



## Article

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# Distributed energy balance, mass balance and climate sensitivity of upper Chandra Basin glaciers, western Himalaya

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**Abstract**

Glacier and snow melt are the primary sources of water for streams, and rivers in upper Indus region of the western Himalaya. However, the magnitude of runoff from this glacierized basin is expected to vary with the available energy in the catchment. Here, we used a physically based energy balance model to estimate the surface energy and surface mass balance (SMB) of the upper Chandra Basin glaciers for 7 hydrological years from 2015 to 2022. A strong seasonality is observed, with net radiation being the dominant energy flux in the summer, while latent and sensible heat flux dominated in the winter. The estimated mean annual SMB of the upper Chandra Basin glaciers is  $-0.51 \pm 0.28$  m w.e.  $a^{-1}$ , with a cumulative SMB of  $-3.54$  m w.e during 7 years from 2015 to 2022. We find that the geographical factors like aspect, slope, size and elevation of the glacier contribute towards the spatial variability of SMB within the study region. The findings reveal that a 42% increase in precipitation is necessary to counteract the additional mass loss resulting from a 1°C increase in air temperature for the upper Chandra Basin glaciers.

**1. Introduction**

The Himalayan region extends over eight countries across Asia (Afghanistan, Bangladesh, Bhutan, China, India, Myanmar, Nepal and Pakistan) and is home to the world's largest ice volume outside the polar regions (Bolch, 2012). These are often called the 'Water Tower of Asia', emphasizing their immense importance as a freshwater resource (Immerzeel, 2020). The meltwater from Himalayan glaciers plays a crucial role in supporting hydropower generation, irrigation and the essential needs of the human population and natural ecosystems (Pritchard, 2019; Immerzeel, 2020; Li, 2022). However, over the past few decades, these glaciers have experienced substantial mass loss (Kaab and others, 2012). Due to the high elevation and challenging terrain of the Himalayas, scientific studies in this region are limited, resulting in gaps in our understanding of various factors including atmospheric conditions, energy fluxes, glacier–atmosphere interaction and glacier dynamics (Mayewski and others, 2020). In particular, calibration and validation of model studies are scarce and filling these knowledge gaps would reduce uncertainty in climate change projections.

The Himalayan region has complex climatic conditions due to the influences of the Indian Summer Monsoon (ISM) system and Western Disturbances (WD), which cause variability in meteorological conditions over the area (Bolch, 2012; Hock and others, 2019; Azam, 2021). In addition, the heterogeneous changes in glacier mass balance are driven primarily by mechanisms associated with altitude-dependent warming, surrounding topography, land–atmospheric interactions and seasonal to interannual variation in atmospheric circulations on a large or regional scale (Kulkarni, 2007; Pepin, 2022; Nair and others, 2023). The higher temperature increases the fraction of solid to liquid precipitation and reduces accumulation (Wang and others, 2013). Further, it intensifies the melt and lengthens the ablation season duration (Arndt and Schneider, 2023). Rising snowlines expose more glacier ice, reducing surface albedo and leading to further melt through snow–albedo feedback (Bolch, 2012). Warming can also alter meltwater refreezing, retention, surface roughness and glacier flow dynamics (Sakai and Fujita, 2017). All these processes interact to change glacier surface energy balance (SEB) and surface mass balance (SMB). Therefore, it is essential to understand the response of glaciers to atmospheric forcing using hydrometeorological observations and numerical models.

To adequately capture the full SEB and SMB, a distributed physically based model incorporating spatial variability in meteorological conditions and surface properties is required (Hock and Holmgren, 2005). Ablation processes are governed by the SEB, consisting of radiative fluxes (net shortwave and longwave radiation), turbulent fluxes (latent and sensible

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heat), latent heat flux from rain and heat conduction into the snow/ice. The heterogeneity of glaciers and their surrounding topography creates significant spatial variability in meteorological variables, leading to complex patterns in energy and mass fluxes (Yu, 2013; Brun, 2019; Patel, 2021). Physically based model studies have demonstrated the importance of implementing distributed approaches to capture the substantial variations in SMB across glaciers (Reijmer and others, 2012). The Western Himalaya, particularly the Chandra Basin in the upper Indus Basin, is an important region for studying glaciers, as it contains several glaciers that are experiencing significant mass loss (Gardelle and others, 2013; Vijay and Braun, 2016; Sharma and others, 2020; Patel, 2021; Pratap, 2023). However, only a few studies from this region have high-resolution data focusing on the variability of meteorological parameters and energy fluxes, limiting our understanding of glacier–atmosphere interactions.

