

Research Article

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Stability of non-proteinogenic amino acids to UV and gamma irradiation

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Abstract

Almost all living organisms on Earth utilize the same 20 amino acids to build their millions of different proteins, even though there are hundreds of amino acids naturally occurring on Earth. Although it is likely that both the prebiotic and the current environment of Earth shaped the selection of these 20 proteinogenic amino acids, environmental conditions on extraterrestrial planets and moons are known to be quite different than those on Earth. In particular, the surfaces of planets and moons such as Mars, Europa and Enceladus have a much greater flux of UV and gamma radiation impacting their surface than that of Earth. Thus, if life were to have evolved extraterrestrially, a different lexicon of amino acids may have been selected due to different environmental pressures, such as higher radiation exposure. One fundamental property an amino acid must have in order to be of use to the evolution of life is relative stability. Therefore, we studied the stability of three different proteinogenic amino acids (tyrosine, phenylalanine and tryptophan) as compared with 20 non-proteinogenic amino acids that were structurally similar to the aromatic proteinogenic amino acids, following ultraviolet (UV) light (254, 302, or 365 nm) and gamma-ray irradiation. The degree of degradation of the amino acids was quantified using an ultra-high performance liquid chromatography-mass spectrometer (UPLC-MS). The result showed that many non-proteinogenic amino acids had either equal or increased stability to certain radiation wavelengths as compared with their proteinogenic counterparts, with fluorinated phenylalanine and tryptophan derivatives, in particular, exhibiting enhanced stability as compared with proteinogenic phenylalanine and tryptophan amino acids following gamma and select UV irradiation.

Introduction

Living organisms are composed of a variety of complex organic compounds, with nucleotides and amino acids being the fundamental building blocks of the larger macromolecules employed in all life forms known on Earth: nucleic acids and proteins (Danchin and Sekowska, 2015). There are 20 different amino acids used to compose the proteins in organisms using the standard genetic code, with two additional amino acids being incorporated into proteins using alternative translation mechanisms (Danchin and Sekowska, 2015). These 20 (or 22) amino acids are termed proteinogenic amino acids, but there are hundreds of non-proteinogenic amino acids produced in Nature, as well as many different amino acids that can be synthesized in the laboratory and artificially inserted into proteins using solid phase synthesis and synthetic biology methods (Xie and Schultz, 2006; Neumann *et al.*, 2010).

Although extraterrestrial life may exist in many different forms, the conservative hypothesis in exploring other terrestrial planets and large satellites is that extended regions of liquid water, conditions that allow for the formation of complex organic molecules and an energy source or gradient must all be present for life, which is similar to life on Earth, to have evolved (Des Marais *et al.*, 2008; Hays *et al.*, 2017). When exploring such potentially habitable environments, the identification of biosignatures in soil and water samples will be the first step in searching for recognizable signs of previous or current life (Hays *et al.*, 2017). Although there are many possible biosignatures and the principles of biological evolution may not be universal, the identification of building blocks similar to our biosphere's molecular machinery would be an extremely strong indication of extraterrestrial life (Des Marais *et al.*, 2008; Burton *et al.*, 2012; Hays *et al.*, 2017). Thus, the identification of DNA, RNA, proteins, or their precursors: nucleotides and amino acids, on potentially habitable planets or satellites would warrant more intense investigation of that site (Des Marais *et al.*, 2008; Hays *et al.*, 2017). One proteinogenic amino acid (glycine) and several amino acid precursor molecules have already been discovered on comets and asteroids (Burton *et al.*, 2012; Elsila *et al.*, 2016). Additionally, over 90 proteinogenic and non-proteinogenic amino acids have been found to exist in multiple

different meteorites, with the number of non-proteinogenic amino acids outnumbering the number of proteinogenic amino acids on meteors (Alyssa and Ralph, 2014; Elsila *et al.*, 2016).

Three likely locations in our Solar System which may satisfy, or have satisfied, the three criteria for a habitable body as discussed above are Mars, Enceladus and Europa (Chyba, 2000; Marion *et al.*, 2003; Bibring *et al.*, 2004; McKay *et al.*, 2008; Tosca *et al.*, 2008; Barge and White, 2017; Steel *et al.*, 2017). However, these locations have a much higher surface exposure to both ultraviolet (UV) and gamma irradiation than Earth (Des Marais *et al.*, 2008; Carlson *et al.*, 2009; Schmidt *et al.*, 2011; Hays *et al.*, 2017). Due to this higher energy irradiation, many postulate that if life were to have evolved it would have done so in subsurface habitats, as thick layers of liquid or solid water provide significant shielding against both of these radiation types (Boston *et al.*, 1992; Cleaves and Miller, 1998; Levy *et al.*, 2000; Weiss *et al.*, 2000; McKay *et al.*, 2008; Parkinson *et al.*, 2008). However, the extent of this shielding is unknown and it is possible that any amino acids on these bodies would still be exposed to a much greater UV and gamma radiation dose than those found on Earth.

Perhaps the first quality an amino acid (or nucleotide) needs to have in order to exist in extraterrestrial living organisms is simply stability. In the case of Mars, for example, amino acids may need enhanced stability to UV and gamma radiation when under a largely carbon dioxide atmosphere. Alternatively, the subsurface oceans of Europa and Enceladus may exhibit high pressures and unique oceanic mineral compositions (Peterson *et al.*, 1969; Kvenvolden *et al.*, 1970; Levy *et al.*, 2000; Ehrenfreund *et al.*, 2001; Cataldo *et al.*, 2008, 2011a, 2011b, 2013a, 2013b; Iglesias-Groth *et al.*, 2011; Sarker *et al.*, 2012). Therefore, in this work, an array of aqueous aromatic proteinogenic and non-proteinogenic amino acids were exposed to a selection of UV and gamma radiation exposures and the degree of degradation of the amino acids was then quantified using UPLC-MS (ultra-high performance liquid chromatography-mass spectrometry) (Swartz, 2005; Xu *et al.*, 2007; Waterval *et al.*, 2009; Want *et al.*, 2010). Although there have been a few previous studies on the UV and gamma radiation stability of certain amino acids, to our knowledge there are neither any previous studies comparing the UV and gamma irradiation stability of the non-proteinogenic amino acids analysed in this study, nor any previous research directly comparing the stability of proteinogenic and non-proteinogenic amino acids with one another after exposure to identical conditions (Peterson *et al.*, 1969; Ehrenfreund and Charnley, 2000; Ehrenfreund *et al.*, 2001; Kate *et al.*, 2005; Ten Kate *et al.*, 2006; Cataldo *et al.*, 2008, 2011a, 2011b; 2013a, 2013b; ten Kate, 2010; Iglesias-Groth *et al.*, 2011; Sarker *et al.*, 2012; Cherubini and Ursini, 2015).

Aromatic amino acids were the focus of this study, with tyrosine, phenylalanine and tryptophan (Tyr, Phe, Trp) representing the proteinogenic aromatic amino acids. Twenty aromatic amino acids that were largely alternatively functionalized structural analogues of tyrosine, tryptophan, or phenylalanine were used as the non-proteinogenic amino acids. This study focused on aromatic amino acids due to their role in biological photoprotection, enzyme mechanisms and hydrophobic core formation in known proteins (Salih *et al.*, 2000; Makwana and Mahalakshmi, 2015). The results showed that some, but not all, non-proteinogenic amino acids tested are as stable as proteinogenic amino acids when exposed to UV and gamma doses, while select (largely fluorinated) non-proteinogenic amino acids are more stable than their proteinogenic amino acid counterparts.

