2015; Pihler-Puzović *et al.* 2015). While both alternatives suffer from depending on the parameter introduced by the regularization itself, i.e. the thickness of the prewetting film or the pressure within the gap, the prewetting film regularization approach allows us to more easily introduce fluid cooling in the problem and is used here.

The dynamics of the spreading for a Newtonian fluid with a constant viscosity has been thoroughly described in an axisymmetric geometry (Michaut 2011; Lister *et al.* 2013; Thorey & Michaut 2014; Hewitt *et al.* 2015) and shows two distinct asymptotic regimes. First, gravity is negligible and the peeling of the front is driven by bending of the overlying layer; the interior is bell shaped, the radius evolves as $\tilde{h}_f^{1/22} t^{7/22}$ and the thickness at the centre as $\tilde{h}_f^{-2/22} t^{8/22}$ where \tilde{h}_f is the prewetting film thickness. When the radius becomes larger than $4\mathcal{L}$, where \mathcal{L} is the flexural wavelength of the upper elastic layer, the weight of the current becomes dominant over the bending terms and the flow enters a gravity current regime (Huppert 1982*b*). In this regime, the thickness profile develops a flat top with bent edges, the radius evolves as $t^{1/2}$ while the thickness tends to a constant. Different analogue experiments of isoviscous flows confirm these theoretical results (Dixon & Simpson 1987; Lister *et al.* 2013).

However, in most geological settings, the isothermal/isoviscous assumptions are not valid. For instance, the viscosity of magmas can vary by several orders of magnitude (Shaw 1972; Lejeune & Richet 1995). Therefore, as the fluid flows and cools down, its composition and crystal content change, which in turn modifies its viscosity and dynamics. Several studies have shown that, in a gravity current, this coupling between cooling and flow results in important deviations from the isoviscous case (Bercovici 1994; Bercovici & Lin 1996; Balmforth, Craster & Sassi 2004; Garel *et al.* 2014).

In this paper, we examine how the spreading of an elastic-plated gravity current is affected by its cooling. In particular, we account for the temperature dependence of the viscosity. This gives rise to a set of two coupled nonlinear equations that we solve numerically. We characterize the flow thermal structure and its effect on the dynamics via its rheology $\eta(T)$ in each regime separately. In both regimes, we identify different thermal phases of propagation that we characterize by different time scales and scaling laws.

2. Theory

2.1. Formulation

We model the axisymmetric flow of fluid below an elastic layer of constant thickness d_c and above a rigid layer (figure 1). To avoid problem at the contact line, we consider a thin, initially cold, prewetting film of thickness \tilde{h}_f and temperature T_0 (figure 1). The case of an initially hot prewetting film is treated in appendix C.

The hot fluid is injected continuously at the base and centre of the current at a constant rate Q_0 and constant temperature T_i through a conduit of diameter a .

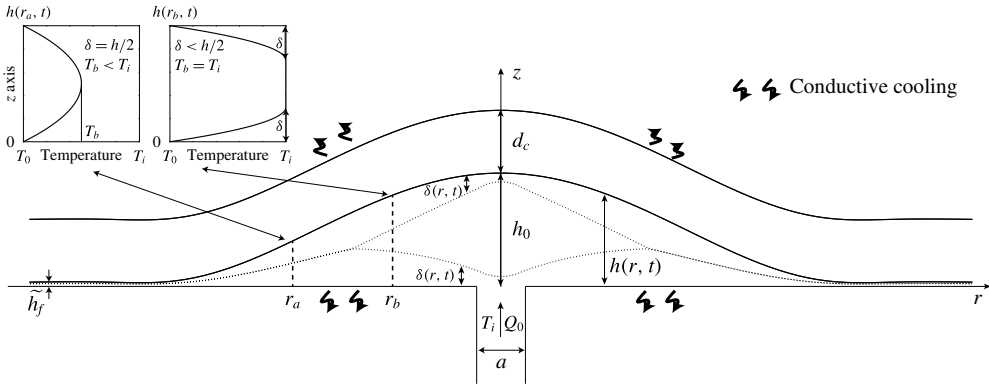


FIGURE 1. Model geometry and parameters. The vertical scale is exaggerated. Left upper panels: temperature profiles with merged ($r = r_a$) and separated ($r = r_b$) thermal boundary layers.

We assume a Poiseuille flow within the cylindrical feeding conduit such that the injection velocity w_i reads

$$w_i(r, t) = \mathcal{H} \left(\frac{a}{2} - r \right) \frac{\Delta P}{4\eta_h Z_c} \left(\frac{a^2}{4} - r^2 \right), \tag{2.1}$$

where $\Delta P/Z_c$ is the pressure gradient driving the flow in the feeding conduit and η_h is the viscosity of the hottest fluid at the temperature T_i . The fluid cools through the top and bottom boundaries by conduction in the surrounding medium, whose temperature is maintained constant and equal to T_0 . For simplicity, heating and melting of the surrounding medium are neglected.

As it cools, the viscosity of the fluid increases following an inverse dependence on the temperature

$$\eta(T) = \frac{\eta_h \eta_c (T_i - T_0)}{\eta_h (T_i - T_0) + (\eta_c - \eta_h)(T - T_0)}, \tag{2.2}$$

where η_c is the viscosity of the coldest fluid at $T = T_0$ (Bercovici 1994).

This rheology has the advantage of restricting strong viscosity variations over a small range of temperature close to T_0 while still capturing the essential behaviour of a viscous fluid, i.e. the viscosity variations are the largest where the temperature is the coldest (Shaw 1972; Marsh 1981; Lejeune & Richet 1995; Giordano, Russell & Dingwell 2008).

2.2. Pressure

The intrusion develops over a radius R that is much larger than its thickness h ($R \gg h$). In the laminar regime and in axisymmetrical coordinates (r, z) , the Navier–Stokes equations within the lubrication approximation are

$$-\frac{\partial P}{\partial r} + \frac{\partial}{\partial z} \left(\eta(T) \frac{\partial u}{\partial z} \right) = 0 \tag{2.3}$$

$$-\frac{\partial P}{\partial z} - \rho_m g = 0, \tag{2.4}$$

where $u(r, z, t)$ is the radial velocity, ρ_m the fluid density, g the standard acceleration due to gravity and $P(r, z, t)$ the pressure within the fluid. Integration of (2.4) gives the dynamic pressure $P(r, z, t)$ within the flow, which, given that the vertical deflection $h(r, t)$ of the upper elastic layer is small compared to its thickness d_c , i.e. $h \ll d_c$, can be written (Michaut 2011)

$$P = \rho_m g(h - z) + B \nabla_r^4 h, \tag{2.5}$$

with

$$\nabla_r^4 h = \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial}{\partial r} \left(\frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial h}{\partial r} \right) \right) \right), \tag{2.6}$$

where $h(r, t)$ is the flow thickness and B is the bending stiffness of the thin elastic layer, that depends on Young’s modulus E , Poisson’s ratio ν^* and on the elastic layer thickness d_c as $B = Ed_c^3 / (12(1 - \nu^{*2}))$.

2.3. Heat transport equation

2.3.1. Local energy conservation

In the laminar regime and in axisymmetrical coordinates (r, z) , the local energy conservation equation within the lubrication assumption is

$$\frac{D}{Dt} (\rho_m C_{p,m} T + \rho_m L(1 - \phi)) = k_m \frac{\partial^2 T}{\partial z^2}, \tag{2.7}$$

where $T(r, z, t)$ is the fluid temperature and ρ_m , k_m and $C_{p,m}$ are the density, thermal conductivity and specific heat of the fluid. Here, we also account for energy release by possible crystallization of the fluid, which is a non-negligible source of heat in the case of magmas; $\phi(r, z, t)$ is the crystal fraction in the melt and L the latent heat of crystallization. In this model, the crystals are only considered as a source/sink of energy as they melt/form at equilibrium during the flow as we assume that the physical properties of the crystal liquid mixture are the same as that of the fluid.

The fluid temperature varies between its liquidus T_L and solidus temperature T_S , i.e. $T_i = T_L$ and $T_0 = T_S$. As T_0 is also the fixed temperature of the surrounding medium in this model, this is equivalent to considering a fluid/wall interface pinned near the bulk freezing temperature T_S of the fluid. In making this assumption, we neglect the heat flux necessary to heat up the wall up to T_S .

Following a common approximation, we assume that the crystal fraction is a linear function of temperature over the melting interval (Hort 1997; Michaut & Jaupart 2006)

$$\phi = \frac{T_L - T}{T_L - T_S} = \frac{T_i - T}{T_i - T_0}. \tag{2.8}$$

With these approximations, the local energy equation (2.7) becomes

$$\frac{\partial T}{\partial t} + u \frac{\partial T}{\partial r} + w \frac{\partial T}{\partial z} = \frac{St}{St + 1} \kappa_m \frac{\partial^2 T}{\partial z^2}, \tag{2.9}$$

where $u(r, z, t)$ and $w(r, z, t)$ are the radial and vertical fluid velocities, $St = (C_{p,m}(T_i - T_0))/L$ is the Stefan number which represents the ratio of sensible heat

between solidus and liquidus to the total energy of the fluid at the liquidus temperature and κ_m is the fluid thermal diffusivity $\kappa_m = k_m / (\rho_m C_{p,m})$. In particular, equation (2.9) shows that considering the energy released by crystallization is simply equivalent to considering a reduced thermal diffusivity $\widetilde{\kappa}_m$, which reads

$$\widetilde{\kappa}_m = \frac{St}{St + 1} \kappa_m, \quad (2.10)$$

in the heat transport equation.

We use an integral balance method of heat transfer theory to approximately solve equation (2.9). In this method, that is developed below, the vertical structure of the temperature field is represented by a known function of depth that approximates the expected solution (Goodman 1958). Previous works on the cooling of lava domes at the surface have shown that such a reduction efficiently reduces the computation time while keeping the full dynamics of the unsimplified equation (2.9) well resolved (Balmforth *et al.* 2004). We use different kinds of vertical temperature profiles and show in appendix A that they all lead to very similar results.

2.3.2. Integral balance solution for the temperature $T(r, z, t)$

We model the cooling of the flow through the growth of two thermal boundary layers: one growing downward from the top and a second growing upward from the base. As we assume a fixed temperature at the boundary, the two thermal boundary layers grow symmetrically and have the same thickness $\delta(r, t)$ (figure 1). A popular approximation for the vertical temperature profile $T(r, z, t)$ is

$$T = \begin{cases} T_b - (T_b - T_0) \left(1 - \frac{z}{\delta}\right)^n & 0 \leq z \leq \delta \\ T_b & \delta \leq z \leq h - \delta \\ T_b - (T_b - T_0) \left(1 - \frac{h-z}{\delta}\right)^n & h - \delta \leq z \leq h, \end{cases} \quad (2.11)$$

where $T_b(r, t)$ is the temperature at the centre of the flow and $n > 1$ (Balmforth *et al.* 2004). This approximation captures the essential behaviour of the thermal structure: cooling is concentrated at the upper and bottom interfaces and cold boundary layers grow into the fluid interior as it flows. In addition, this profile assures the continuity of the temperature and heat flux within the flow.

While higher-order contributions to the temperature field, i.e. with $n > 2$, may also exist in certain situations, a parabolic profile is the most natural choice and we use $n = 2$ in the following (Bercovici 1994; Bercovici & Lin 1996). Nonetheless, the case of higher-order profiles are treated in appendix A and are shown not to influence the flow dynamics at least within the level of description adopted here.

While the integral balance solution in (2.11) depends on two variables T_b and δ that have to be consistently determined, these two variables are not independent. Indeed, in the flow region where thermal boundary layers exist, i.e. where $\delta < h/2$, the temperature T_b is equal to the injection temperature T_i and δ is the unknown (figure 1, $r = r_b$). In contrast, in the flow region where the thermal boundary layers have connected, which eventually arises at the current front, $\delta = h/2$ and T_b becomes the unknown (figure 1, $r = r_a$).