In this study, we have investigated and quantified the distributed SEB and SMB of the upper Chandra Basin glaciers for 7 hydrological years from 2015 to 2022 to understand how glacier SMB is affected by climatic (meteorological variability) and non-climatic (topographic characteristics) factors. In this study, we apply the COSIPY model (COupled Snowpack and Ice SEB and mass-balance model in PYthon) (Sauter and others, 2020) to simulate the SEB and SMB of selected glaciers. There are four automatic weather stations (AWS) located within the study area which are in close proximity to the glaciers, one AWS is situated on the glacier, while the other three are located within  $\sim 3$  km distance. The data obtained from these AWS are used to extrapolate the hourly meteorological variables across the glacier surface, combined with a resampled digital elevation model (DEM), to derive the spatial distribution of SEB and SMB. Furthermore, the model is calibrated using in situ point SMB measurements and then validated over Sutri Dhaka Glacier and Samudra Tapu Glacier. Sources of uncertainty related to meteorological forcing, model parameters and surface characteristics are quantified through Monte Carlo analysis. Furthermore, we analysed both climatic and non-climatic factors influencing the SMB of glaciers in the study region. The sensitivity of glacier SMB to climatic variations is crucial, and therefore, we conducted perturbation experiments to assess the impacts of changes in air temperature and precipitation.

## 2. Study area

The Chandra Basin is one of the major sub-basins of the Chenab River within the Indus River system, which is located in the central crystalline axis of the Pir Panjal range in Lahaul-Spiti, Himachal Pradesh, India (Fig. 1). The Chandra River originates from the southern slopes of the Baralacha Pass and flows for  $\sim 125$  km through the basin before joining the Bhaga River at Tandi (Fig. 1). The Chandra Basin covers a geographical area of 2446 km<sup>2</sup>. The elevation ranges from 2800 to 6592 m a.s.l. and the mean slope is  $\sim 26^\circ$ . It has 211 glaciers, covering an area of 631 km<sup>2</sup> ( $\sim 26\%$  of the total basin area) (Fig. 1). The basin is located within the transition zone between monsoon and arid climate and experiences the ISM during summer (July–September) and the Northern Hemisphere Mid Latitude WD during winter (December–April) (Bookhagen and Burbank, 2010). However,  $\sim 60$ – $80\%$  of the annual precipitation occurs during the winter mainly in the form of snowfall, whereas the remaining 20–40% falls during the summer monsoon season (Koul and Ganjoo, 2010; Oulkar, 2022).

The Chandra Basin glaciers have large differences in size, surface characteristics and orientation/aspect (Fig. 1b). A significant