Materials and methods

Instrumentation and materials

An Acquity H Class UPLC-MS (Waters, Milford, MA) instrument, including PDA and QDa detectors (Waters, Milford, MA), Empower V3 software (Orlando, FL) and an autosampler with Waters screw top 12 32 mm² vials with PTFE/Silicone septa (Milford, MA), was used for liquid chromatography and mass spectrometry analyses of all samples. A BEH C8 1.7 μ m, 2.1 \times 100 mm² column (Water, Milford, MA) was used to separate amino acids from degradation products. Gamma irradiations were performed using a Shepherd 109-68R and a Gammacell 220 cobalt-60 gamma irradiator. The dose rates were measured as 10.6 and 0.104 krad min⁻¹, using Fricke dosimetry, for the gamma radiolysis experiments (Fricke, 1934). A three wavelength UV-Lamp (254/302/365 nm) with 3 \times 8 watt UV tubes from Thermo Scientific (Waltham, MA) was used for irradiation experiments.

All proteinogenic amino acids were from a Sigma-Aldrich L-amino acids analytical standard kit and non-proteinogenic amino acids were purchased from Sigma-Aldrich (St. Louis, MO), Santa Cruz Biotechnology (Dallas, TX), Enamine (Kiev, Ukraine), Bachem (Bubendorf, Switzerland), Iris Biotech (Marktredwitz, Germany), Abblis (Houston, TX), ACROS (Geel, Belgium), or MP Biomedical (Santa Ana, CA). Amino acids were suspended in Ultrapure filtered water that was purified using the Milli-Q (Billerica, MA) Advantage A10 water purification system and the ultrapure BioPak polisher. This ultrapure water was also used in solvent preparation. Optima LC/MS formic acid was purchased from Fisher Chemical (Hampton, NH) and Chromasolv MS grade acetonitrile was purchased from Fisher Scientific (Hampton, NH). Round bottom polypropylene cryogenic vials were purchased from Corning (Corning, NY).

Preparation and exposure of amino acid samples

A stock solution was initially prepared of all amino acid samples and aliquots from that initial stock solution was subsequently used for experiments in order to normalize the initial concentration. Amino acid stock solutions were 100 μ M concentration, all amino acids fully dissolved in deionized, ultrapure water and solutions were stored in glass flasks in 4°C refrigeration. This stock was aliquoted into either cryogenic vials or UPLC glass autosampler vials in 1.8 ml aliquots for exposure to different temperatures, UV light, or gamma rays. However, the 24 h UV-B treatment of the 100 μ M tryptophan analogues resulted in degradation such that the concentration of the remaining amino acids was below the detection limit of the single quadrupole MS. Therefore, the 24 h UV-B tryptophan analogue samples (Fig. 1) were from a 1 mM concentration stock solution instead of a 100 μ M. Untreated samples, the controls, which the treated samples were compared with in order to determine the degree of amino acid degradation caused by the treatment, were pipetted into glass autosampler vials and run on the UPLC-MS. For UV and gamma irradiation experiments and high-temperature experiments, 1.8 ml of this initial stock of the amino acid was pipetted into borosilicate autosampler vials under ambient atmospheric conditions. During the UV and gamma irradiation experiments, the temperature in the laboratory ranged from 20°C to 25°C. However, the temperature within the vials during UV or gamma irradiation was not monitored. These capped vials were then immersed in an 80°C hot water bath for 24 h for the high

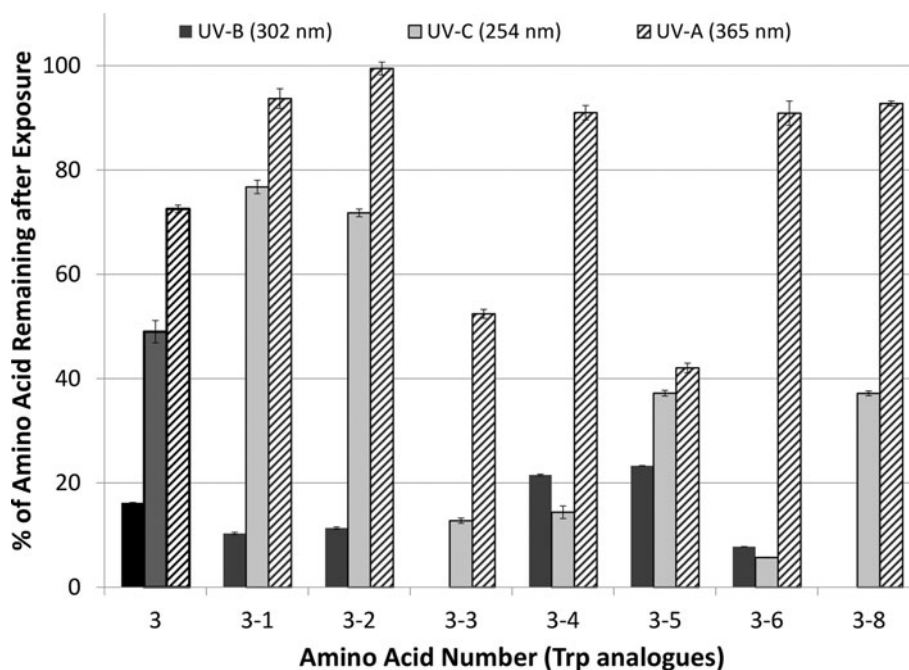


Fig. 1. Stability of select aqueous tryptophan analogous amino acids (from Figs. 2 and 3) to UV-A (365 nm), UV-B (302 nm) and UV-C (254 nm) irradiation when under ambient atmospheric conditions. The darker triplet of bands (3) is the proteinogenic amino acid (Trp). All others (3-1 to 3-8) are non-proteinogenic amino acids whose structures are shown in Figs. 2 and 3. The y-axis shows the % of the amino acid remaining after exposure to the different UV wavelengths.

temperature experiments and then analysed for degradation on the UPLC-MS. For the UV irradiation, the glass vials of the amino acids were placed under the UV lamp (4.5 cm from the light source) at either 254, 302, or 365 nm, which was enclosed within a box under a fume hood for 24 h and then analysed with UPLC-MS. Luminous emittance from the UV lamp at the three different wavelengths was measured using a Fisher Scientific Traceable dual Display light meter. The sampling time for the light meter was 0.4 s and the mercury lamp setting was employed, with the lamp reporting $\pm 4\%$ accuracy and ± 1 lux resolution for the lux range that was determined. Illuminance readings using the light meter were taken at three different locations under the UV-lamp, with UV-A illumination yielding 350 ± 20 lux, UV-B illumination yielding 145 ± 21 lux and UV-C illumination yielding 104 ± 20 lux. For the gamma irradiations, all glass vials were placed within a large glass beaker (400 ml) and this beaker was placed inside the appropriate gamma source for the required time for 399 or 3.99 krad dosage.

The cold temperature samples were pipetted into cryogenic vials (1.8 ml), in triplicate and immersed in liquid nitrogen for 24 h, thawed and pipetted into glass autosampler vials. Cold temperature controls for the amino acids were also produced in which the amino acid was pipetted into cryogenic vials, stored at 4°C for 24 h and then pipetted into glass auto-sampler vials.