2.3.3. Integral balance equation

We integrate the local energy conservation equation (2.9) separately over the two thermal boundary layers. The integration over the bottom thermal layer, from $z=0$ to $z=\delta$ gives

$$\begin{aligned} & \frac{\partial}{\partial t}(\delta(\bar{T} - T_b)) + \frac{1}{r} \frac{\partial}{\partial r}(r\delta(\bar{uT} - \bar{u}T_b)) + \delta \left(\frac{\partial T_b}{\partial t} + \bar{u} \frac{\partial T_b}{\partial r} \right) \\ & = -\widetilde{\kappa}_m \left. \frac{\partial T}{\partial z} \right|_{z=0} + w_i(T_i - T_b), \end{aligned} \tag{2.12}$$

where the bar indicates a vertical average over the bottom thermal boundary layer

$$\bar{f} = \frac{1}{\delta} \int_0^\delta f \, dz, \tag{2.13}$$

$T_b(r, t)$ is the temperature at $z=\delta$ and we have used the nullity of the thermal gradient at $z=\delta$ and the local mass conservation

$$\frac{1}{r} \frac{\partial ru}{\partial r} + \frac{\partial w}{\partial z} = 0. \tag{2.14}$$

The integration over the top thermal layer, from $z=h-\delta$ to $z=h$ gives

$$\begin{aligned} & \frac{\partial}{\partial t}(\delta(\bar{T} - T_b)) + \frac{1}{r} \frac{\partial}{\partial r}(r\delta(\bar{uT} - \bar{u}T_b)) + \delta \left(\frac{\partial T_b}{\partial t} + \bar{u} \frac{\partial T_b}{\partial r} \right) \\ & = \widetilde{\kappa}_m \left. \frac{\partial T}{\partial z} \right|_{z=h}, \end{aligned} \tag{2.15}$$

where, in addition to (2.14) and the fact that the thermal gradient at $z=h-\delta$ is equal to zero, we have used the kinematic boundary condition at $z=h(r, t)$

$$\frac{\partial h}{\partial t} + u \frac{\partial h}{\partial r} = w. \tag{2.16}$$

Adding (2.12) and (2.15) and using (2.11) with $n=2$ to derive the conductive fluxes, we finally obtain the heat balance equation which reads

$$\begin{aligned} & \frac{\partial}{\partial t}(\delta(\bar{T} - T_b)) + \frac{1}{r} \frac{\partial}{\partial r}(r\delta(\bar{uT} - \bar{u}T_b)) + \delta \left(\frac{\partial T_b}{\partial t} + \bar{u} \frac{\partial T_b}{\partial r} \right) \\ & = -2\widetilde{\kappa}_m \frac{(T_b - T_0)}{\delta} + \frac{w_i}{2}(T_i - T_b). \end{aligned} \tag{2.17}$$

2.3.4. Final heat transport equation

Equation (2.17) can be rewritten depending on the unknown in the temperature field δ or T_b .

When $T_b = T_i$ and δ is the variable, equation (2.17) becomes

$$\frac{\partial}{\partial t}(\delta(\bar{T} - T_i)) + \frac{1}{r} \frac{\partial}{\partial r}(r\delta(\bar{uT} - \bar{u}T_i)) = -2\widetilde{\kappa}_m \frac{(T_i - T_0)}{\delta}. \tag{2.18}$$

When $\delta = h/2$ and T_b is the variable, equation (2.17) becomes

$$\begin{aligned} \frac{\partial}{\partial t}(h(\bar{T} - T_b)) + \frac{1}{r} \frac{\partial}{\partial r}(rh(\overline{uT} - \bar{u}T_b)) + h \left(\frac{\partial T_b}{\partial t} + \bar{u} \frac{\partial T_b}{\partial r} \right) \\ = -8\widetilde{\kappa}_m \frac{(T_b - T_0)}{h} + w_i(T_i - T_b). \end{aligned} \tag{2.19}$$

Expanding the derivatives and using a global statement of mass conservation, equation (2.19) simplifies and reads

$$\frac{\partial}{\partial t}(h\bar{T}) + \frac{1}{r} \frac{\partial}{\partial r}(rhu\bar{T}) = -8\widetilde{\kappa}_m \frac{(T_b - T_0)}{h} + w_i T_i. \tag{2.20}$$

Interestingly, equation (2.20) is a particular case of (2.18) if the boundary layers have merged and $\delta = h/2$. We can then use (2.18) as a simplification of (2.17) to describe the heat transport within the flow in both cases.

We finally rewrite (2.18) using a new variable $\xi = \delta(T_i - \bar{T})$ which encloses both unknowns δ and T_b since \bar{T} depends on T_b following $\bar{T} = (2T_b + T_0)/3$

$$\frac{\partial \xi}{\partial t} + \frac{1}{r} \frac{\partial}{\partial r}(r\bar{u}\xi) - \frac{1}{r} \frac{\partial}{\partial r}(r\delta(\overline{uT} - \bar{u}\bar{T})) = 2\widetilde{\kappa}_m \frac{(T_i - T_0)}{\delta}. \tag{2.21}$$

The second term on the left-hand side of (2.21) contains advection by the vertically integrated radial velocity profile while the third term contains a correction accounting for the vertical structure of the temperature field. The term on the right is the loss of heat by conduction in the surrounding medium.

In addition, this formulation in terms of a unique variable ξ allows us to calculate T_b or δ directly from the expression of ξ using either $\delta = h/2$ or $T_b = T_i$ respectively

$$T_b(r) = \begin{cases} T_i & \text{if } \xi \leq \xi_t \\ \frac{3T_i - T_0}{2} - \frac{3\xi}{h} & \text{if } \xi > \xi_t \end{cases} \quad \delta(r) = \begin{cases} 3\xi & \text{if } \xi \leq \xi_t \\ h(r, t)/2 & \text{if } \xi > \xi_t, \end{cases} \tag{2.22a,b}$$

with $\xi_t = h(T_i - T_0)/6$.

2.4. Equation of motion

To obtain an equation for the flow thickness, we first note that, since the boundary conditions are the same at $z=0$ and $z=h$ and the viscosity and velocity u possess the same symmetry, the vertical structure of the temperature field (2.11) is symmetric around $h/2$. Taking advantage of this symmetry, we integrate once (2.3) using $\partial u/\partial z|_{z=h/2} = 0$ to get

$$\frac{\partial u}{\partial z} = \frac{1}{\eta(r, z, t)} \frac{\partial P}{\partial r} \left(z - \frac{h}{2} \right), \tag{2.23}$$

where $\eta(r, z, t)$ is given by (2.2). Using no-slip boundary conditions at the top and bottom of the flow, the global mass conservation can be rewritten as

$$\frac{\partial h}{\partial t} = \frac{1}{r} \frac{\partial}{\partial r} \left(r \int_0^h \frac{\partial u}{\partial z} z \, dz \right) + w_i, \tag{2.24}$$

and therefore, inserting (2.23) into (2.24) and using (2.5), we obtain the equation for the flow thickness evolution

$$\frac{\partial h}{\partial t} = \frac{1}{r} \frac{\partial}{\partial r} \left(r \left(\rho_m g \frac{\partial h}{\partial r} + B \frac{\partial}{\partial r} (\nabla_r^4 h) \right) I_1(h) \right) + w_i, \tag{2.25}$$

where

$$I_1(z) = \int_0^z \frac{1}{\eta(r, y, t)} \left(y - \frac{h}{2} \right) y \, dy. \tag{2.26}$$

2.5. Average quantities

Solving the equations of motion (2.25) and heat transport (2.21) requires determining the average quantities \bar{u} , \bar{uT} and $\bar{u\bar{T}}$.

Integration of (2.23) using the no-slip boundary condition at $z=0$ gives

$$u(r, z, t) = \frac{\partial P}{\partial r} I_0(z), \tag{2.27}$$

where

$$I_0(z) = \int_0^z \frac{1}{\eta(r, y, t)} \left(y - \frac{h}{2} \right) \, dy. \tag{2.28}$$

The average velocity over a thermal boundary layer \bar{u} then reads

$$\begin{aligned} \bar{u} &= \frac{1}{\delta} \int_0^\delta u \, dz = u(r, \delta, t) - \frac{1}{\delta} \int_0^\delta \frac{\partial u}{\partial z} z \, dz \\ &= \frac{1}{\delta} \frac{\partial P}{\partial r} (\delta I_0(\delta) - I_1(\delta)), \end{aligned} \tag{2.29}$$

where $P(r, z, t)$ is given by (2.5) and we have used (2.23).

The average rate of heat advected \bar{uT} over a thermal boundary layer reads

$$\begin{aligned} \bar{uT} &= \frac{1}{\delta} \int_0^\delta uT \, dz = \frac{1}{\delta} \left([uG(z)]_0^\delta - \int_0^\delta G(z) \frac{\partial u}{\partial z} \, dz \right) \\ &= \frac{1}{\delta} \frac{\partial P}{\partial r} \left(G(\delta) I_0(\delta) - \int_0^\delta \frac{1}{\eta(r, y, t)} \left(y - \frac{h}{2} \right) G(y) \, dy \right), \end{aligned} \tag{2.30}$$

where

$$G(z) = T_0 z + \frac{z^2}{\delta} (T_b - T_0) - \frac{z^3}{3\delta^2} (T_b - T_0) \tag{2.31}$$

is an antiderivative of $T(z)$ for $0 \leq z \leq \delta$. More conveniently, equation (2.30) can be rewritten in terms of $I_1(z)$ and a new integral $I_2(z)$ as

$$\bar{uT} = \frac{1}{\delta} \int_0^\delta uT \, dz = \frac{1}{\delta} \frac{\partial P}{\partial r} (G(\delta) I_0(\delta) - T_0 I_1(\delta) - I_2(\delta)), \tag{2.32}$$

where

$$I_2(z) = \int_0^z \frac{1}{\eta(r, y, t)} \left(y - \frac{h}{2} \right) \frac{y^2}{3\delta^2} (3\delta - y) (T_b - T_0) \, dy. \tag{2.33}$$

Therefore, using (2.29), (2.32) and $\bar{T} = (2T_b + T_0)/3$, we finally get

$$\bar{uT} - \bar{u}\bar{T} = \frac{1}{\delta} \frac{\partial P}{\partial r} \left(\frac{2}{3} (T_b - T_0) I_1(\delta) - I_2(\delta) \right). \tag{2.34}$$

2.6. Dimensionless equations

We adopt $T = T_0 + (T_i - T_0)\theta$ to obtain a dimensionless temperature $\theta(r, z, t)$, i.e. the injection temperature T_i is scaled to 1 and the far-field temperature T_0 to 0. The dimensionless integral balance approximation (2.11) with $n = 2$ becomes

$$\theta(z) = \begin{cases} \Theta_b \left(1 - \left(1 - \frac{z}{\delta}\right)^2\right) & 0 \leq z \leq \delta \\ \Theta_b & \delta \leq z \leq h - \delta \\ \Theta_b \left(1 - \left(1 - \frac{h-z}{\delta}\right)^2\right) & h - \delta \leq z \leq h, \end{cases} \tag{2.35}$$

where $\Theta_b = (T_b - T_0)/(T_i - T_0)$. Equations (2.21) and (2.25) are non-dimensionalized using a horizontal scale \mathcal{L} , a vertical scale \mathcal{H} and a time scale \mathcal{T} given by

$$\mathcal{L} = \left(\frac{B}{\rho_m g}\right)^{1/4} \tag{2.36}$$

$$\mathcal{H} = \left(\frac{12\eta_h Q_0}{\rho_m g \pi}\right)^{1/4} \tag{2.37}$$

$$\mathcal{T} = \frac{\pi \mathcal{L}^2 \mathcal{H}}{Q_0}, \tag{2.38}$$

where \mathcal{L} represents the flexural wavelength of the upper elastic layer (Michaut 2011), \mathcal{H} the characteristic thickness of an isoviscous constant flux gravity current with viscosity η_h (Huppert 1982a) and \mathcal{T} the characteristic time to fill up a cylindrical flow of radius \mathcal{L} and thickness \mathcal{H} at a constant rate Q_0 . In addition, we can define a horizontal velocity scale $\mathcal{U} = \mathcal{L}/\mathcal{T} = (\rho_m g \mathcal{H}^3)/(12\eta_h \mathcal{L})$ and a pressure scale $\rho_m g \mathcal{H}$.

The dimensionless variable ξ reads $\xi = \delta(1 - \theta)$ and the dimensionless model summarizes as followed

$$\frac{\partial h}{\partial t} - \frac{12}{r} \frac{\partial}{\partial r} \left(r I_1(h) \frac{\partial P}{\partial r}\right) = \mathcal{H} \left(\frac{\gamma}{2} - r\right) \frac{32}{\gamma^2} \left(\frac{1}{4} - \frac{r^2}{\gamma^2}\right) \tag{2.39}$$

$$\frac{\partial \xi}{\partial t} + \frac{1}{r} \frac{\partial}{\partial r} (r \bar{u} \xi - \Sigma) = 2Pe^{-1} \frac{\Theta_b}{\delta}, \tag{2.40}$$

with

$$\Theta_b(r) = \begin{cases} 1 & \text{if } \xi \leq \xi_t = \frac{h}{6} \\ \frac{3}{2} - \frac{3\xi}{h} & \text{if } \xi > \xi_t = \frac{h}{6} \end{cases} \quad \delta(r) = \begin{cases} 3\xi & \text{if } \xi \leq \xi_t = \frac{h}{6} \\ h(r, t)/2 & \text{if } \xi > \xi_t = \frac{h}{6} \end{cases} \tag{2.41a,b}$$

$$\bar{u} = \frac{12}{\delta} \frac{\partial P}{\partial r} (\delta I_0(\delta) - I_1(\delta)) \tag{2.42}$$

$$\Sigma = \frac{\partial P}{\partial r} (8I_1(\delta)\Theta_b - 12I_2(\delta)) \tag{2.43}$$

and

$$I_0(z) = \int_0^z (\nu + (1 - \nu)\theta(r, y, t)) \left(y - \frac{h}{2}\right) dy \tag{2.44}$$

$$I_1(z) = \int_0^z (v + (1 - v)\theta(r, y, t)) \left(y - \frac{h}{2}\right) y \, dy \tag{2.45}$$

$$I_2(z) = \int_0^z (v + (1 - v)\theta(r, y, t)) \left(y - \frac{h}{2}\right) \frac{y^2}{3\delta} (3\delta - y) \, dy, \tag{2.46}$$

where $P = h + \nabla_r^4 h$ is the dimensionless dynamic pressure, γ the dimensionless size of the conduit is set to 0.02 in the following (Michaut 2011), \mathcal{H} is the Heaviside function and h_f , Pe and ν are the three dimensionless numbers that control the dynamics of the flow

$$h_f = \frac{\tilde{h}_f}{\mathcal{H}} \tag{2.47}$$

$$Pe = \frac{\mathcal{H}^2}{\tilde{\kappa}_m \mathcal{T}} \tag{2.48}$$

$$\nu = \frac{\eta_h}{\eta_c}, \tag{2.49}$$

h_f is the dimensionless thickness of the prewetting film; it is the sole parameter governing the evolution of the dimensionless isoviscous current (Lister *et al.* 2013). For practical purposes, we use $h_f = 5 \times 10^{-3}$ in the main text and show results and scalings for variable h_f in appendix C. Pe is a reduced Péclet number which incorporates the energy released by crystallization; it compares the vertical diffusion of heat to the horizontal advection in the interior. Finally, ν is the viscosity contrast, i.e. the ratio between the hottest and coldest viscosity.