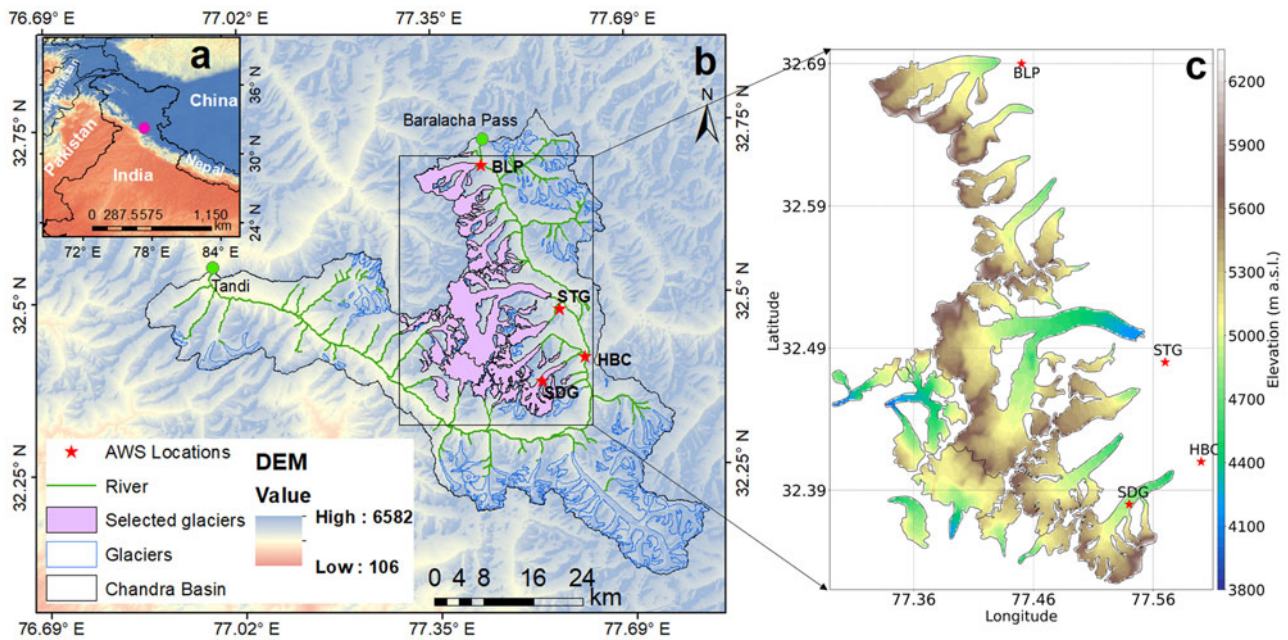
variability in the SEB, particularly the net shortwave radiation within the basin, is caused by both non-climatic and climatic parameters (Patel, 2021; Oulkar, 2022). Therefore, we have categorized the glaciers in the upper Chandra Basin based on the orientation and size of the glaciers (Fig. 1c). The selected glaciers in the upper Chandra Basin are mostly oriented towards the north-east, south and southwest, which has been valuable for conducting orientation analysis. In contrast, the remaining glaciers exhibit substantial variation in orientation, and including these glaciers would have increased the computational time and complexity of the analysis. The selected glaciers cover an area of  $\sim 298$  km<sup>2</sup>,  $\sim 47\%$  of the total glacierized area of the Chandra Basin. In the present study, we have considered only those glaciers which cover an area  $\geq 1$  km<sup>2</sup>, a criterion employed to mitigate uncertainties and ensure a more precise examination of the glacier SEB and SMB. In addition, all these selected glaciers are debris-free glaciers and partially covered with debris over the lower ablation zone. Within the catchment of this study region, we have four AWS installed to monitor the hydrometeorological conditions (Fig. 1 and Fig. S1 in the Supplementary). One AWS is situated on the glacier, while the other three are located within a distance of 3 km. The study area covers 18 glaciers in the upper Chandra Basin which also include benchmark glaciers such as Samudra Tapu and Sutri Dhaka. We have in situ SMB measurements for Sutri Dhaka Glacier and Samudra Tapu Glacier from 2015 to 2022 (Fig. S3 in the Supplementary). These in situ observations are used to validate and estimate the uncertainty of modelled SMB. This will significantly strengthen the effectiveness of the SEB and SMB model simulation and offer a fast and efficient method to evaluate the model performance.

## 3. Data collection and methodology

### 3.1. Meteorological data

The locations and elevations of each AWS are provided in Figure S1 and Table S1. The AWS network includes various meteorological sensors connected to Campbell CR1000 and CR1000X data loggers stored within watertight enclosures. The sensor specifications, accuracy and meteorological variables are detailed in Table 1. The sensors collect data at intervals of 10 min, 30 min and daily, which are stored by the data logger and retrieved during field expeditions. The quality control checks applied to the AWS data include a range test to identify values outside acceptable ranges, an internal consistency test, a time series consistency test to detect sudden jumps or spikes in the data and a spatial consistency test, which involves cross-referencing data with nearby stations to ensure data consistency (Estévez and others, 2011). Additionally, outliers were identified using statistical analysis, such as detecting values beyond 3 SD from the mean (Estévez and others, 2011).