UPLC-MS analysis of amino acid samples

The following solvents were used in the UPLC separation on a C8 column: Solvent B: Acetonitrile with 0.1% (v/v) formic acid and Solvent C: Water with 0.1% (v/v) formic acid. The amino acids were separated from reaction products using the following chromatographic conditions: solvents and columns were held at 30°C and the solvent flow rate was 0.4 ml min^{-1} . The solvent gradient began at 5% Solvent B and 95% Solvent C from initial to 8 min and changed to 40% Solvent B and 60% Solvent C from 8.0 to 8.5 min, followed by 2.5 min of 40% Solvent B and 60% Solvent C flow. Solvent composition was then changed to 5%

Solvent B and 95% Solvent C at 11 min for 1 min. The PDA detector was sampled at 20 points per second, recording absorbance at $\lambda = 254 \text{ nm}$ with a 4.8 nm resolution. The QDa detector cone voltage was set at 15 V, using a positive scan and 0.8 kV capillary voltage, sampling at 10 points per second. The mass spectrometer was a single quadrupole, 30–300 Da mass range, 0.2 Da accuracy and 0.7 Da mass resolution. All samples were performed in triplicate.

Tyrosine, phenylalanine and tryptophan were the proteinogenic amino acids tested in ultra-cold or hot temperatures and the initial round of UV-B exposure. Aza-tryptophan, L5-hydroxy-tryptophan, 5-methyl-tryptophan, 5-fluoro-tryptophan, 5-cyano-tryptophan, p-CN-phenylalanine, fluoro-phenylglycine (Fig. 2 = p-F-Phe), p-benzoyl-phenylalanine, p-azido-phenylalanine, H-4-(7-hydroxy-4-coumarinyl), 3-(2-oxo-1,2-dihydro-4-quinolinyl) alanine (oxo-ala in Fig. 2) and 4-boronic acid-phenylalanine, were the 12 non-proteinogenic amino acids tested for temperature degradation and the first round of UV-B exposure (Fig. 2). Based on the results from the first round of UV-B degradation, eight additional non-proteinogenic amino acids (Fig. 3: m-F-Phe, 4-Cl-Phe, p-Br-Phe, m-F-Tyr, 3-Cl-Tyr, 4-F-Trp, 5-Cl-Trp, 5-Br-Trp) were tested during the second round of UV irradiation with three wavelengths of UV light (254, 302, or 365 nm) and two doses of gamma-ray irradiation (399 and 3.99 krad).

Data analysis of UPLC-MS results

The proteinogenic amino acids from the analytical standard kit, tyrosine, tryptophan and phenylalanine exhibited ideal chromatograms, while the non-proteinogenic amino acids varied in this respect, likely due to the reduced purity of these samples. The main peak in the non-proteinogenic amino acids was confirmed to be the amino acid via mass spectrum and all non-proteinogenic amino acids purchased were stated to be above 90% purity. The LC peak corresponding to the amino acid was identified and confirmed by extracting the mass spectrum of that particular peak and comparing it with the known molecular weight of the

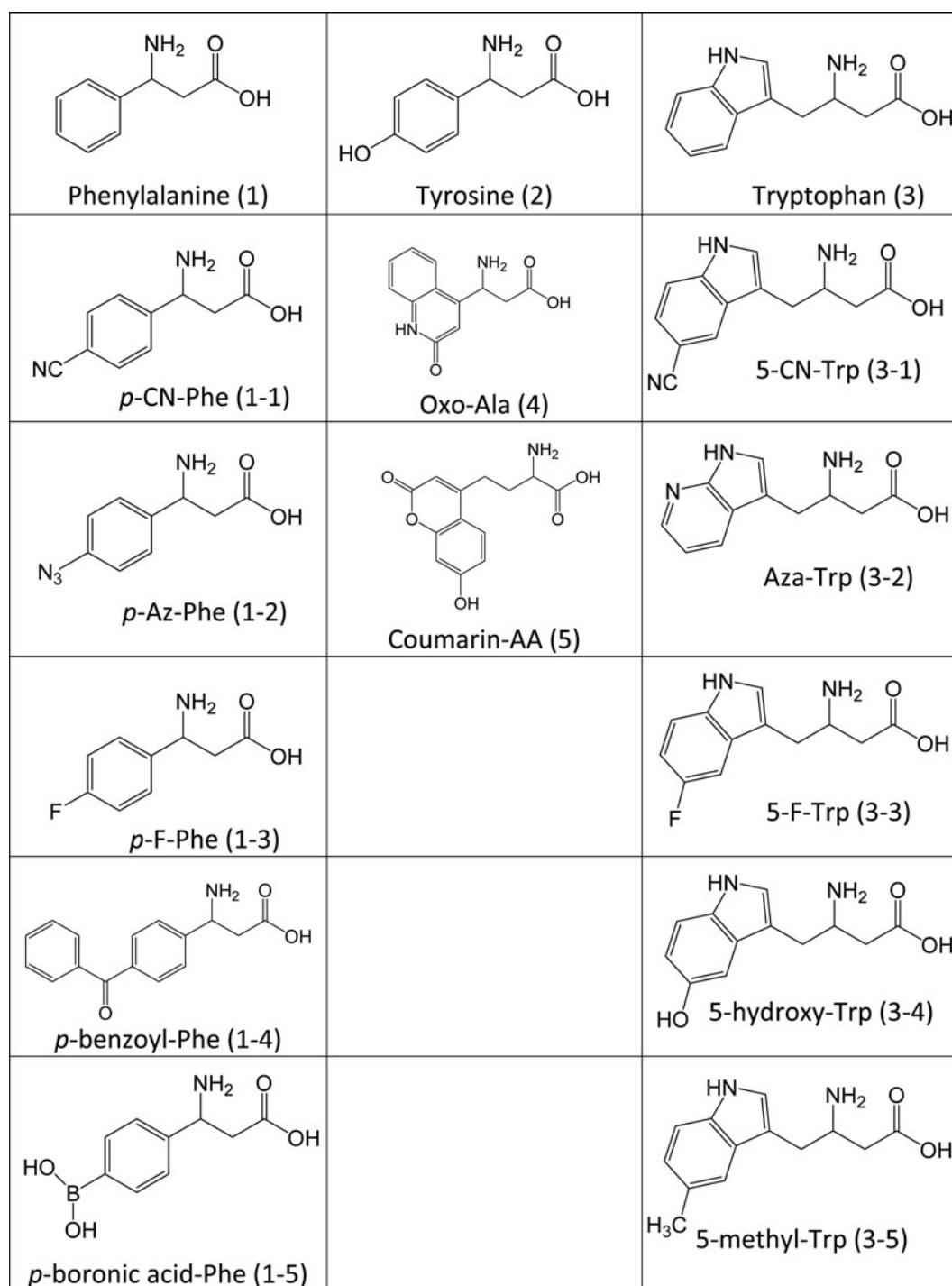


Fig. 2. Structures of amino acids tested against 24 h +80°C exposure and 24 h UV-B exposure. Select amino acids that exhibited UV-B stability were also tested against UV-A, UV-C and gamma irradiation exposure.

amino acid and the mass spectrum of the untreated, control amino acid (see Fig. 4 and Supplementary Information). The peak area of the intact amino acid following condition exposure was then compared with the peak area of the initial amino acid that had no treatment and all measurements were done in triplicate. The differences in integrated LC peak areas of the control and treated samples were used to calculate the percentage of amino acid degradation reported and the identity of LC peaks was confirmed with mass spectrum analysis.