The expression of $I_0(\delta)$, $I_1(h)$, $I_1(\delta)$ and $I_2(\delta)$, as well as the numerical scheme used to solve equations (2.39) and (2.40), are given in appendix B for $n = 2$.

2.7. Basic behaviour of an isothermal flow

For a constant injection rate, a small prewetting film thickness, i.e. $h_f \ll 1$ and a viscosity contrast ν set to 1, the numerical resolution of (2.39) shows the two classical asymptotic spreading regimes that were previously described by Michaut (2011), Lister *et al.* (2013) and Hewitt *et al.* (2015).

At early times, when $R \ll \mathcal{L}$, gravity is negligible and the interior has a uniform pressure $P = \nabla_r^4 h$. The flow is bell shaped and its thickness is given by

$$h(r, t) = h_0(t) \left(1 - \frac{r^2}{R^2(t)}\right)^2, \tag{2.50}$$

with $h_0(t)$ the flow thickness at the centre. In this regime, Lister *et al.* (2013) have shown that the spreading is controlled by the propagation of a peeling by bending wave at the flow front whose velocity c , which critically depends on the flow viscosity η , reads

$$c = \frac{dR}{dt} = \frac{B\tilde{h}_f^{1/2}}{12\eta} \left(\frac{\kappa}{1.35}\right)^{5/2} \tag{2.51}$$

in dimensional form, where κ is the curvature of the interior solution. Using $\eta = \eta_h$ in (2.51), the dimensionless flow radius and height are given by (figure 2)

$$h_0(t) = 0.65h_f^{-1/11} t^{8/22}, \tag{2.52}$$

$$R(t) = 2.13h_f^{1/22} t^{7/22}. \tag{2.53}$$

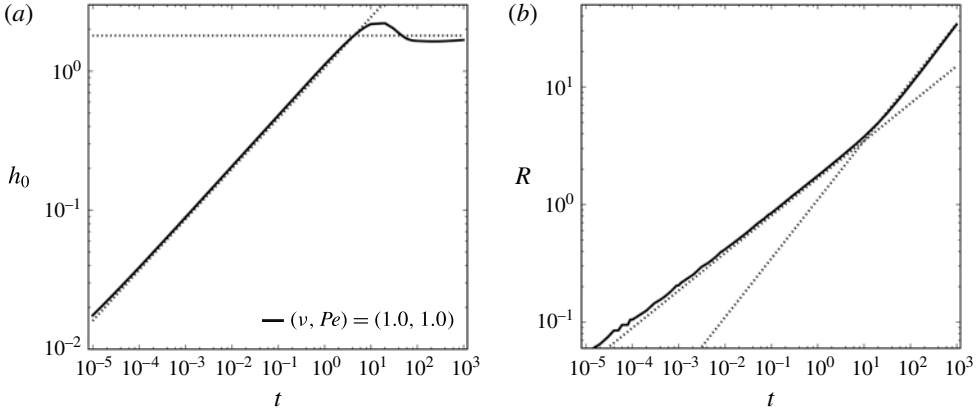


FIGURE 2. (a) Dimensionless thickness at the centre h_0 versus dimensionless time t . Dotted lines: scaling laws in the bending regime $h_0 = 0.65h_f^{-1/11}t^{8/22}$ and in the gravity regime where h_0 tends to a constant. (b) Dimensionless radius R versus dimensionless time t . Dotted lines: scaling laws in the bending regime $R = 2.13h_f^{1/22}t^{7/22}$ and in the gravity current regime $R = 1.10t^{1/2}$.

In contrast, when the radius R becomes larger than $4\mathcal{L}$ ($R \gg \mathcal{L}$), the fluid weight becomes the dominant pressure contribution, i.e. $P = h$ and the current enters a classical gravity current regime where the dimensional radius is given by

$$R(t) = 0.715 \left(\frac{\rho g Q_0^3}{12\eta} \right)^{1/8} t^{1/2} \tag{2.54}$$

and the thickness h_0 tends to a constant (Huppert 1982b; Michaut 2011; Lister *et al.* 2013). Taking $\eta = \eta_h$ in (2.54), the dimensionless radius $R(t)$ is given by (figure 2)

$$R(t) = 1.10t^{1/2}. \tag{2.55}$$

In the following, we study the effect of cooling on the flow dynamics in both regimes separately.

3. Evolution in the bending regime

We first concentrate on the case in which only bending contributes to the pressure. We then numerically solve and study the system composed by (2.39) and (2.40) where we remove the gravitational contribution in the pressure P , i.e. $P = \nabla_r^4 h$. The thin film of fluid is initially cold.

3.1. Qualitative description

3.1.1. Thermal structure for an isoviscous flow, effect of Pe

The current cools by conduction and thermal boundary layers form at the contact with the surrounding medium. These boundary layers first connect at the tip of the flow, where the small thickness induces an important cooling (figure 3). A region of cold fluid forms at the front and slowly grows with time.

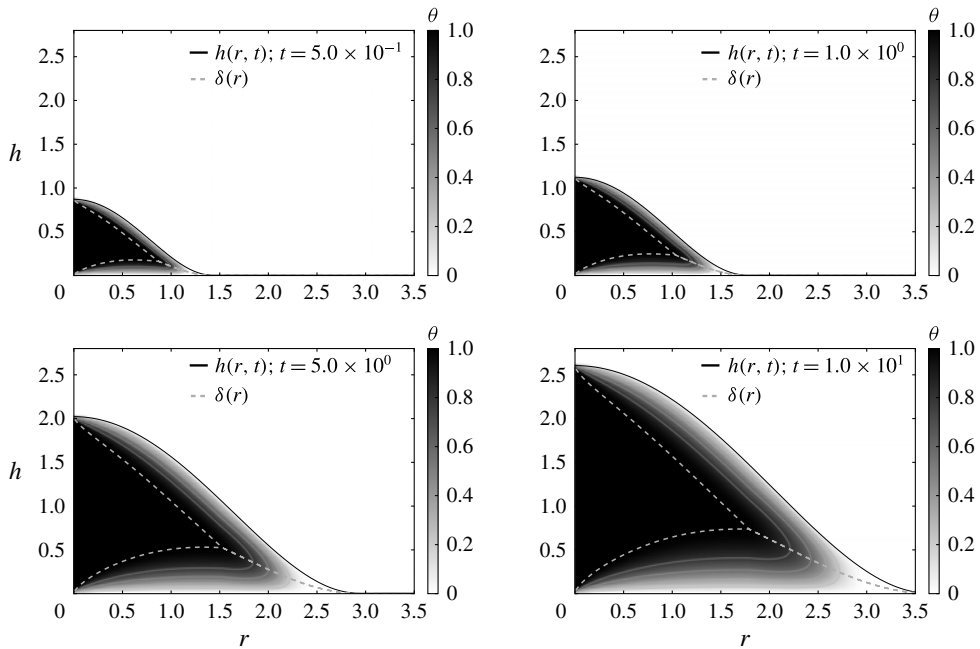


FIGURE 3. Snapshots of the flow thermal structure $\theta(r, z, t)$ at different times indicated on the plot. Dashed lines represent the thermal boundary layers. Solid grey lines are isotherms for $\theta = 0.2, 0.4, 0.6$ and 0.8 . Here, the temperature field is decoupled from the dynamics, i.e. $\nu = 1$ and $Pe = 100$.

The flow is composed of a hot core, that contains a significant amount of heat and is lately referred to as the flow thermal anomaly, and a cold front. The radius $R_c(t)$ of the thermal anomaly is defined as the radius where $\Theta_b = 0.01$. As the current thickens with time, a balance between advection and diffusion of heat is never reached in the current and the extent of the thermal anomaly grows with time. However, it spreads slower than the current itself and the extent of the cold fluid region at the tip, i.e. the region defined by $(R - R_c)(t)$, grows. For instance, for $Pe = 100$, while the region of cold fluid extends over approximately 10% of the current at $t = 0.5$, it extends over approximately 20% at $t = 10$ (figure 3). The smaller the number Pe , the more important the conductive cooling and the larger the cold region is (figure 4).

3.1.2. Thickness and temperature profile, effect of ν

When accounting for the temperature dependence of the viscosity, the region of cold fluid at the tip is marked by a higher viscosity which enhances flow thickening at the expense of spreading. The larger the viscosity contrast, the larger the aspect ratio h_0/R (figure 4). For instance, for the same value of $Pe = 1$, while the aspect ratio is 0.7 for $\nu = 1$ at $t = 10$, it is 4.2 at the same time for $\nu = 10^{-3}$ (figure 4). Nevertheless, the shape of the flow remains essentially self-similar, i.e. well described by (2.50) and cannot be differentiated from the shape of an isoviscous current if the thickness and the radial coordinates are rescaled by the thickness at the centre $h_0(t)$ and radius $R(t)$ (figure 5).

While the flow thermal structure is similar to the isoviscous case (figure 4), the important thickening induced by the viscosity increase tends to limit heat loss to the

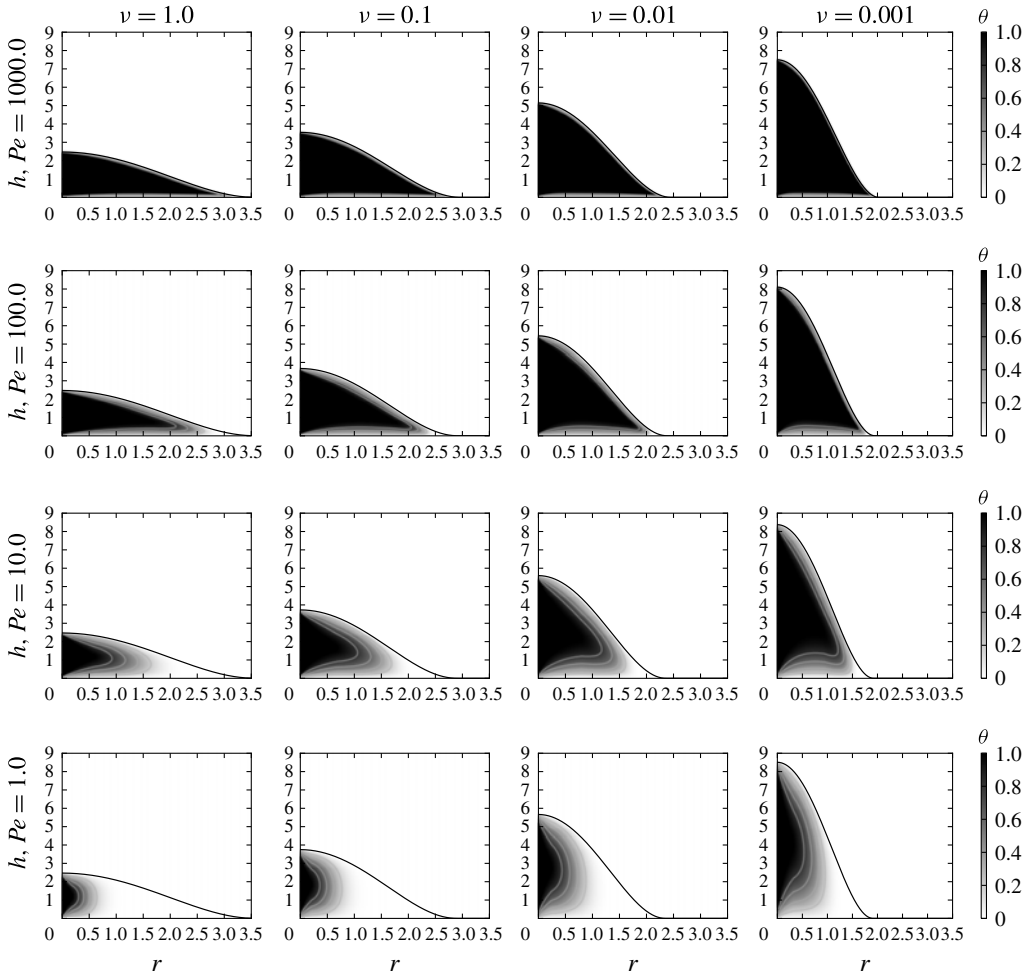


FIGURE 4. Snapshots of the flow thermal structure $\theta(r, z, t)$ for different set (ν, Pe) with $\nu = 1, 0.1, 0.01$ and 0.001 and $Pe = 1, 10, 100$ and 1000 at $t = 10$. While Pe controls the thermal structure of the flow, it has only a small influence on the flow aspect ratio which is controlled by ν .

surrounding and to increase the size of the thermal anomaly at a given time. For instance, for $Pe = 1$ at $t = 10$, while the thermal anomaly extends over approximately 30% of the flow for $\nu = 1$, it extends over more than 50% for $\nu = 10^{-3}$ (figure 4).