To estimate the SEB and SMB of the glaciers in the upper Chandra Basin, we used hourly meteorological data from the Himansh Base Camp (HBC) AWS located at 4052 m elevation. The HBC has the longest data from October 2015 to September 2022 with only minor data gaps owing to power cuts compared to other AWS (Table S1 in the Supplementary). To fill the data gap, we have used hourly data which has a spatial resolution of  $0.1^\circ \times 0.1^\circ$ , obtained from the European Centre for Medium-Range Weather Forecasts reanalysis (ERA5-land: <https://cds.climate.copernicus.eu/cdsapp#!/dataset/reanalysis-era5-land?tab=overview>) (Hersbach, 2020) at the nearest surface grid points of HBC AWS using bias correction. The bias correction method involved



**Figure 1.** Location map (a) of the study region in the western Himalaya, (b) the Chandra Basin with selected glaciers (pink colour) and (c) elevation gradient of the selected glaciers. The red star represents automatic weather stations (AWS); Himansh Base Camp (HBC, 4052 m a.s.l.), Sutri Dhaka Glacier (SDG, 4864 m a.s.l.), Samudra Tapu Glacier (STG, 4513 m a.s.l.) and Baralacha Pass (BLP, 4904 m a.s.l.).

**Table 1.** List of the AWS sensors and parameters used in the study

Parameter	Sensor type	Measurement range	Sensor accuracy	Height from the surface (m)
Air temperature	Campbell HC2S3	-50°C to +60°C	±0.1°C	2
Relative humidity	Campbell HC2S3	0–100% RH	±0.8% RH	2
Wind speed & wind direction	RM Young Sensor 05103	0–100 ms <sup>-1</sup>	±0.3 ms <sup>-1</sup> & ±3° direction	4
Solar radiation (incoming & outgoing)	Kipp & Zonen CNR4	0–2000 Wm <sup>-2</sup>	± 10%-day total	4
Longwave radiation (incoming & outgoing)	Kipp & Zonen CNR4		± 10%-day total	4
Precipitation	OTT Pluvio <sup>2</sup>	12–1800 mm/h	± 0.05 mm	0
Air pressure	Vaisala CS106	500–1100 hPa	±1.5 hPa (-40 to +60°C)	2

fine-tuning the ERA5-1 and reanalysis data using a linear regression model to align with the observed data from the HBC AWS (Maraun and Widmann, 2018). We calculated the bias by comparing the reanalysis data with the observed data during overlapping periods and applied a correction factor based on the SD. This statistical adjustment ensured that the corrected data closely matched the characteristics of the observed data, improving the accuracy of the gap-filled data. Cloud cover over the study area is estimated using the method from Van den Broeke and others (2006). This method calculates cloud cover as a fraction between 0 and 1 based on net longwave radiation and air temperature measurements.

To simulate the SEB and SMB of the selected glaciers, we use a 100 m resolution DEM derived from the 30 m ASTER GDEM V2 dataset. The ASTER GDEM V2 data are obtained from the Earth Remote Sensing Data Analysis Centre (ERSDAC) (Tachikawa, 2011). The lapse rate/vertical gradient is highly sensitive to surface characteristics and local microclimate, and thus significant variations are expected in the extrapolated data. To ensure accurate distributed data, we have partitioned the selected upper Chandra Basin glaciers into four zones: CB1, CB2, CB3 and CB4 (Fig. S2 in the Supplementary). This division is based on the proximity of these zones to four installed AWS (HBC, STG-Samudra Tapu Glacier, SDG-Sutri Dhaka Glacier and BLP-Baralacha Pass) (Fig. S2 in the Supplementary). Meteorological data obtained from

these AWS are utilized to estimate lapse rates for the corresponding zones, thereby representing the vertical gradients of air temperature and relative humidity (Table S1 in the Supplementary). Lapse rates are computed on a mean annual basis by regression of the AWS data against station elevation for the common observation period of AWS data. The lapse rates are then applied to extrapolate the point HBC AWS data to the entire DEM based on the elevation at each 100 m grid cell to obtain distributed SEB and SMB of the glaciers (Table 2). The precipitation and air pressure gradient data used in this analysis are obtained from previous study by Oulkar (2022). These distributed meteorological data are used as inputs to the SEB model to simulate hourly SMB for the 7 hydrological years.