Results

Stability of amino acids at varying temperature

As expected, no amino acids tested exhibited any statistically significant degradation due to 24 h immersion in liquid nitrogen (−196°C). Similarly, after exposure to 80°C temperature for 24 h, most of the amino acids in aqueous 100 μM solution did not undergo any statistically significant degradation (Supplementary Information Fig. 1). The following amino acids did degrade

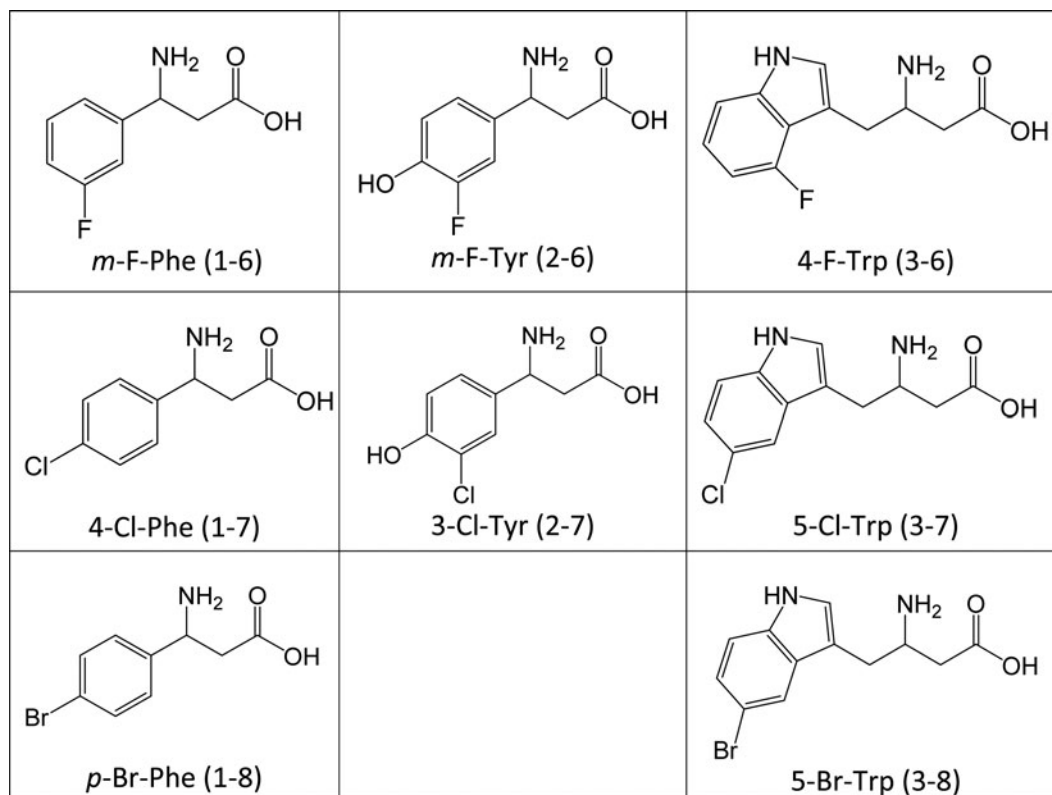


Fig. 3. Structures of additional amino acids tested against UV-B, UV-A, UV-C and gamma irradiation exposure

significantly (>4% degradation) after the 24 h 80°C exposure, listed in order of decreasing degradation: L-5-hydroxy-tryptophan by 12.6% ($\pm 0.9\%$), tryptophan by 8.7% ($\pm 0.2\%$) and oxo-alanine by 7.7% ($\pm 4.9\%$). Several of the tryptophan non-proteinogenic amino acid analogues showed slightly enhanced stability to high-temperature incubation when compared with tryptophan. 5-CN-Trp, Aza-Trp, 5-F-Trp and 5-methyl-Trp showed a slightly increased stability as compared with Trp after 80°C incubation, but no more than a 10% increase in any case.

Stability of amino acids following UV irradiation

Tyrosine, phenylalanine, tryptophan, aza-tryptophan, L5-hydroxy-tryptophan, 5-methyl-tryptophan, 5-fluoro-tryptophan, 5-cyano-tryptophan, p-CN-phenylalanine, fluoro-phenylglycine, p-benzoyl-phenylalanine, p-azido-phenylalanine, H-4-(7-hydroxy-4-coumarinyl), 3,(2-oxo-1,2,-dihydro-4-quinolinyl) alanine (oxo-ala in tables) and 4-boronic acid-phenylalanine were initially tested for degradation after 24 h UV-B exposure when under ambient atmospheric conditions (Supplementary Information Fig. 2). The -CN and -F group on Phe exhibited the most stability of the non-proteinogenic phenylalanine derivatives, while the azide and benzoyl functionalized Phe analogues resulted in nearly a complete transformation from the original amino acid with UV-B irradiation. Aza-Trp and the F- and hydroxy groups on the non-proteinogenic Trp analogues appeared to slightly enhance stability to UV-B irradiation, while the CN group reduced stability and a methyl group in the 5 position of Trp made no significant difference in UV-B stability.

Based on the results of the initial UV-B irradiation experiments, additional non-proteinogenic amino acids were tested with 24 h exposure to UV-B irradiation, as well as UV-A and UV-C

irradiation under ambient atmospheric conditions. Different halogen substitutions on Phe, Tyr and Trp were selected due to the altered stability of fluoro-substituted non-proteinogenic amino acids in the first round of UV-B experiments and previous research showing an increased stability of engineered proteins following incorporation of fluorinated non-proteinogenic amino acids (Buer *et al.*, 2012; Buer and Marsh, 2012; Biava and Budisa, 2014; Marsh, 2014). As can be seen in Fig. 5, phenylalanine (1) did not undergo >10% degradation following exposure to UV-A, UV-B, or UV-C irradiation and tyrosine (2) did not significantly degrade with UV-A or UV-C exposure, but degraded by over 50% with UV-B irradiation. The non-proteinogenic amino acids p-CN-Phe (1-1), p-F-Phe (1-3) and *m*-F-Phe (1-6) all had similar stability to UV-A and UV-C irradiation as the proteinogenic Phe, but p-CN-Phe was less stable to UV-B irradiation. 3-amino-Tyr was similarly stable to all UV wavelengths as its proteinogenic tyrosine (2) counterpart. The coumarin amino acid and the oxo-ala amino acid (5,4), which are not simple analogues of proteinogenic amino acids, were not particularly stable at any UV wavelengths tested. Sample LC chromatograms and the mass spectra of UV-A, UV-B and UV-C results are provided in the Supplementary Information.

As can be seen in Fig. 1, the tryptophan analogues were much less stable at all UV wavelengths, especially UV-B. 5-CN-Trp (3-1) and Aza-Trp (3-2) were more stable than the proteinogenic tryptophan (Trp) following UV-A and UV-C irradiation, but less stable to UV-B irradiation. 5-hydroxy-Trp, 4-F-Trp and 5-Br-Trp were all more stable than Trp at the UV-A wavelength, but less so at the UV-C wavelengths. Both 5-hydroxy-Trp and 5-methyl-Trp were slightly more stable than Trp at UV-B wavelengths, although 5-methyl-Trp had decreased stability at UV-A and UV-C wavelengths.