As expected, a larger Péclet number leads to a larger thermal anomaly (figure 4). However, although different Péclet numbers cause very different thermal structures, the influence of the Péclet number on the flow morphology is small, much smaller than the effect of the viscosity contrast ν (figure 4). For instance, at $t = 10$ for $\nu = 10^{-3}$, the thermal anomaly is still attached to the tip of the current for $Pe = 1000$ whereas it makes approximately 50% of the current for $Pe = 1$; but, the thickness h_0 and the radius R in both cases differ only by a few per cent (figure 4). This confirms that, in this regime, the spreading of the flow is not controlled by the mean temperature or average viscosity of the flow, but by the local dynamics at the flow front, as suggested by Lister *et al.* (2013).

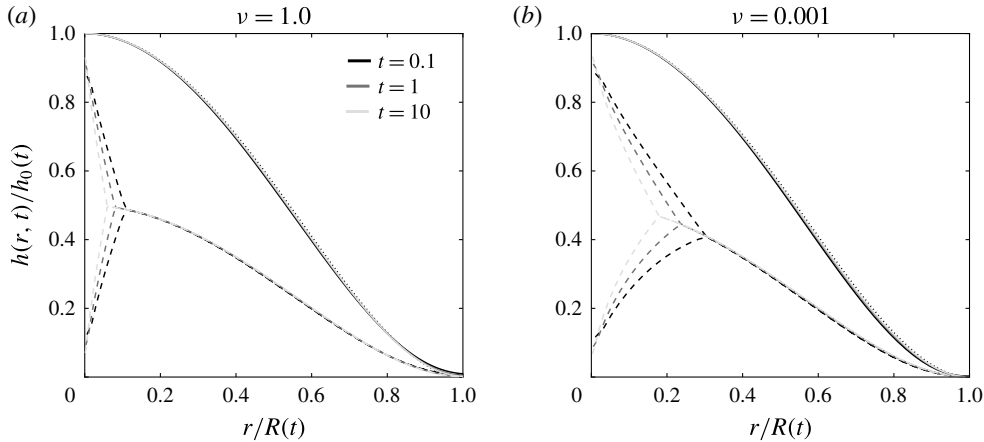


FIGURE 5. (a) Thickness normalized by the thickness at the centre $h(r, t)/h_0(t)$ versus radial axis normalized by the current radius $r/R(t)$ at different times indicated on the plot for $Pe = 1.0$ and $\nu = 1.0$. The solid lines represent the thickness profiles and the dashed lines represent the thermal boundary layers. The predicted morphology (2.50) (dotted line) is also plotted for comparison. (b) Same plot but for $\nu = 10^{-3}$.

3.2. Quantitative description

3.2.1. Evolution of the thickness and the radius

In this bending dominated regime, the dynamics shows three different spreading phases. The thickness as well as the radius first evolve like an isoviscous flow, i.e. $h_0 \propto t^{8/22}$ (2.52) and $R \propto t^{7/22}$ (2.53) (figure 6). In a second phase, thickening occurs at the expense of spreading and the dynamics deviates from the isoviscous case. Finally, the current returns to an isoviscous-like dynamics but offset from the hot isoviscous scaling law by a factor that depends on the viscosity contrast (figure 6).

These different propagation phases thus reflect variations in the spreading rate of the current. As described in § 2.7, the propagation speed of the front is governed by local conditions in the peeling region and is given in dimensional form by (2.51); it decreases when the viscosity η increases.

Inserting the cold viscosity η_c in place of η into (2.51), we find that the dimensionless thickness and radius for a cold isoviscous current evolve as

$$h_0 = 0.65\nu^{-2/11}h_f^{-1/11}t^{8/22}, \quad (3.1)$$

$$R = 2.13\nu^{1/11}h_f^{1/22}t^{7/22}. \quad (3.2)$$

Both (3.1) and (3.2) fit very well the third asymptotic phase of the dynamics (figure 6).

3.2.2. Flow effective viscosity

The effective viscosity of the current, i.e. the dimensionless viscosity η_e which controls the current spreading rate in (2.51), is obtained by substituting ν by $1/\eta_e(t)$ in (3.1) and reads

$$\eta_e(t) = \left(\frac{h_0(t)t^{-8/22}}{0.65h_f^{-1/11}} \right)^{11/2}, \quad (3.3)$$

where $h_0(t)$ is given by our simulations.

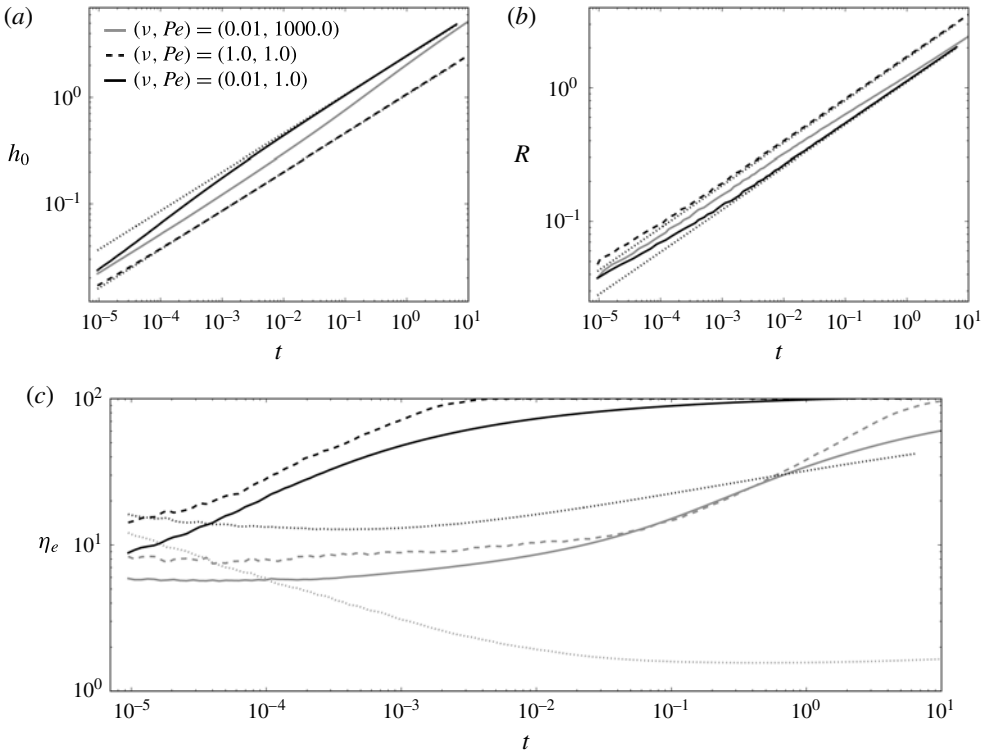


FIGURE 6. (a) Dimensionless thickness at the centre h_0 versus dimensionless time t for different sets (ν, Pe) indicated on the plot. Dotted lines: scaling laws $h_0 = 0.65h_f^{-1/11}\nu^{-2/11}t^{8/22}$ for $\nu = 1.0$ and 0.01 . (b) Dimensionless radius R versus dimensionless time t for the same sets (ν, Pe) . Dotted lines: scaling laws $R = 2.13h_f^{1/22}\nu^{1/11}t^{7/22}$ for $\nu = 1.0$ and 0.01 . (c) Dimensionless effective viscosity versus dimensionless time t for different Pe and $\nu = 0.01$. Solid lines: dimensionless effective viscosity η_e defined by (3.3). Dotted lines: dimensionless average flow viscosity defined by $\overline{\eta_a(t)} = (1/V(t)) \int_0^{R(t)} \int_0^{h(r,t)} r\eta(\theta) dr dz$ where $V(t)$ is the current volume. Dashed lines: dimensionless average front viscosity η_f defined by (3.4).

As suggested by the results of § 3.2.1, the effective viscosity is first low (figure 6c). It rapidly increases in the second phase of propagation and finally tends to the cold viscosity $1/\nu$ in the third phase. The effective viscosity is however much larger than the average flow viscosity (figure 6c).

We calculate the average viscosity $\eta_f(t)$ over a front region of size L in between $R(t) - L$ and $R(t)$

$$\eta_f = \frac{1}{V_f} \int_{R-L}^R \int_0^h r\eta(\theta) dr dz, \tag{3.4}$$

where $V_f(t)$ is the volume of this region. The numerical evaluation of $\eta_f(t)$ for $L = 4L_p(t)$, where L_p is the peeling length scale defined by Lister *et al.* (2013), given by

$$L_p = 1.08h_f^{13/22}\eta_e^{-2/11}t^{3/22}, \tag{3.5}$$

fits relatively well the behaviour of the effective viscosity η_e (figure 6c).

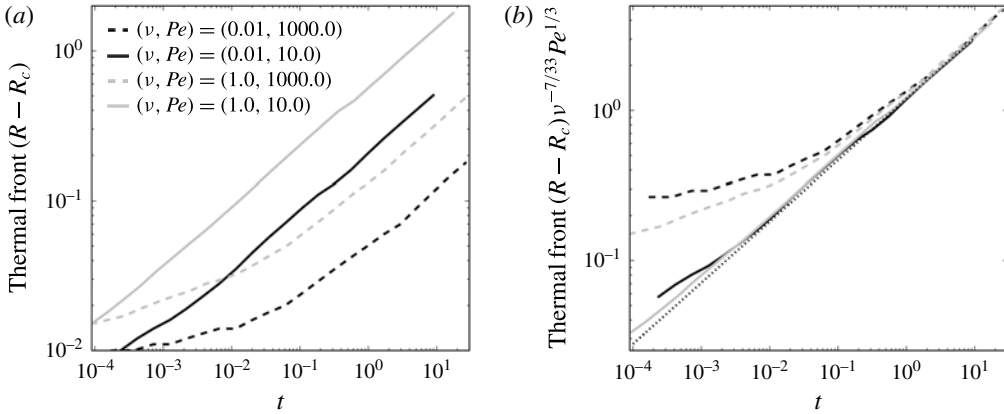


FIGURE 7. (a) Extent of the cold fluid region $R(t) - R_c(t)$ versus dimensionless time for different combinations (ν, Pe) indicated on the plot. (b) Same plot but where we rescale the extent of the cold fluid region by $Pe^{-1/3}\nu^{7/33}$. Dotted line: scaling law $(R(t) - R_c(t))Pe^{1/3}\nu^{-7/33} = 2.1h_f^{7/66}t^{9/22}$.

Therefore, the effective viscosity, and thus the spreading rate, are controlled by the average viscosity of a small region at the front of size $L = O(L_p)$. Note that in the first flow phase, the effective viscosity η_e , and thus the average viscosity of the peeling region, is larger than 1 as the current has to accommodate the cold fluid initially present in the prewetting film. The effective viscosity is initially equal to 1 in the case of an initially hot prewetting film (see appendix C).

Average local thermal conditions in the peeling region thus control the spreading rate in the bending regime.

3.2.3. Evolution of the thermal condition at the tip

The thermal anomaly is first advected at the same velocity as the current itself, i.e. $R(t) = R_c(t)$ (figure 7a). After a time that depends on Pe and ν , the flow front leaves the thermal anomaly behind and $R(t) - R_c(t)$ increases with time (figure 7).

In the bending regime, the interior pressure is constant and the thickness profile $h(r, t)$ is given by (2.50) (figure 5). The radius of the thermal anomaly $R_c(t)$ is theoretically taken as corresponding to the radius in the flow where heat advection locally balances heat loss, i.e.

$$\frac{d}{dt}(\Theta_b h) \sim Pe^{-1} \frac{\Theta_b}{h}. \tag{3.6}$$

Using the thickness profile (2.50), (3.6) becomes

$$\alpha^2 \left(1 + \frac{R_c}{R}\right)^2 \left(\Theta_b \frac{dh_0}{dt} + h_0 \frac{d\Theta_b}{dt}\right) + \frac{4h_0 R_c^2 \Theta_b}{R^3} \frac{dR}{dt} \alpha \left(1 + \frac{R_c}{R}\right) \sim \frac{Pe^{-1} \Theta_b}{\alpha^2 \left(1 + \frac{R_c}{R}\right)^2 h_0}, \tag{3.7}$$

where $\alpha(t) = (R(t) - R_c(t))/R(t)$ is the normalized region beyond $r = R_c(t)$. In the limit $\alpha \ll 1$, i.e. $R_c/R \sim 1$, the time derivative is locally dominated by its advective

part ($\alpha\alpha$) and we finally get

$$\alpha^3 \sim \frac{Pe^{-1}}{h_0^2(t)} \frac{R}{\frac{\partial R}{\partial t}}. \quad (3.8)$$

Substituting $h_0(t)$ and $R(t)$ by their respective scaling laws (3.1) and (3.2), the size evolution of the normalized cold front region α reads

$$\alpha(t) \sim Pe^{-1/3} \nu^{4/33} h_f^{2/33} t^{1/11}, \quad (3.9)$$

which is equivalent to

$$R(t) - R_c(t) \sim Pe^{-1/3} \nu^{7/33} h_f^{7/66} t^{9/22}. \quad (3.10)$$

The predicted scaling law for the evolution of the cold fluid region (3.10) closely fits the numerical simulations for $\nu < 1$ and for different Péclet numbers (figure 7). In particular, this scaling perfectly fits the simulations with a numerical prefactor equal to 2.1.