### 3.2. Methodology

We have used the COSIPY model, which is an open-source (<https://github.com/cryotools/cosipy>), physically based model that simulates the SEB, SMB and subsurface processes for glaciers (Sauter and others, 2020). The model is based on energy and mass conservation and estimates of SEB and SMB. It combines a SEB model with a multilayer snow and ice model to calculate energy fluxes, subsurface processes and SMB at a given resolution (Huintjes, 2015; Sauter and others, 2020). For the present study, the model is forced with HBC AWS hourly observations of incoming

**Table 2.** Details of model parameters (altitudinal gradient/lapse rates) and extrapolation methods

Variables	Parameters/Model			
	CB1	CB2	CB3	CB4
Air temperature lapse rate ( $^{\circ}\text{C m}^{-1}$ )	-0.0035	-0.0031	-0.0044	-0.0037
Relative humidity lapse rate ( $\% \text{ m}^{-1}$ )	0.049	0.032	0.022	0.031
Precipitation gradient ( $\% \text{ m}^{-1}$ )	0.12 (Oulkar, 2022)			
Air pressure gradient ( $\text{hPa m}^{-1}$ )	-0.034 (Oulkar, 2022)			
Cloud covers	Estimated based on <i>Lnet</i> and <i>Tair</i> (Van den Broeke and others, 2006)			
Incoming shortwave radiation	Solar radiation model (Georg and others, 2016)			
Longwave radiation	Stefan–Boltzmann law (Klok and Oerlemans, 2002)			
DEM: digital elevation model (m)	100 × 100 m resolution			

Details of CB1, CB2, CB3 and CB4 are given in Fig. S2 in the Supplementary. The air temperature and relative humidity lapse rates are computed for the common observation period from October 2020 to September 2022 using data from all AWS.

shortwave radiation, incoming longwave radiation, air temperature, total precipitation, relative humidity, surface pressure, wind speed and cloud cover fraction during 2015–22.

### 3.2.1. SEB model

Hourly HBC AWS data and 100 m DEM are used for the COSIPY model to derive the spatial distribution of SEB and SMB of the upper Chandra Basin. The total surface energy flux ( $\text{Wm}^{-2}$ ) at the glacier surface is calculated within the COSIPY model at each time step (hourly) for each grid based on the principle of energy conservation (Oerlemans, 2001):

$$Q = S_{\text{in}}(1 - \alpha) + L_{\text{in}} + L_{\text{out}} + H_{\text{se}} + H_{\text{la}} + Q_{\text{G}} \quad (1)$$

where  $Q$  is the total energy available for melt,  $S_{\text{in}}$  is the incoming shortwave radiation,  $\alpha$  is the surface albedo,  $L_{\text{in}}$  is the incoming longwave radiation,  $L_{\text{out}}$  is the outgoing longwave radiation,  $H_{\text{se}}$  is the turbulent sensible heat flux,  $H_{\text{la}}$  is the turbulent latent heat flux and  $Q_{\text{G}}$  is the ground heat flux. The heat flux resulting from liquid precipitation has a negligible effect; therefore, it is not considered in the model (Huintjes, 2015; Sauter and others, 2020; Oulkar, 2022). The sign convention for the energy flux terms is that the positive values represent an energy gain from the surface, while negative values represent an energy loss from the surface. Here,  $Q$  is determined through an iteration using a Newton–Raphson optimization scheme to solve the SEB equation (Sauter and others, 2020). The surface temperature is the primary variable associated with the energy fluxes. If the resultant surface temperature exceeds  $0^{\circ}\text{C}$ , the excess energy is available for melt ( $Q > 0$ ) and the surface temperature is reset to  $0^{\circ}\text{C}$ .