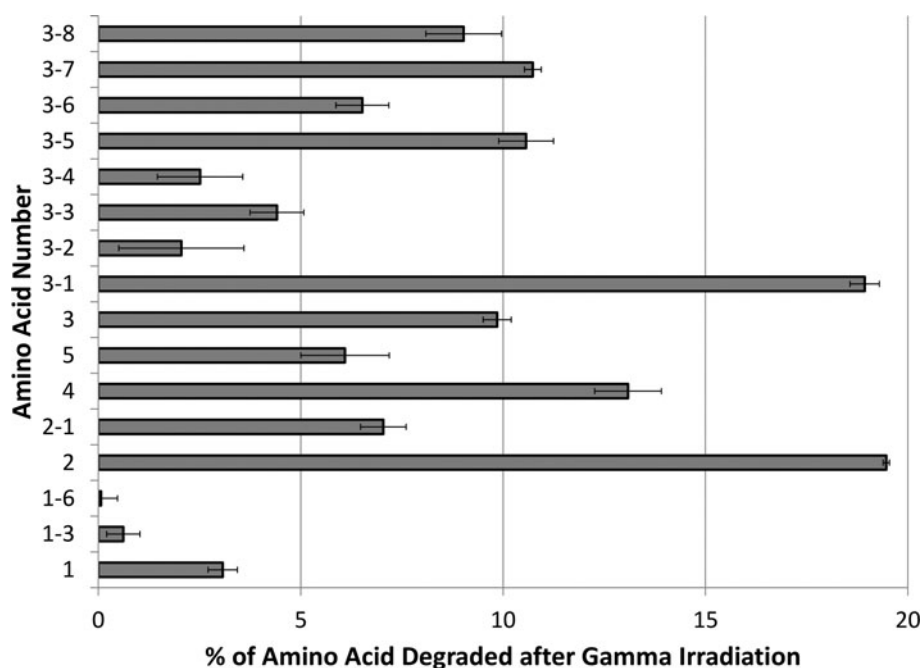


Fig. 4. Stability of select aqueous amino acids (from Figs. 2 and 3) to gamma irradiation when under ambient atmospheric conditions (3.99 krad). The x-axis shows the % of amino acid degraded by the gamma irradiation.

Stability of amino acids following gamma irradiation

The higher gamma dose at ambient atmospheric conditions, 399 krad, resulted in nearly complete degradation of most proteinogenic and non-proteinogenic amino acids tested (1, 2, 4, 5, 3-1, 3-5, 3-7, 3-8). 51 and 96% of the p-F-Phe (1-3) and p-boronic amino acids fully degraded with the high gamma irradiation, as compared with full degradation of the proteinogenic Phe. The 3-amino-Tyr was also more stable under high gamma irradiation than native tyrosine, with only 35% of 3-amino-Tyr (2-1) degrading as compared with 100% of tyrosine degrading following a 399 krad gamma dose. The 5-F-Trp, 5-hydroxy-Trp and 4-F-Trp all had slightly enhanced stability under high gamma doses as compared with the proteinogenic Trp, with 85, 91 and 95% of these non-proteinogenic amino acids being degraded after a 399 krad dose, as compared with 100% of native Trp degrading. The lower gamma dose at ambient atmospheric conditions was two orders of magnitude less than the high gamma dose, at 3.99 krad and at this dose, all tested amino acids maintained significant (>80% intact) stability (Fig. 6). Sample LC chromatograms and the mass spectra of all 3.99 krad gamma exposed amino acids are provided in the Supplementary Information.

Fluorine functionalization of Phe and Trp slightly increased the stability of these non-proteinogenic amino acids to gamma irradiation as compared with the proteinogenic Phe and Trp (Fig. 4). Likewise, amino-, and hydroxyl-functionalization of Tyr and Trp also slightly increased the stability of these non-proteinogenic amino acids to 3.99 krad gamma irradiation, as compared with proteinogenic Tyr and Trp (Fig. 4). However, the increased stability was not a large increase, ranging from 2 to 12%.

Discussion

This study focused on one class of amino acids, the aromatic amino acids, mainly phenylalanine (Phe), tryptophan (Trp) and tyrosine (Tyr) derivatives. The initial round of non-proteinogenic amino acids was chosen in order to represent a wide range of

functional substitutions on the Phe, Tyr, or Trp scaffold, as well as test a couple of truly alternative, but still aromatic, non-proteinogenic amino acids (4 and 5).

As can be seen in Figs. 1 and 5, different amino acids had altered stability depending on the wavelength of UV light. In most cases, the least energetic UV-A radiation (365 nm) predictably caused the least amount of degradation in both the proteinogenic and non-proteinogenic amino acids. The absorbance spectrum from the UPLC-PDA of Tyr, Trp, Phe and all of the non-proteinogenic amino acids, except 4 and 5, show that these amino acids do not absorb at 365 nm radiation to any significant degree (see Supplementary Information for absorbance spectra of all amino acids). The greatest degree of UV degradation was seen with UV-B irradiation, instead of UV-C irradiation, even though UV-C rays are of higher energy than UV-B rays. Inspection of the absorbance spectra of Phe, Tyr and Trp, however, show a 'valley' in the absorbance spectrum of these amino acids in the 254 nm range (Supplementary Information; Beaven and Holiday, 1952; Edelhoich, 1967). The UV-C lamp used in this study specified a wavelength of 254 nm. Although Tyr, Trp and Phe all absorb very strongly at shorter UV-C wavelengths (<220 nm) this low wavelength of UV-C radiation was not utilized in this study. The UV-B wavelength used was 302 nm, which is on the edge of absorbance of tryptophan, but outside the range of phenylalanine's absorption (Supplementary Information, Beaven and Holiday, 1952). The results mimic this, with Phe barely degrading with this UV-B dose, Trp strongly degrading and Tyr being somewhere in the middle.

The non-proteinogenic phenylalanine analogues (Fig. 5) did not show a dramatic change in UV-stability as compared with native Phe. An exception to this is the p-CN-Phe (1-1), which had significantly decreased stability to UV-B irradiation as compared with native Phe. However, the Trp analogous non-proteinogenic amino acid stability varied significantly depending on UV wavelength exposure. Inspection of the absorbance spectrum of these amino acids obtained with the UPLC-PDA do not lend to a simple correlation of absorption spectra with stability (see Supplementary

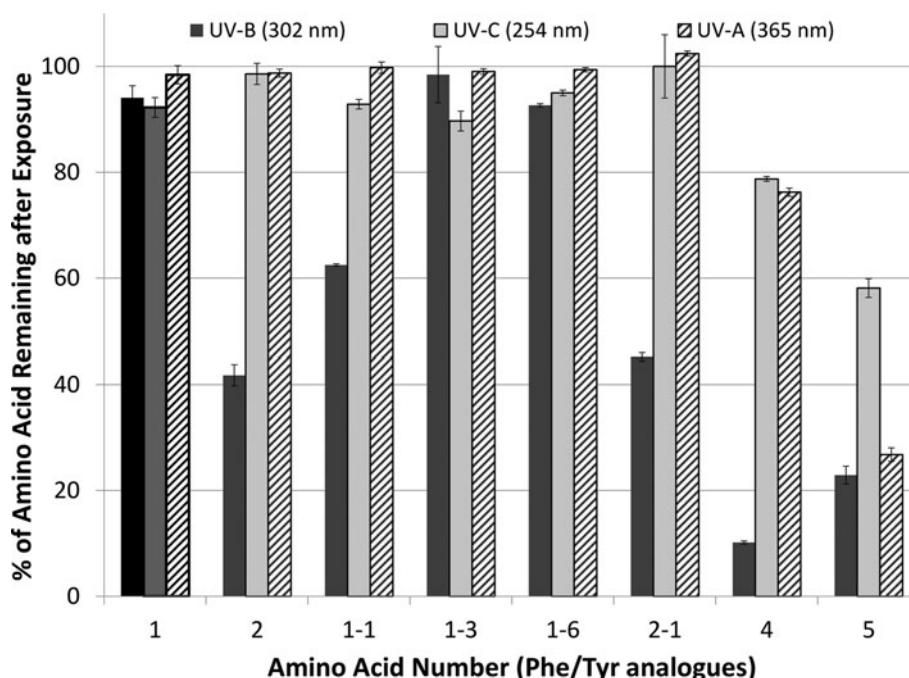


Fig. 5. Stability of select aqueous phenylalanine and tyrosine analogous amino acids (from Figs. 2 and 3) to UV-A (365 nm), UV-B (302 nm) and UV-C (254 nm) irradiation when under ambient atmospheric conditions. The darker triplet of bands (1) is the proteinogenic amino acid (Phe). Tyrosine (2) is also a proteinogenic amino acid. All others (1-1 to 5) are non-proteinogenic amino acids whose structures are shown in Figs. 2 and 3. The y-axis shows the % of the amino acid remaining after exposure to the different UV wavelengths.