3.2.4. Summary of the bending regime dynamics

The spreading rate of the current is controlled by the average viscosity in the peeling region which depends on the average local thermal conditions at the tip.

At the initiation of the flow, heat advection is larger than conductive heat losses in the peeling region, the thermal anomaly reaches the flow front and the spreading rate is similar to the one of an isoviscous hot current.

Once heat loss in the peeling region becomes larger than heat advection, the current tip leaves the thermal anomaly behind. The average temperature of the peeling region decreases and the effective viscosity rapidly increases. Setting $(R - R_c)(t)$ (3.10) equal to the peeling length scale $L_p(t)$ (3.5), taking $\nu = 1$ and inverting for time, we found that the time t_{b2} to reach the second thermal phase reads

$$t_{b2} = 7 \times 10^{-3} Pe^{11/9} h_f^{16/9}, \quad (3.11)$$

where the numerical prefactor is chosen from the numerical solutions. When rescaling the time of the simulations by t_{b2} , the different simulations enter the second phase simultaneously (figure 8a).

Finally, when the peeling region becomes entirely cold, i.e. $R(t) - R_c(t) \gg L_p(t)$, the flow behaves as an isoviscous cold current. Repeating the same exercise as before while keeping the viscosity contrast ν , we found that the time t_{b3} for the flow to enter this third phase reads

$$t_{b3} = 14Pe^{11/9} h_f^{16/9} \nu^{-1/9}, \quad (3.12)$$

where the numerical prefactor is chosen from the numerical solutions. When rescaling the time of the simulations by t_{b3} , the different simulations enter the third phase simultaneously (figure 8b).

4. Evolution in the gravity current regime

To study the late time behaviour, we concentrate on the case where only the weight of the fluid contributes to the pressure. We numerically solve and study the system composed by (2.39) and (2.40) where we remove the bending contribution in the pressure P , i.e. $P = h$. We follow the same framework as in § 3.

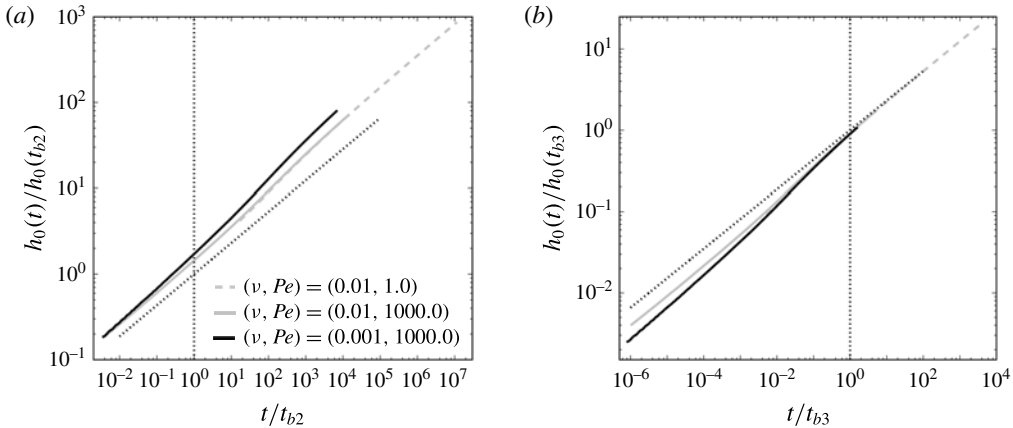


FIGURE 8. (a) Dimensionless thickness at the centre $h_0/h_0(t_{b2})$ versus dimensionless time t/t_{b2} , where t_{b2} is the time to enter the second thermal phase of the bending regime (3.11), for different sets (ν, Pe) indicated on the plot. Dotted line: scaling laws $h_0(t)/h_0(t_{b2}) = (t/t_{b2})^{8/22}$. (b) Dimensionless thickness at the centre $h_0/h_0(t_{b3})$ versus dimensionless time t/t_{b3} , where t_{b3} is the time to enter the third thermal phase of the bending regime (3.12), for different sets (ν, Pe) indicated on the plot. Dotted line: scaling laws $h_0(t)/h_0(t_{b3}) = (t/t_{b3})^{8/22}$.

4.1. Qualitative description

4.1.1. Thermal structure for an isoviscous flow, effect of Pe

As in the bending regime, the bulk of the fluid first expands at the injection temperature and $R_c(t) = R(t)$. As the bottom and top cool by conduction, thermal boundary layers form at the contact with the surrounding medium and connect at the tip of the current. However, in the gravity current regime, for a constant viscosity, the thickness of the current tends to a constant. Therefore, when conduction in the surrounding medium balances the input of heat at the centre and the flow front has already left the thermal anomaly behind, its extent approaches a steady state (figure 9).

The radius of the steady-state thermal anomaly R_c also largely depends on Pe in this regime: the larger the number Pe , the larger the radius R_c (figure 10).

4.1.2. Thickness and temperature profile, effect of ν

For a current with a viscosity that depends on temperature, as soon as the cooling becomes more effective and the thermal anomaly detaches from the current tip, the spreading slows down and flow thickening is enhanced (figure 10). For instance, for $Pe = 1$, while the aspect ratio h_0/R is approximately 0.12 for $\nu = 1$ at $t = 200$, it is ~ 1 for $\nu = 10^{-3}$ (figure 10). The shape of the current is not self-similar and the front steepens when the viscosity increases in comparison to the isoviscous case as noted by Bercovici (1994). However, when the current becomes much larger than the thermal anomaly, the current side slumps to become less steep (figure 10) and recovers a shape similar to an isoviscous flow with a cold viscosity.

The thermal structure is similar to the isoviscous case ($\nu = 1$). In particular, after a time that depends on Pe , the thermal anomaly approaches a steady-state profile (figure 10). As in the bending regime, the thickening at the centre limits heat loss to the surrounding for large values of the viscosity contrast ν . Therefore, the extent of the thermal anomaly in the steady state is slightly larger for a larger viscosity contrast. For instance, for $Pe = 10$ at $t = 200$, while the thermal anomaly extends over less than 2 for $\nu = 1$, it reaches ~ 3 for $\nu = 10^{-3}$.

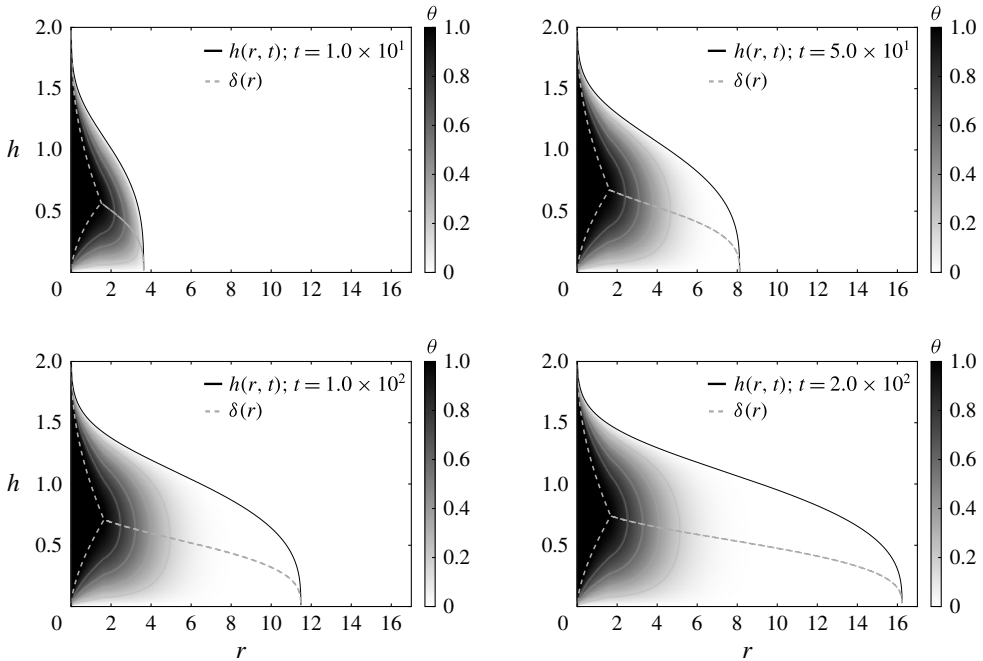


FIGURE 9. Snapshots of the flow thermal structure $\theta(r, z, t)$ at different times indicated on the plot. Dashed lines: thermal boundary layers. Here, $\nu = 1$ and $Pe = 100$.

The flow morphology is much more sensitive to Pe in the gravity current regime than in the bending regime and different Pe lead to different current morphologies for a given ν (figure 10). For instance, for $\nu = 10^{-3}$ at $t = 200$, the thermal anomaly is still attached to the tip for $Pe = 10^3$ and the aspect ratio of the flow h_0/R is close to 0.15. In contrast, for $Pe = 1$, the thermal anomaly radius is less than 30% of the current radius and the dimensionless aspect ratio of the flow is much larger $h_0/R = 1.15$ (figure 10).

4.2. Quantitative description

4.2.1. Evolution of the thickness and radius

As in the bending regime, the dynamics in the gravity current regime shows three different spreading phases. The thickness as well as the radius first follow the isoviscous scaling laws for a given hot viscosity η_h , i.e. h_0 approaches a constant and $R \propto t^{1/2}$ (figure 11). In a second phase, the thickness rapidly increases and the spreading slows down. Finally, the current returns to an isoviscous-like dynamics but offset from the hot isoviscous scaling law by a factor that depends on the viscosity contrast (figure 11). In particular, replacing η by η_c instead of η_h in (2.54), we find that the dimensionless radius $R(t)$ in the third flow phase matches the one of a cold viscosity current (4.1) (figure 11)

$$R(t) = 1.10\nu^{1/8}t^{1/2}. \quad (4.1)$$

As in the bending regime, these spreading rate variations reflect variations in the flow effective viscosity, from the hot to the cold viscosity. However, in that case, the

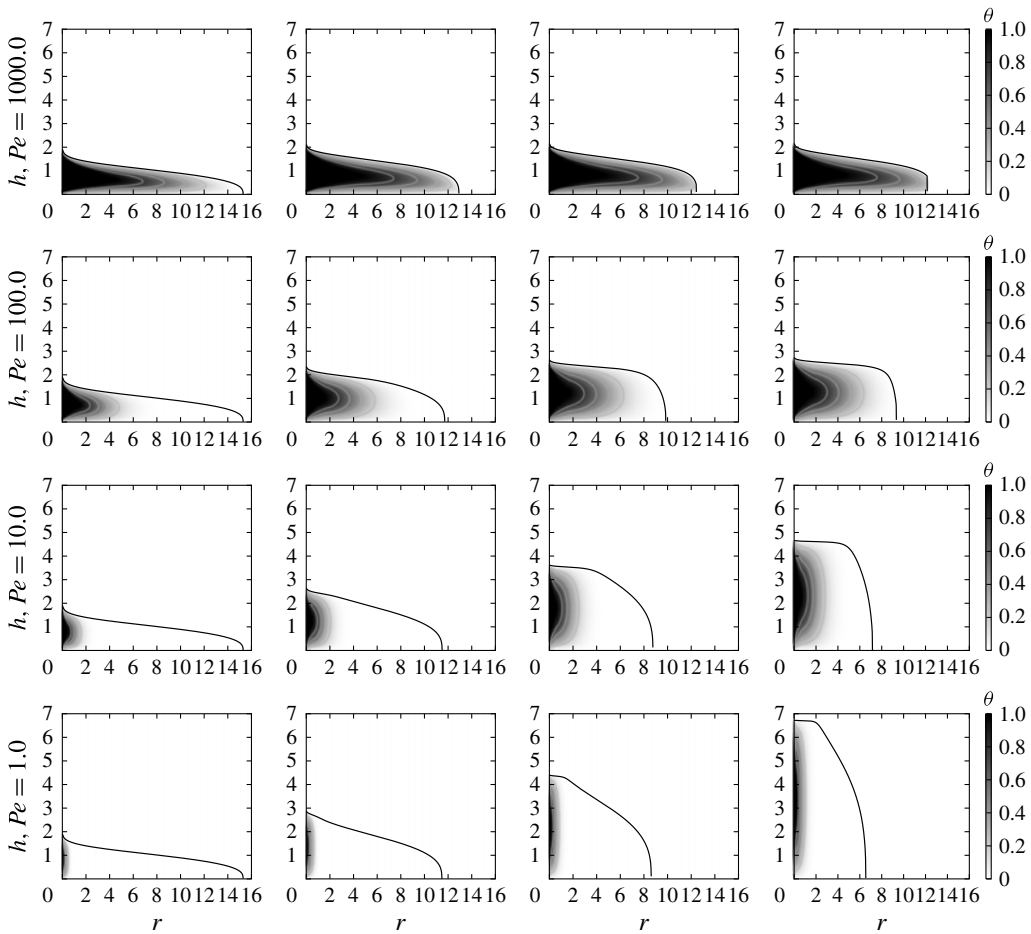


FIGURE 10. Snapshots of the flow thermal structure $\theta(r, z, t)$ for different sets (ν, Pe) with $\nu = 1, 0.1, 0.01$ and 0.001 and $Pe = 1, 10, 100$ and 1000 at $t = 200$.

effective viscosity is the average flow viscosity. Using the predicted scaling law for the radius $R(t)$ (4.1) as a function of ν_e to derive the effective viscosity $1/\nu_e$, we show that the current average viscosity matches the effective flow viscosity almost perfectly (figure 11c).

4.2.2. Characterization of the thermal anomaly

The thermal anomaly is first advected at the same velocity as the current itself, i.e. $R_c(t)/R(t) = 1$ (figure 12a). After a time that depends on Pe and ν , the flow front leaves the thermal anomaly behind and the anomaly extent slowly approaches a steady-state profile as the viscosity becomes closer to the cold viscosity (figures 9, 10 and 12).

At the steady-state radius R_c of the thermal anomaly, a balance between heat advection and diffusion in the surrounding medium gives

$$\Theta_b \frac{U_0}{R_c} \sim \frac{Pe^{-1}}{h_0^2} \Theta_b, \quad (4.2)$$

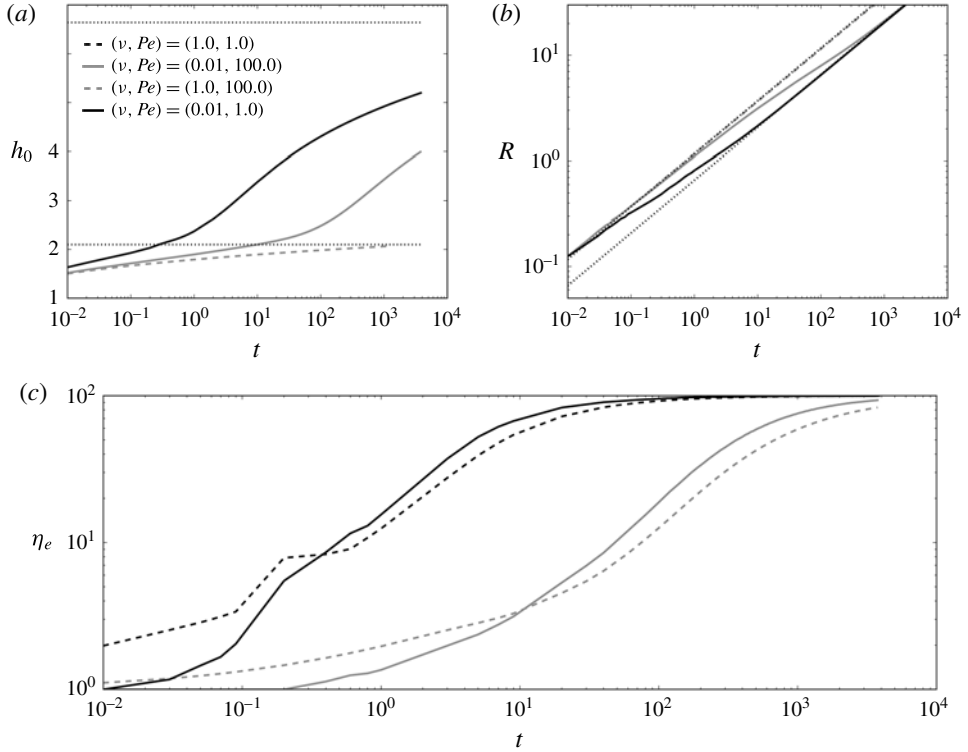


FIGURE 11. (a) Dimensionless thickness at the centre h_0 versus dimensionless time t for different sets (ν, Pe) indicated on the plot. Dotted lines represent the scaling laws $h_0 = 2.1\nu^{-1/4}$ for $\nu = 1.0$ and 10^{-2} . (b) Dimensionless radius R versus dimensionless time t for the same sets (ν, Pe) . Dotted lines represent the scaling laws $R = 1.1\nu^{1/8}t^{1/2}$ for $\nu = 1.0$ and 10^{-2} . (c) Dimensionless effective viscosity versus dimensionless time t for different Pe and $\nu = 0.01$. Solid lines: effective viscosity $\eta_e(t) = (R(t)/1.1t^{1/2})^{-8}$. Dashed lines: average flow viscosity defined by $\overline{\eta_a(t)} = (1/V(t)) \int_0^{R(t)} \int_0^{h(r,t)} r\eta(\theta) dr dz$ where $V(t)$ is the current volume. Note that the irregularities at early times, which are enhanced by the log–log representation, come from a slight decrease of the effective viscosity when the current begins to thicken.

where U_0 is a mean dimensionless velocity of advection. For a gravity current, in contrast to the bending regime, the thickness h_0 approaches a constant. Taking U_0 as a horizontal redistribution of the dimensionless injection rate at $r = R_c$, i.e. $U_0 = 1/(R_ch_0)$, and using the scaling for the thickness of a cold isoviscous current $h_0 \sim \nu^{-1/4}$, we obtain

$$R_c \sim Pe^{1/2}\nu^{-1/8} \tag{4.3}$$

and hence

$$\frac{R_c}{R(t)} \sim Pe^{1/2}\nu^{-1/4}t^{-1/2}, \tag{4.4}$$

where we have used (4.1).

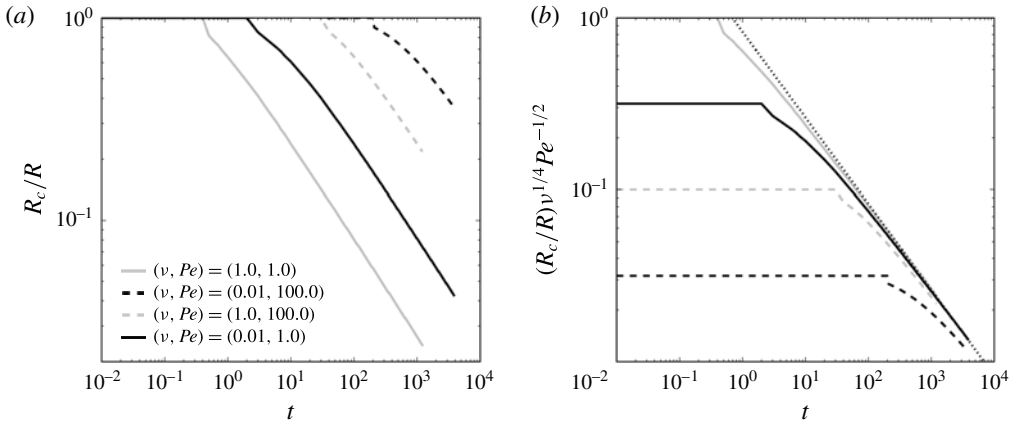


FIGURE 12. (a) Normalized thermal anomaly radius $R_c(t)/R(t)$ versus dimensionless time for different combinations (ν, Pe) indicated on the plot. (b) Same plot but where we rescale the normalized thermal anomaly radius $R_c(t)/R(t)$ by $Pe^{1/2}\nu^{-1/4}$. Dotted line: $Pe^{-1/2}\nu^{1/4}R_c(t)/R(t) = 0.8t^{-1/2}$.

This scaling law fits our simulation for a prefactor equal to 0.8 (figure 12): when the thermal anomaly enters the steady state, its radius remains constant and the normalized radius $R_c(t)/R(t)$ evolves as the inverse of the current radius, i.e. as $t^{-1/2}$ (figure 12). Furthermore, both the dependence with Pe and ν vanish when rescaling $R_c/R(t)$ by $Pe^{1/2}\nu^{-1/4}$ in the steady state (figure 12b).

4.2.3. Summary of the gravity regime dynamics

In the gravity regime, the current spreading rate is controlled by the flow average viscosity. Therefore, it strongly depends on the extent of the thermal anomaly.

At flow initiation, the thermal anomaly is advected at the same velocity as the current itself and the current spreads with a hot viscosity η_h . Once heat loss over the current balances the heat input at the centre, the current tip leaves the thermal anomaly behind, the average viscosity increases and the spreading rate decreases. The time t_{g2} to enter this second phase scales with the time to cool the hot current ($\nu = 1$) by conduction and reads

$$t_{g2} = 0.65Pe, \quad (4.5)$$

where the numerical prefactor is chosen from the numerical solutions. Indeed, when rescaling the time of the simulations by t_{g2} , the different simulations enter the second phase simultaneously (figure 13a).

Finally, when the thermal anomaly becomes small compared to the current, i.e. $R_c/R \ll 1$, the average flow temperature is close to zero and the current behaves as an isoviscous cold current. The time t_{g3} to enter this third phase scales as the time to cool an isoviscous cold gravity current and reads

$$t_{g3} = 10Pe\nu^{-1/2}, \quad (4.6)$$

where the numerical prefactor is chosen from the numerical solutions. Indeed, when rescaling the time of the simulations by t_{g3} , the different simulations enter the third phase simultaneously (figure 13b). The time scales t_{g2} (respectively t_{g3}) can be derived by inverting (4.4) for t for a fixed value of R_c/R and setting $\nu = 1$ (respectively keeping ν) in a similar way that we derive t_{b2} (respectively t_{b3}) in § 3.2.4.

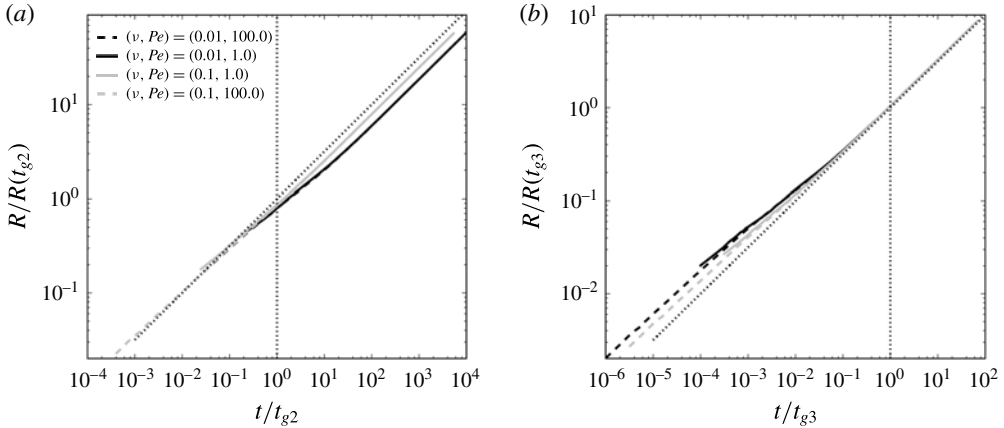


FIGURE 13. (a) Dimensionless radius at the centre $R/R(t_{g2})$ versus dimensionless time t/t_{g2} , where t_{g2} (4.5) is the time to enter the second thermal phase of the gravity regime, for different sets of (ν, Pe) indicated on the plot. Dotted line: scaling law $R/R(t_{g2}) = (t/t_{g2})^{1/2}$. (b) Dimensionless radius $R/R(t_{g3})$ versus dimensionless time t/t_{g3} , where t_{g3} (4.6) is the time to enter the third thermal phase of the gravity regime, for different sets (ν, Pe) indicated on the plot. Dotted line: scaling law $R/R(t_{g3}) = (t/t_{g3})^{1/2}$.

5. Different evolutions with bending and gravity

For an isoviscous flow with $h_f \ll h \ll d_c$, in between the bending and gravity regime, Lister *et al.* (2013) also describe a short intermediate regime where the peeling by bending continues to control the propagation but where the flow shows an interior flat-topped region due to the increasing effect of gravity. For simplicity, we only consider the two asymptotic regimes. At the transition, the isoviscous current is characterized by $R \sim 4$ and for $h_f = 0.005$, $h_0 \sim 2$ and $t \sim 10$.

For a temperature-dependent viscosity current and for any values of h_f , ν and Pe , the current always transitions to the gravity regime when $R \sim 4$ (figure 14). However, the current thickness as well as the time to reach this transition naturally depends on the bending thermal phase of the current at the transition, i.e. on the combination of (ν, Pe) considered. For instance, for $\nu = 0.01$ and a small value of $Pe = 1.0$, the current transitions to the gravity regime while in the third bending thermal phase. The transition occurs much later and the current is much thicker than for an isoviscous current ($t_i \sim 50$ and $h_0(t_i) \sim 8$, figure 14). In contrast, for $\nu = 0.01$ and a large value of $Pe = 10^5$, the current spreads in the first thermal bending phase for a longer period of time; it reaches the transition sooner at a smaller thickness while in the second thermal bending phase, i.e. $t_i \sim 30$ and $h_0(t_i) \sim 5$.