Information). For example, most substitutions at carbon 5 or 4 on tryptophan (3-3, 3-4, 3-5, 3-8) resulted in a red-shifted absorbance spectrum, as compared with native tryptophan. This slightly increased the amount of absorbed UV-B at 302 nm of these amino acids, which may correlate to decreased stability to the UV-B irradiation as compared with native tryptophan. While this was the case for 3-3 and 3-8, which had halogen functional groups at position 5, amino acids 3-4 and 3-5, which had a hydroxyl or methyl group at carbon position 5, had slightly increased stability to UV-B irradiation (Fig. 1). The changes in stability of the tryptophan analogues to UV-B irradiation correlated more closely as to whether an electron donating or an electron withdrawing group was added to tryptophan. For example, adding any electron withdrawing group to the tryptophan at carbon 5 or 4 (-CN, F, or Br) decreased the stability of that amino acid to UV-B degradation, whereas adding an electron donating group at carbon 5 (CH₃ and OH) increased the stability of that amino acid to UV-B irradiation (Fig. 4).

Gamma rays have the smallest wavelength and the greatest energy of any wave in the electromagnetic spectrum. Gamma rays are ionizing radiation and as such have the ability to break molecular bonds in many compounds (such as DNA strands), as well as producing highly reactive radicals from the radiolysis of water molecules (Fricke, 1934; Henner *et al.*, 1982). There are many astronomical events that produce gamma rays and galactic cosmic rays and strong solar energetic particles can produce gamma rays and neutrons when impacting the surface of planets and moons (Dartnell *et al.*, 2007; Cooper *et al.*, 2008).

The gamma dosages in this study were chosen due to the Mars rover Curiosity radiation assessment detector data, which recorded approximately 20 mrad radiation exposure daily combined with 2–3 annual proton event spikes, each of which bombarded Mars with approximately a 2000 mrad dose (Hassler *et al.*, 2013). This resulted in approximately 13 300 mrad of radiation per 365 days observed on the surface of Mars. Since microbial life on Earth is estimated to have taken at least 300 million year to evolve, 300 million (Earth) years on Mars would yield a

radiation dose of 3.99×10^6 krad (Cavicchioli, 2002; Schulze-Makuch *et al.*, 2002; Griffin, 2013). Assuming atmospheric or surface shielding to block much of that, 399 krad was equal to 1/10 000 of the initial dose and 3.99 krad was equal to one millionth of the initial dose calculated and these values (399 and 3.99 krad) were the doses of gamma radiation that select amino acid solutions were exposed to using a cobalt-60 source.

All amino acids listed in Fig. 6 were tested with both high (399 krad) and low (3.99 krad) gamma-ray exposure. Gamma-ray exposure to water creates hydroxyl radicals, hydrated electrons and hydrogen atoms via radiolysis of water (Buxton *et al.*, 1988). At the low (100 μ M) aqueous concentrations of amino acids tested, it is assumed that these radicals are the species that are affecting and degrading the amino acids, and not the gamma rays directly (Buxton *et al.*, 1988). Many of the amino acids tested completely degraded following the 399 krad gamma exposure. However, as mentioned in the Results section, p-F-Phe (1-3), p-boronic amino acid (1-5), 3-amino-Tyr (2-1), 5-F-Trp (3-3), 5-hydroxy-Trp (3-4) and 4-F-Trp (3-6) all had enhanced stability under high gamma doses as compared with the proteinogenic Phe, Tyr and Trp.

As expected, the lower gamma dosage did not completely destroy the original amino acids or create as many unresolvable side products that interfered with LC peak interpretation (Bonifačić *et al.*, 1998). No amino acids tested (Fig. 4) degraded by more than 20% following the low gamma dosage since this dose equates to approximately 24 μ M of reactive radical species from the 3.99 krad of gamma irradiation. Similarly to what was seen following the high gamma irradiation dosage, fluorination of both Phe and Trp slightly increased the stability of these non-proteinogenic amino acids to gamma degradation, as compared with the proteinogenic Phe and Trp (Figs. 4 and 6). Likewise, amino-, and hydroxyl- functionalization of Tyr and Trp also slightly increased the stability of these non-proteinogenic amino acids to 3.99 krad gamma irradiation, as compared with proteinogenic Tyr and Trp (Fig. 4). However, the increased stability was not a large increase, ranging from 2 to 12%.

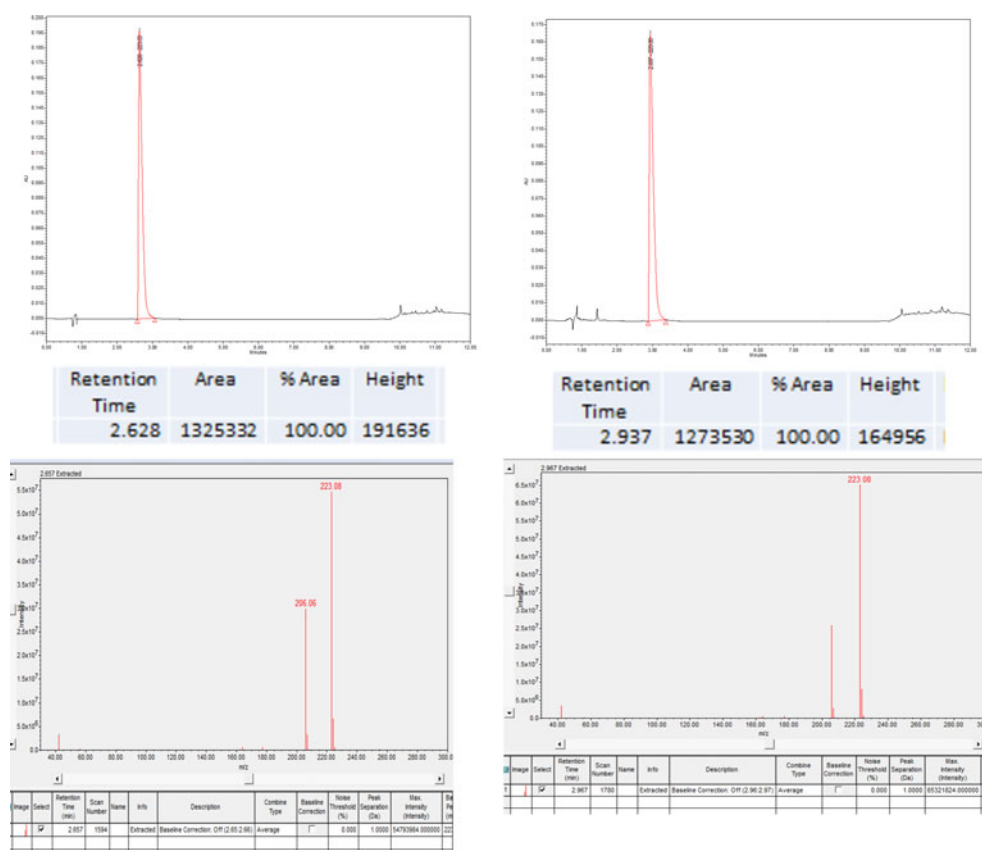


Fig. 6. Sample LC and MS chromatogram of 5-F-Trp (3-3) before (left) and after (right) exposure to 3.99 krad of gamma radiation. The values reported in Fig. 4 are the average of three trials. The LC absorbance at 254 nm was integrated and the mass spectrum extracted at the corresponding retention time in order to positively identify the non-degraded amino acid peak. 5-F-Trp had 95.6% ($\pm 0.66\%$) of the original amino acid remaining (so 4.4% degraded) following 3.99 krad of gamma radiation, as compared with proteinogenic tryptophan, which had 90.1% ($\pm 0.35\%$) of the original amino acid remaining (so 9.9% degraded) following 3.99 krad of gamma radiation.

Interestingly, protein researchers and synthetic biologists have also found that substituting fluorinated non-proteinogenic amino acids for regular (non-fluorinated) amino acids often significantly increases the stability of the resulting engineered proteins due to the increased stability of the resulting hydrophobic core, suggesting fluorine may have a unique role in constructing super-stable proteins (Schröder and Meesters, 2005; Buer and Marsh, 2012; Buer *et al.*, 2012). One of the reasons it is assumed that fluorine is not used in biological building blocks on Earth because the unusually strong C-F bond, as well as the stronger hydrogen bonding capability of the F-H functional group, creates molecules that have an energetic barrier that is too high for typical metabolic reactions that require rapid and dynamic bond breakages with minimal energy expenditure. However, extraterrestrial environments with high radiation exposures may favour biological macromolecules with the stronger bond formation and enhanced macromolecular stability that fluorinated monomers would impart to biological polymers.

Conclusions and future work

Many non-proteinogenic acids tested herein have equal, or greater, stability than their proteinogenic counterparts when exposed to varying wavelengths of UV radiation and gamma radiation. Increased stability of non-proteinogenic amino acids to UV and gamma irradiation was not dramatic (<12% increase), but it is

notable that similar to thermal and denaturing stability studies in synthetic biology, fluorination of both tryptophan and phenylalanine increases their stability to gamma radiation degradation (Buer and Marsh, 2012; Buer *et al.*, 2012). These results suggest that certain non-proteinogenic amino acids may be more stable than the proteinogenic aromatic amino acids under high radiation conditions that may be found extraterrestrially, making them possible alternative building blocks for the evolution of extraterrestrial life. Future work will further characterize these amino acids under more intermediate gamma irradiation exposure, analyse the main transformation products, as well as testing the relative UV and gamma stability of non-aromatic amino acids.

Supplementary material. The supplementary material for this article can be found at <https://doi.org/10.1017/S1473550418000381>

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References

- Alyssa KC and Ralph EP (2014) Nature's starships. I. Observed abundances and relative frequencies of amino acids in meteorites. *Astrophysical Journal* **783**, 140.
- Barge LM and White LM (2017) Experimentally testing hydrothermal vent origin of life on Enceladus and other icy/ocean worlds. *Astrobiology* **17**, 820–833.
- Beaven GH and Holiday ER (1952) Ultraviolet absorption spectra of proteins and amino acids. In Anson ML, Bailey K and Edsall JT (eds), *Advances in Protein Chemistry*, Vol. 7. Elsevier Publishing, pp. 319–386.
- Biava H and Budisa N (2014) Evolution of fluorinated enzymes: an emerging trend for biocatalyst stabilization. *Engineering in Life Sciences* **14**, 340–351.
- Bibring J-P, Langevin Y, Poulet F, Gendrin A, Gondet B, Berthé M, Soufflot A, Drossart P, Combes M, Bellucci G, Moroz V, Mangold N and Schmitt B, Omega team, t. (2004) Perennial water ice identified in the south polar cap of Mars. *Nature* **428**, 627.
- Bonifačić M, Štefanić I, Hug GL, Armstrong DA and Asmus K-D (1998) Glycine decarboxylation: the free radical mechanism. *Journal of the American Chemical Society* **120**, 9930–9940.
- Boston PJ, Ivanov MV and McKay CP (1992) On the possibility of chemosynthetic ecosystems in subsurface habitats on Mars. *Icarus* **95**, 300–308.
- Buer BC and Marsh ENG (2012) Fluorine: a new element in protein design. *Protein Science* **21**, 453–462.
- Buer BC, Meagher JL, Stuckey JA and Marsh ENG (2012) Structural basis for the enhanced stability of highly fluorinated proteins. *Proceedings of the National Academy of Sciences of the United States of America* **109**, 4810–4815.
- Burton AS, Stern JC, Elsila JE, Glavin DP and Dworkin JP (2012) Understanding prebiotic chemistry through the analysis of extraterrestrial amino acids and nucleobases in meteorites. *Chemical Society Reviews* **41**, 5459–5472.
- Buxton GV, Greenstock CL, Helman WP and Ross AB (1988) Critical review of rate constants for reactions of hydrated electrons, hydrogen atoms and hydroxyl radicals ($\cdot\text{OH}/\cdot\text{O}$ —in Aqueous Solution. *Journal of Physical and Chemical Reference Data* **17**, 513–886.
- Carlson R, Calvin W, Dalton J, Hansen G, Hudson R, Johnson R, McCord T and Moore M (2009) Europa's surface composition. In Pappalardo RT, McKinnon WB, Khurana KK and Dotson R (eds), *Europa*. University of Arizona Press, pp. 283–327.
- Cataldo F, Ursini O and Angelini G (2008) Radioracemization and radiation-induced chiral amplification of chiral terpenes measured by optical rotatory dispersion (ORD) spectroscopy. *Radiation Physics and Chemistry* **77**, 961–967.
- Cataldo F, Angelini G, Iglesias-Groth S and Machado A (2011a) Solid state radiolysis of amino acids in an astrochemical perspective. *Radiation Physics and Chemistry* **80**, 57–65.
- Cataldo F, Ragni P, Iglesias-Groth S and Machado A (2011b) A detailed analysis of the properties of radiolyzed proteinaceous amino acids. *Journal of Radioanalytical and Nuclear Chemistry* **287**, 903–911.
- Cataldo F, Angelini G, Hafez Y and Iglesias-Groth S (2013a) Solid state radiolysis of non-proteinaceous amino acids in vacuum: astrochemical implications. *Journal of Radioanalytical and Nuclear Chemistry* **295**, 1235–1243.
- Cataldo F, Iglesias-Groth S, Angelini G and Hafez Y (2013b) Stability toward high energy radiation of Non-proteinogenic amino acids: implications for the origins of life. *Life* **3**, 449–473.
- Cavicchioli R (2002) Extremophiles and the search for extraterrestrial life. *Astrobiology* **2**, 281–292.
- Cherubini C and Ursini O (2015) Amino acids chemical stability submitted to solid state irradiation: the case study of leucine, isoleucine and valine. *SpringerPlus* **4**, 541.
- Chyba CF (2000) Energy for microbial life on Europa. *Nature* **403**, 381.
- Cleaves HJ and Miller SL (1998) Oceanic protection of prebiotic organic compounds from UV radiation. *Proceedings of the National Academy of Sciences of the United States of America* **95**, 7260–7263.
- Cooper J., Sittler Jr E., Hartle R. and Sturmer S. (2008) In European Planetary Science Congress, Vol. 1.
- Danchin A and Sekowska A (2015) The logic of metabolism. *Perspectives in Science* **6**, 15–26.
- Dartnell LR, Desorgher L, Ward J and Coates A (2007) Modelling the surface and subsurface Martian radiation environment: implications for astrobiology. *Geophysical Research Letters* **34**, 1–6.
- Des Marais DJ, Nuth JA, Allamandola LJ, Boss AP, Farmer JD, Hoehler TM, Jakosky BM, Meadows VS, Pohorille A, Runnegar B and Spormann AM (2008) The NASA astrobiology roadmap. *Astrobiology* **8**, 715–730.
- Edelhoc H (1967) Spectroscopic determination of tryptophan and tyrosine in proteins*. *Biochemistry* **6**, 1948–1954.
- Ehrenfreund P and Charnley SB (2000) Organic molecules in the interstellar medium, comets, and meteorites: a voyage from dark clouds to the early earth. *Annual Review of Astronomy and Astrophysics* **38**, 427–483.
- Ehrenfreund P, Bernstein MP, Dworkin JP, Sandford SA and Allamandola LJ (2001) The photostability of amino acids in space. *Astrophysical Journal* **550**, L95–L99.
- Elsila JE, Aponte JC, Blackmond DG, Burton AS, Dworkin JP and Glavin DP (2016) Meteoritic amino acids: diversity in compositions reflects parent body histories. *ACS Central Science* **2**, 370–379.
- Fricke H (1934) The reduction of oxygen to hydrogen peroxide by the irradiation of its aqueous solution with X-rays. *Journal of Chemical Physics* **2**, 556–557.
- Griffin DW (2013) The quest for extraterrestrial life: what about the viruses? *Astrobiology* **13**, 774–783.
- Hassler DM, Zeitlin C, Wimmer-Schweingruber RF, Ehresmann B, Rafkin S, Eigenbrode JL, Brinza DE, Weigle G, Böttcher S and Böhm E (2013) Mars' surface radiation environment measured with the Mars science laboratory's curiosity rover. *Science* **343**, 1244797–1–1244797-5.
- Hays LE, Graham HV, Des Marais DJ, Hausrath EM, Horgan B, McCollom TM, Parenteau MN, Potter-McIntyre SL, Williams AJ and Lynch KL (2017) Biosignature preservation and detection in Mars analog environments. *Astrobiology* **17**, 363–400.
- Henner WD, Grunberg SM and Haseltine WA (1982) Sites and structure of gamma radiation-induced DNA strand breaks. *Journal of Biological Chemistry* **257**, 11750–11754.
- Iglesias-Groth S, Cataldo F, Ursini O and Machado A (2011) Amino acids in comets and meteorites: stability under gamma radiation and preservation of the enantiomeric excess. *Monthly Notices of the Royal Astronomical Society* **410**, 1447–1453.
- Kate IL, Garry JR, Peeters Z, Quinn R, Foing B and Ehrenfreund P (2005) Amino acid photostability on the Martian surface. *Meteoritics & Planetary Science* **40**, 1185–1193.
- Kvenvolden K, Lawless J, Pering K, Peterson E, Flores J, Ponnampereuma C, Kaplan IR and Moore C (1970) Evidence for extraterrestrial amino-acids and hydrocarbons in the Murchison meteorite. *Nature* **228**, 923–926.
- Levy M, Miller SL, Brinton K and Bada JL (2000) Prebiotic synthesis of adenine and amino acids under Europa-like conditions. *Icarus* **145**, 609–613.
- Makwana KM and Mahalakshmi R (2015) Implications of aromatic–aromatic interactions: from protein structures to peptide models. *Protein Science* **24**, 1920–1933.
- Marion GM, Fritsen CH, Eicken H and Payne MC (2003) The search for life on Europa: limiting environmental factors, potential habitats, and earth analogues. *Astrobiology* **3**, 785–811.
- Marsh ENG (2014) Fluorinated proteins: from design and synthesis to structure and stability. *Accounts of Chemical Research* **47**, 2878–2886.
- McKay CP, Porco CC, Altheide T, Davis WL and Kral TA (2008) The possible origin and persistence of life on Enceladus and detection of biomarkers in the plume. *Astrobiology* **8**, 909–919.
- Neumann H, Wang K, Davis L, Garcia-Alai M and Chin JW (2010) Encoding multiple unnatural amino acids via evolution of a quadruplet-decoding ribosome. *Nature* **464**, 441.