Overall, the time for the current to reach the transition t_i is the time for its radius to reach $R(t) = 4$. Setting (3.2) equal to 4, we obtain $t_i = 6.5\eta_e^{2/7}h_f^{-1/7}$ where η_e is the effective viscosity of the current. This transition time depends on the effective viscosity and is bounded by two values $t_i^h < t_i < t_i^c$. If the current transitions to the gravity regime while in the first thermal bending phase, then $\eta_e \approx 1$ and $t_i^h = 6.5h_f^{-1/7}$; if the current transitions to the gravity regime while in the third thermal bending phase, then $\eta_e \approx \nu^{-1}$ and $t_i^c = 6.5\nu^{-2/7}h_f^{-1/7}$ (table 1).

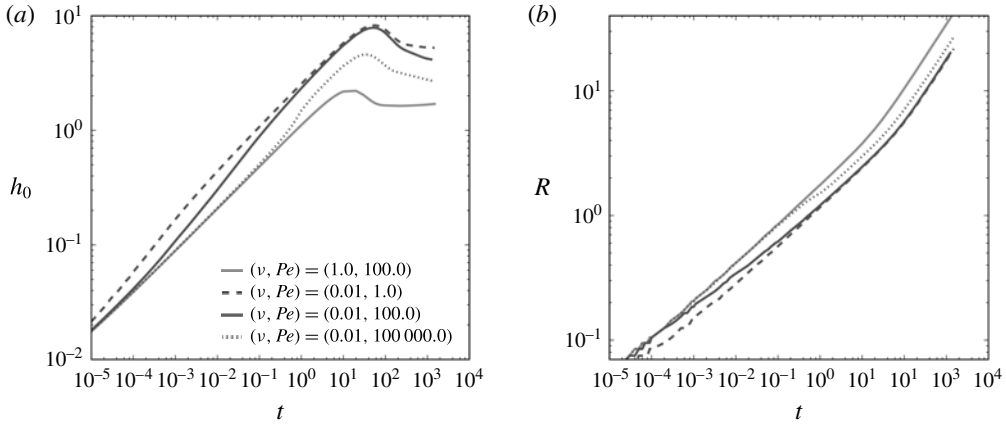


FIGURE 14. (a) Dimensionless thickness at the centre h_0 versus dimensionless time for different sets (ν, Pe) indicated on the plot and an initially hot prewetting film. The grey line represents the isoviscous case $\nu = 1$. (b) Same plot but for the dimensionless radius R . These numerical solutions were obtained by solving (2.39) and (2.40) with the pressure P containing both the gravity and the bending contributions.

Name	From	To	Expression
t_t	Bending	Gravity	$6.5\eta_e^{2/7}h_f^{-1/7}$
t_t^h	Bending	Gravity	$6.5h_f^{-1/7}$
t_t^c	Bending	Gravity	$6.5\nu^{-2/7}h_f^{-1/7}$
Bending regime			
t_{b2}	Phase 1	Phase 2	$7 \times 10^{-3}Pe^{11/9}h_f^{16/9}$
t_{b3}	Phase 2	Phase 3	$14Pe^{11/9}\nu^{-1/9}h_f^{16/9}$
Gravity regime			
t_{g2}	Phase 1	Phase 2	$0.65Pe$
t_{g3}	Phase 2	Phase 3	$10Pe\nu^{-1/2}$

TABLE 1. Summary of the different transition times. t_t is the transition time between bending and gravity which is bound by t_t^h , when the current transitions in the first bending thermal phase, and t_t^c , when the current transitions in the third bending thermal phase. t_{b2} (respectively t_{b3}) represents the time to transition from phase 1 to phase 2 (respectively from phase 2 to phase 3) in the bending regime. t_{g2} (respectively t_{g3}) represents the time to transition from phase 1 to phase 2 (respectively from phase 2 to phase 3) in the gravity regime.

The subsequent evolution in the gravity regime also depends on the combinations of (ν, Pe) considered. Indeed, in contrast to the bending regime where the effective viscosity is that of a small region at the tip, the effective viscosity is the average flow viscosity in the gravity regime. Therefore, the flow effective viscosity can drastically decrease when entering the gravity regime. In particular, a current in the i th bending thermal phase can transition in the j th gravity thermal phase with $j \leq i$.

Overall, five evolution scenarios are possible depending on the combination (ν, Pe) considered and are depicted in the phase diagram proposed in figure 15.

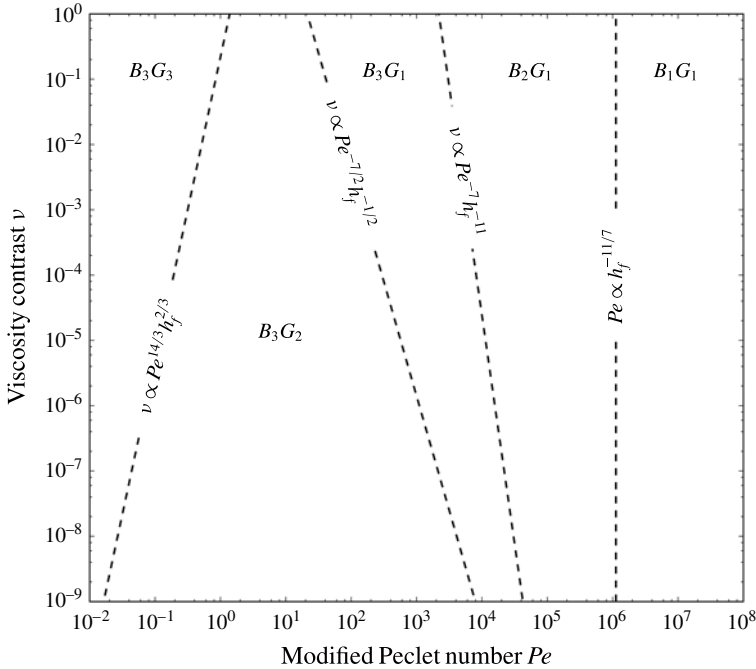


FIGURE 15. Phase diagram for the evolution with bending and gravity for different combinations (ν, Pe) and a given value of $h_f = 0.005$. B_iG_j refers to the case where the current transitions from the i th bending thermal phase to the j th gravity thermal phase where i and $j \in \{1, 2, 3\}$.

6. Summary and conclusions

Isothermal elastic-plated gravity currents show two asymptotic regimes. At early times, the gravity is negligible and the peeling of the front is driven by the bending of the overlying layer. In contrast, at late times, the own flow weight becomes the driving pressure and the current evolves in a gravity current regime. In this study, we have developed a theory for the evolution of an elastic-plated gravity current with a temperature-dependent viscosity and studied the response of the flow to its cooling in each regime separately.

In the bending regime, we show that the flow effective viscosity, or equivalently, its spreading rate, is critically controlled by average local thermal conditions in the peeling region. We identify three main propagation phases. In a first phase, the thermal anomaly grows as the current radius, the peeling region remains relatively hot and the dynamics is close to the hot isoviscous case. In a second phase, a cold front grows and the thermal anomaly detaches from the radius, the temperature in the peeling region rapidly decreases, the current slows down and thickens. Finally, in a third phase, the peeling region is entirely cold and the dynamics returns to an isoviscous propagation but with a cold viscosity. We propose a scaling law for the behaviour of the cold fluid region $(R - R_c)(t)$ in terms of the dimensionless numbers of the system (Pe, ν) and characterize the transition time scales in between the different thermal bending phases.

The evolution of the spreading rate is similar in the gravity regime with three different thermal phases: a first phase of hot isoviscous spreading followed by a second phase of important thickening and a third phase of isoviscous spreading

but with a cold viscosity. However, the spreading rate depends on the average flow viscosity in this regime. In particular, flow propagation is controlled by the behaviour of the thermal anomaly itself. We propose a scaling law for the relative size of the thermal anomaly $(R/R_c)(t)$ in terms of Pe and ν and characterize the different thermal phase transitions in this regime as well.

The overall evolution of an elastic-plated gravity current therefore depends on the relative phase changes within each regime and on the transition between the bending and the gravity regime itself which occurs when $R(t) \sim 4$, whatever the value of h_f , ν and Pe . We finally provide a general phase diagram which summarizes the different evolution scenarios as a function of the dimensionless parameters.

Acknowledgements

We thank three anonymous reviewers for their helpful comments on the manuscript. This work was supported by the UnivEarths Labex program at Sorbonne Paris Cité (ANR-10-LABX-0023 and ANR-11-IDEX-0005-02) as well as by PNP/INSU/CNES.

Appendix A. Influence of the vertical temperature structure

The vertical temperature field used in this study (2.35) assumes a quadratic decrease of the temperature within the thermal boundary layers and hence, a parabolic temperature profile where the two thermal boundary layers have merged. Here, we test whether higher-order contributions (i.e. $n > 2$ in (2.11)) would modify the loss and advection of heat and thereby test the accuracy of our initial assumptions. We show in the following that the details of the vertical temperature structure does not influence our results.

When considering $n > 2$, the vertical temperature profile tends to flatten at the centre and sharpen toward the top and bottom thermal boundary layers. Therefore, while the fluid is more efficiently advected toward the tip of the current over a somewhat larger central zone, heat loss at the boundary is also larger than for the case $n = 2$ and, in the end, both effects compensate for each other. The scaling analysis proposed for the extent of the cold fluid region $(R - R_c)(t)$ (§ 3.2.3) in the bending regime and of the thermal anomaly $(R_c/R)(t)$ in the gravity regime (§ 4.2.2) still apply. The numerical prefactors observed for these quantities are only slightly modified (figure 16). For instance, the extent of the cold fluid region appears slightly larger for larger values of n , because the heat loss is then increased, especially for large Pe (figure 16). This slight difference has clearly a negligible effect on the evolution of the spreading rate which appears independent of the choice of the vertical temperature structure (figure 16).

Overall, the precise shape of the temperature profile, as long as the symmetry is preserved and the boundary layers accounted for, has only little influence on our results. Since the boundary conditions are symmetric here, this simplified model allows to account for a vertical temperature structure including a hotter core within the flow, which was not possible in Balmforth *et al.* (2004) given the asymmetry of the considered cooling process; this constituted the main difference in between the thermal structure obtained from the simplified model and the full heat transport equation.

Appendix B. Numerical scheme

The coupled nonlinear partial differential equations (2.39) and (2.40) are solved on a grid much larger than the flow itself and that is shifted at the centre to avoid problems arising from the axisymmetrical geometry. To solve equations (2.39)

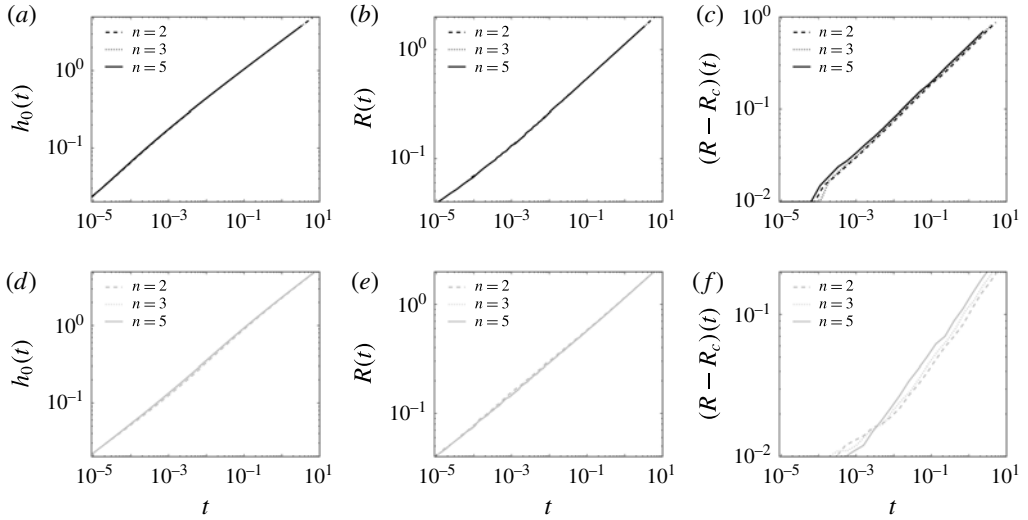


FIGURE 16. (a–c) Dimensionless thickness at the centre h_0 (a,d), radius (b,e) and extent of the cold fluid region $R(t) - R_c(t)$ (c,f) versus dimensionless time t for different exponents n indicated on the plot, $\nu = 0.01$ and $Pe = 1.0$. (d–f) Same plots but for $Pe = 1000$.

and (2.40), we use a finite-difference scheme for spatial discretization coupled with an implicit backward Euler scheme in time. In addition, since each equation is nonlinear, we use a Newton–Raphson method to iterate towards the solution at each time step for both equations. Unless specified differently, we begin the computation with $h = h_f$, $\Theta_b = 0$ and $\delta = h/2$ over the whole domain. In addition, we impose

$$\left. \frac{\partial h}{\partial r} \right|_{r=0} = \left. \frac{\partial P}{\partial r} \right|_{r=0} = 0 \tag{B 1}$$

and $h = h_f$ at the end of the grid.

The expressions of $I_0(\delta)$, $I_1(h)$, $I_1(\delta)$ and $I_2(\delta)$ are the following

$$I_0(\delta) = \frac{\delta}{12}(\nu(6\delta - 6h) + (1 - \nu)\Theta_b(5\delta - 4h)) \tag{B 2}$$

$$I_1(h) = \frac{1}{60}(5h^3\nu + (1 - \nu)\Theta_b(-4\delta^3 + 10\delta^2h - 10\delta h^2 + 5h^3)) \tag{B 3}$$

$$I_1(\delta) = \frac{\delta^2}{120}(\nu(40\delta - 30h) + (1 - \nu)\Theta_b(36\delta - 25h)) \tag{B 4}$$

$$I_2(\delta) = \frac{\delta^2\Theta_b}{2520}(\nu(462\delta - 315h) + (1 - \nu)\Theta_b(428\delta - 280h)) \tag{B 5}$$

and therefore, equations (2.42) and (2.43) reduce to

$$\bar{u} = \frac{\delta}{10}(\nu(20\delta - 30h) + (1 - \nu)\Theta_b(14\delta - 15h)) \tag{B 6}$$

$$\Sigma = \frac{\delta^2\Theta_b}{210}(\nu(98\delta - 105h) + (1 - \nu)\Theta_b(76\delta - 70h)). \tag{B 7}$$

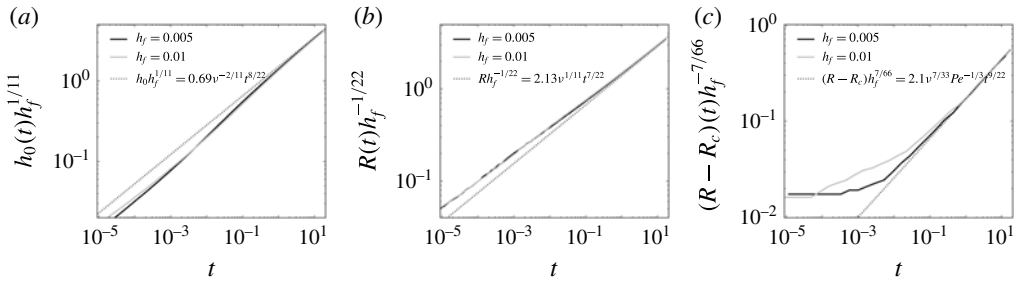


FIGURE 17. (a) Dimensionless thickness at the centre $h_0h_f^{1/11}$ versus dimensionless time t for different h_f indicated on the plot, $\nu = 0.01$ and $Pe = 100$. Dashed line represents the scaling law $h_0h_f^{1/11} = 0.65\nu^{-2/11}t^{8/22}$. (b) Dimensionless radius R versus dimensionless time t for the same h_f . Dashed line represents the scaling law $Rh_f^{-1/22} = 2.13\nu^{1/11}t^{7/22}$. (c) Extent of the cold fluid region $(R(t) - R_c(t))h_f^{-7/66}$ versus dimensionless time for the same h_f . Dashed line: scaling law $(R(t) - R_c(t))h_f^{-7/66} = 2.1Pe^{-1/3}\nu^{7/33}t^{9/22}$.

Appendix C. Influence of the initial temperature and thickness of the film

The thickness of the film is one of the parameters, with Pe and ν , which controls the spreading rate in the bending regime. For instance, the scaling law (3.1) predicts that the thickness h_0 decreases with $h_f^{-1/11}$ and that the radius increases with $h_f^{1/22}$ (3.2). We also show that the extent of the cold fluid region $(R - R_c)(t)$ increases with $h_f^{7/66}$ (3.10). We show there that these scaling laws are in agreement with our numerical simulations for different values of h_f : the different curves collapse when rescaling by h_f elevated to the appropriate power (figure 17).

In the main text, we assume that the prewetting film is initially cold, i.e. $\Theta_b = 0$ and $\delta = h/2$ everywhere. However, the thermal state of the prewetting film at $t = 0$ certainly influences the spreading evolution. In particular, during the first thermal bending phase, the effective viscosity is constant and slightly larger than 1 as the peeling region accounts for the presence of cold fluid in the film at the tip.

When considering an initially hot prewetting film, i.e. $\Theta_b = 1$ and $\delta = 10^{-4}$ at $t = 0$, the effective viscosity of the current is initially closer to 1, i.e. $\eta_e(t) \sim 1$ (figure 18, $Pe = 1000$). The film then cools on a time that scales with $Pe h_f^2$, which is smaller than the time t_{b2} for the current to enter the second thermal bending phase. The effective viscosity increases slightly as the peeling region includes the presence of cold fluid in the film at the tip and the evolution then collapses with the one of an initially cold prewetting film. For instance, for $Pe = 1000.0$, the film has cooled at $t \sim 0.01$ and the current transitions to the second thermal bending phase at $t \sim 0.03$.

In summary, the first thermal bending phase splits in two different phases when considering an initially hot film: a first phase where the current spreads with $\eta_e \sim 1$ which lasts over a time scaling with $Pe h_f^2$ and a second phase where the current spreads with η_e slightly larger than 1 once the film is entirely cold.

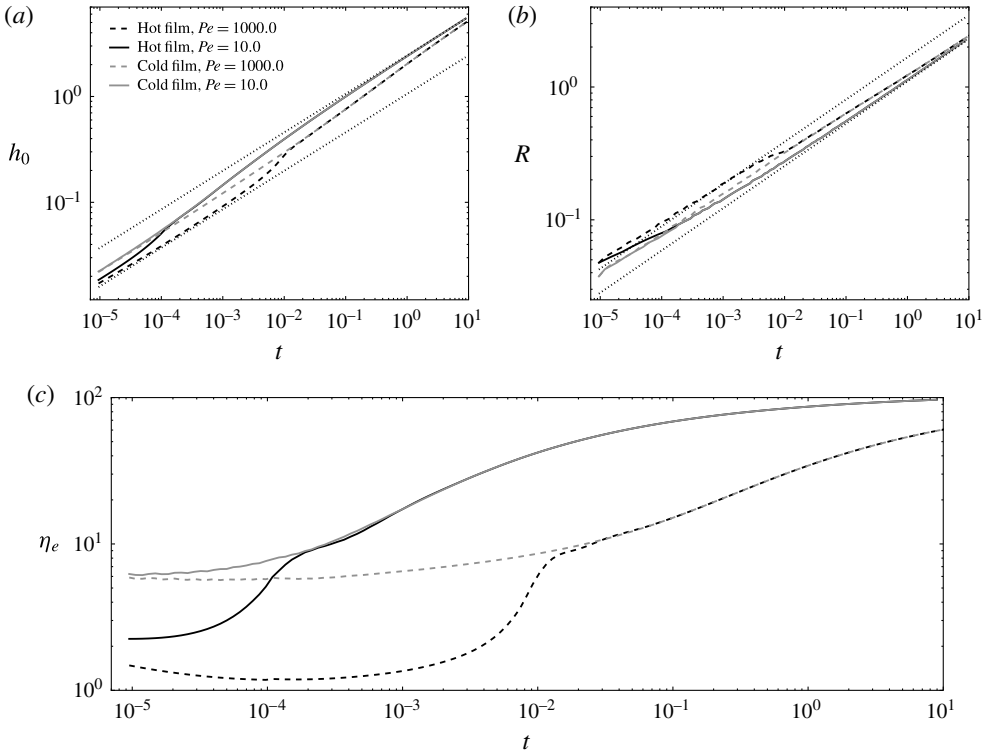


FIGURE 18. (a) Dimensionless thickness at the centre h_0 versus dimensionless time t for an initially cold and hot prewetting film, $Pe = 10.0$ and $Pe = 1000.0$ and $\nu = 0.01$. Dotted lines: scaling laws $h_0 = 0.65h_f^{-1/11}\nu^{-2/11}t^{8/22}$ for $\nu = 1.0$ and 0.001 . (b) Dimensionless radius R versus dimensionless time t for initially cold and hot prewetting films and the same sets of parameters (ν , Pe). Dotted lines: scaling laws $R = 2.13h_f^{1/22}\nu^{1/11}t^{7/22}$ for $\nu = 1.0$ and 0.001 . (c) Dimensionless effective viscosity versus dimensionless time t for initially cold and hot prewetting films and the same set of parameters (ν , Pe). The initially cold film is set with $\Theta_b = 0$ and $\delta = h/2$. The initially hot film is set with $\Theta_b = 1$ and $\delta = 10^{-4}$.

REFERENCES

- BALMFORTH, N. J., CRASTER, R. V. & SASSI, R. 2004 Dynamics of cooling viscoplastic domes. *J. Fluid Mech.* **499**, 149–182.
- BERCOVICI, D. 1994 A theoretical model of cooling viscous gravity currents with temperature-dependent viscosity. *Geophys. Res. Lett.* **21** (12), 1177–1180.
- BERCOVICI, D. & LIN, J. 1996 A gravity current model of cooling mantle plume heads with temperature-dependent buoyancy and viscosity. *J. Geophys. Res.* **101** (B2), 3291–3309.
- BERTOZZI, A. L. 1998 The mathematics of moving contact lines in thin liquid films. *Not. Am. Math. Soc.* **45** (6), 689–697.
- BUNGER, A. P. & CRUDEN, A. R. 2011 Modeling the growth of laccoliths and large mafic sills: Role of magma body forces. *J. Geophys. Res.* **116** (B2), B02203.
- DIXON, J. M. & SIMPSON, D. G. 1987 Centrifuge modelling of laccolith intrusion. *J. Struct. Geol.* **9** (1), 87–103.
- FLITTON, J. C. & KING, J. R. 2004 Moving-boundary and fixed-domain problems for a sixth-order thin-film equation. *Eur. J. Appl. Maths* **15** (06), 713–754.

- GAREL, F., KAMINSKI, E., TAIT, S. & LIMARE, A. 2014 An analogue study of the influence of solidification on the advance and surface thermal signature of lava flows. *Earth Planet. Sci. Lett.* **396**, 46–55.
- GIORDANO, D., RUSSELL, J. K. & DINGWELL, D. B. 2008 Viscosity of magmatic liquids: a model. *Earth Planet. Sci. Lett.* **271** (1), 123–134.
- GOODMAN, T. R. 1958 The heat-balance integral and its application to problems involving a change of phase. *Trans. ASME* **85**, 335–342.
- HEWITT, I. J., BALMFORTH, N. J. & DE BRUYN, J. R. 2015 Elastic-plated gravity currents. *Eur. J. Appl. Maths* **26** (1), 1–31.
- HORT, M. 1997 Cooling and crystallization in sheet-like magma bodies revisited. *J. Volcanol. Geotherm. Res.* **76** (3), 297–317.
- HOSOI, A. & MAHADEVAN, L. 2004 Peeling, healing, and bursting in a lubricated elastic sheet. *Phys. Rev. Lett.* **93** (13), 137802.
- HUPPERT, H. E. 1982a Flow and instability of a viscous current down a slope. *Nature* **300** (5891), 427–429.
- HUPPERT, H. E. 1982b The propagation of two-dimensional and axisymmetric viscous gravity currents over a rigid horizontal surface. *J. Fluid Mech.* **121**, 43–58.
- LEJEUNE, A. M. & RICHEL, P. 1995 Rheology of crystal-bearing silicate melts: an experimental study at high viscosities. *J. Geophys. Res.* **100** (B3), 4215–4229.
- LISTER, J. R., PENG, G. G. & NEUFELD, J. A. 2013 Viscous control of peeling an elastic sheet by bending and pulling. *Phys. Rev. Lett.* **111** (15), 154501.
- MARSH, B. D. 1981 On the crystallinity, probability of occurrence, and rheology of lava and magma. *Contrib. Mineral. Petrol.* **78** (1), 85–98.
- MICHAUT, C. 2011 Dynamics of magmatic intrusions in the upper crust: theory and applications to laccoliths on Earth and the Moon. *J. Geophys. Res. Solid Earth* **116**, B05205.
- MICHAUT, C. & JAUPART, C. 2006 Ultra-rapid formation of large volumes of evolved magma. *Earth Planet. Sci. Lett.* **250** (1), 38–52.
- PENG, G. G., PIHLER-PUZOVIĆ, D., JUEL, A., HEIL, M. & LISTER, J. R. 2015 Displacement flows under elastic membranes. Part 2. Analysis of interfacial effects. *J. Fluid Mech.* **784**, 512–547.
- PIHLER-PUZOVIĆ, D., JUEL, A., PENG, G. G. & HEIL, M. 2015 Displacement flows under elastic membranes. Part 1. Experiments and direct numerical simulations. *J. Fluid Mech.* **784**, 487–511.
- SHAW, H. R. 1972 Viscosities of magmatic silicate liquids; an empirical method of prediction. *Am. J. Sci.* **272** (9), 870–893.
- SLIM, A. C., BALMFORTH, N. J., CRASTER, R. V. & MILLER, J. C. 2009 Surface wrinkling of a channelized flow. *Proc. R. Soc. Lond. A* **465** (2101), 123–142.
- THOREY, C. & MICHAUT, C. 2014 A model for the dynamics of crater-centered intrusion: Application to lunar floor-fractured craters. *J. Geophys. Res.* **119**, 286–312.