- Parkinson CD, Liang M-C, Yung YL and Kirschvink JL (2008) Habitability of Enceladus: planetary conditions for life. *Origins of Life and Evolution of the Biosphere* **38**, 355–369.
- Peterson DB, Holian J and Garrison WM (1969) Radiation chemistry of the alpha-amino acids. Gamma. Radiolysis of solid cysteine. *Journal of Physical Chemistry* **73**, 1568–1572.
- Salih A, Larkum A, Cox G, Kühl M and Hoegh-Guldberg O (2000) Fluorescent pigments in corals are photoprotective. *Nature* **408**, 850.
- Sarker PK, Takahashi J-i, Kawamoto Y, Obayashi Y, Kaneko T and Kobayashi K (2012) Photostability of isovaline and its precursor 5-ethyl-5-methylhydantoin exposed to simulated space radiations. *International Journal of Molecular Sciences* **13**, 1006.
- Schmidt B, Blankenship D, Patterson G and Schenk P (2011) Active formation of chaos terrain over shallow subsurface water on Europa. *Nature* **479**, 502.
- Schröder HF and Meesters RJW (2005) Stability of fluorinated surfactants in advanced oxidation processes—a follow up of degradation products using flow injection–mass spectrometry, liquid chromatography–mass spectrometry and liquid chromatography–multiple stage mass spectrometry. *Journal of Chromatography. A* **1082**, 110–119.
- Schulze-Makuch D, Irwin LN and Guan H (2002) Search parameters for the remote detection of extraterrestrial life. *Planetary and Space Science* **50**, 675–683.
- Steel EL, Davila A and McKay CP (2017) Abiotic and biotic formation of amino acids in the Enceladus ocean. *Astrobiology* **17**, 862–875.
- Swartz ME (2005) UPLC™: an introduction and review. *Journal of Liquid Chromatography & Related Technologies* **28**, 1253–1263.
- ten Kate IL (2010) Organics on Mars? *Astrobiology* **10**, 589–603.
- Ten Kate IL, Garry JR, Peeters Z, Foing B and Ehrenfreund P (2006) The effects of Martian near surface conditions on the photochemistry of amino acids. *Planetary and Space Science* **54**, 296–302.
- Tosca NJ, Knoll AH and McLennan SM (2008) Water activity and the challenge for life on early Mars. *Science* **320**, 1204–1207.
- Want EJ, Wilson ID, Gika H, Theodoridis G, Plumb RS, Shockcor J, Holmes E and Nicholson JK (2010) Global metabolic profiling procedures for urine using UPLC-MS. *Nature Protocols* **5**, 1005.
- Waterval WH, Scheijen JL, Ortmans-Ploemen MM, Habets-van der Poel CD and Bierau J (2009) Quantitative UPLC-MS/MS analysis of underivatized amino acids in body fluids is a reliable tool for the diagnosis and follow-up of patients with inborn errors of metabolism. *Clinica Chimica Acta* **407**, 36–42.
- Weiss BP, Yung YL and Nealson KH (2000) Atmospheric energy for subsurface life on Mars? *Proceedings of the National Academy of Sciences* **97**, 1395–1399.
- Xie J and Schultz PG (2006) A chemical toolkit for proteins—an expanded genetic code. *Nature Reviews. Molecular Cell Biology* **7**, 775–782.
- Xu RN, Fan L, Rieser MJ and El-Shourbagy TA (2007) Recent advances in high-throughput quantitative bioanalysis by LC-MS/MS. *Journal of Pharmaceutical and Biomedical Analysis* **44**, 342–355.