The  $S_{\text{in}}$  is modelled using the approach of Georg and others (2016) which accounts for the effects of topographic shading, slope and aspect on the solar radiation received at the glacier surface. The decrease of  $S_{\text{in}}$  through the snowpack is computed using an exponential extinction function based on depth, with separate coefficients for snow and ice (Sauter and others, 2020). The surface  $\alpha$  is calculated as a function of snowfall, snow depth, snow age and ice albedo following Oerlemans and Knap (1998). The albedo decreases exponentially from fresh snow to firn albedo over a period of 6 days (Huintjes, 2015; Sauter and others, 2020). The surface albedo parameters in our model are set based on the values

determined by Molg and Scherer (2012). The Stefan–Boltzmann law is used to estimate the modelled  $L_{\text{out}}$  (Klok and Oerlemans, 2002). The turbulent heat fluxes (sensible and latent) are computed using the bulk aerodynamic method with stability corrections, based on measurements of air temperature, humidity and wind speed at 2 m above the surface (Oerlemans, 2001). Stability corrections for stratification are included based on the bulk Richardson number. The surface roughness length for momentum, heat and moisture transfer evolve from fresh snow to ice values based on time or snow depth (Sauter and others, 2020). The  $Q_{\text{G}}$  is comprised of heat conduction fluxes combined with fluxes from the portion of shortwave radiation that penetrates the snowpack/ice.

The values of site-specific parameters within the model are taken from previous studies (Klok and Oerlemans, 2002; Huintjes, 2015; Oulkar, 2022) for SEB and SMB simulations at the upper Chandra Basin glaciers. The air temperature and relative humidity lapse rates are estimated within the present study, while precipitation and air pressure gradients are adopted from Oulkar (2022) (Table 2). Further, the subsurface density profile is initialized on 1 October with constant densities of  $870 \text{ kg m}^{-3}$  for glacier ice and  $490 \text{ kg m}^{-3}$  for snow, based on values used in previous studies (Oulkar, 2022; Pratap and others, 2019).

### 3.2.2. SMB model

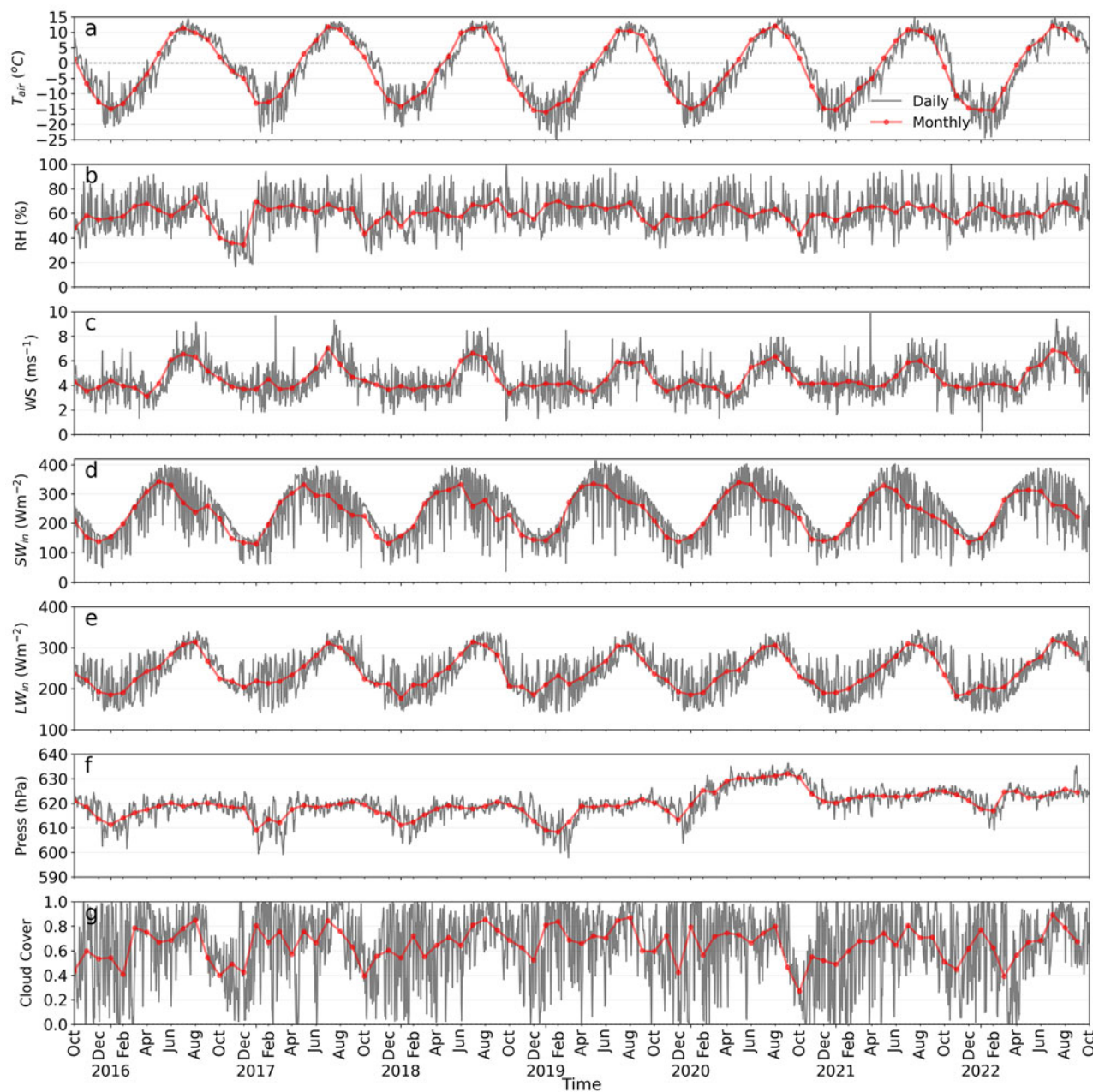
The SMB is estimated as follows:

$$\text{SMB} = \text{Ab} + \text{Ac} + \text{Aci} + \text{Abi} \quad (2)$$

where  $\text{Ab}$  is the surface ablation from melt and sublimation,  $\text{Ac}$  is an accumulation from snowfall and deposition,  $\text{Aci}$  is the internal accumulation from refreezing and  $\text{Abi}$  is the internal ablation from subsurface melt. Solid precipitation (snowfall) contributes directly to accumulation. Liquid precipitation (rainfall) and meltwater percolate through the snowpack following a tipping bucket approach based on the liquid water holding capacity (Sauter and others, 2020). Refreezing occurs when snow temperature is below  $0^{\circ}\text{C}$  and liquid water is present from rainfall or melt. Excess meltwater that reaches the bottom of the snowpack contributes to runoff. The model uses a dynamic mesh with variable layering to represent the vertical profile of snow and ice properties, including temperature, density, liquid water content and ice fraction (Sauter and others, 2020). Layer thickness adapts over time, with thinner layers near the surface and increasing thickness with depth. For more information, the COSIPY model is described in detail by Sauter and others (2020), including the parametrizations, underlying physical principles, model structure and optimization approach.

### 3.2.3. Model uncertainty assessment

The uncertainty in the model output comes from different sources, such as measurements of meteorological variables, uncertainty of energy fluxes, model parameters, threshold values, etc. Therefore, to assess model uncertainty, we perform 500 Monte Carlo simulations varying model parameters and thresholds by 10% and meteorological inputs within measurement uncertainty ranges, following van der Veen (2002) and Machguth and others (2008). Model uncertainty is quantified by running the repeated simulations at all observed point locations on Samudra Tapu Glacier and Sutri Dhaka Glacier (Fig. S3 in the Supplementary). The mean SMB of 500 Monte Carlo simulations is within the uncertainty range of the corresponding in situ values. Based on the results of the 500 simulations compared against the observed SMB, there is a well match between the uncertainty range and observed SMB (Fig. S4 in



**Figure 2.** Observed daily mean values of (a) air temperature ( $T_{air}$ , °C), (b) relative humidity (RH, %), (c) wind speed ( $m\ s^{-1}$ ), (d) incoming shortwave radiation ( $SW_{in}$ ,  $Wm^{-2}$ ), (e) incoming longwave radiation ( $LW_{in}$ ,  $Wm^{-2}$ ), (f) pressure (hPa), and (g) cloud cover over the Chandra Basin glaciers at the site HBC AWS for study period from October 2015 to September 2022.

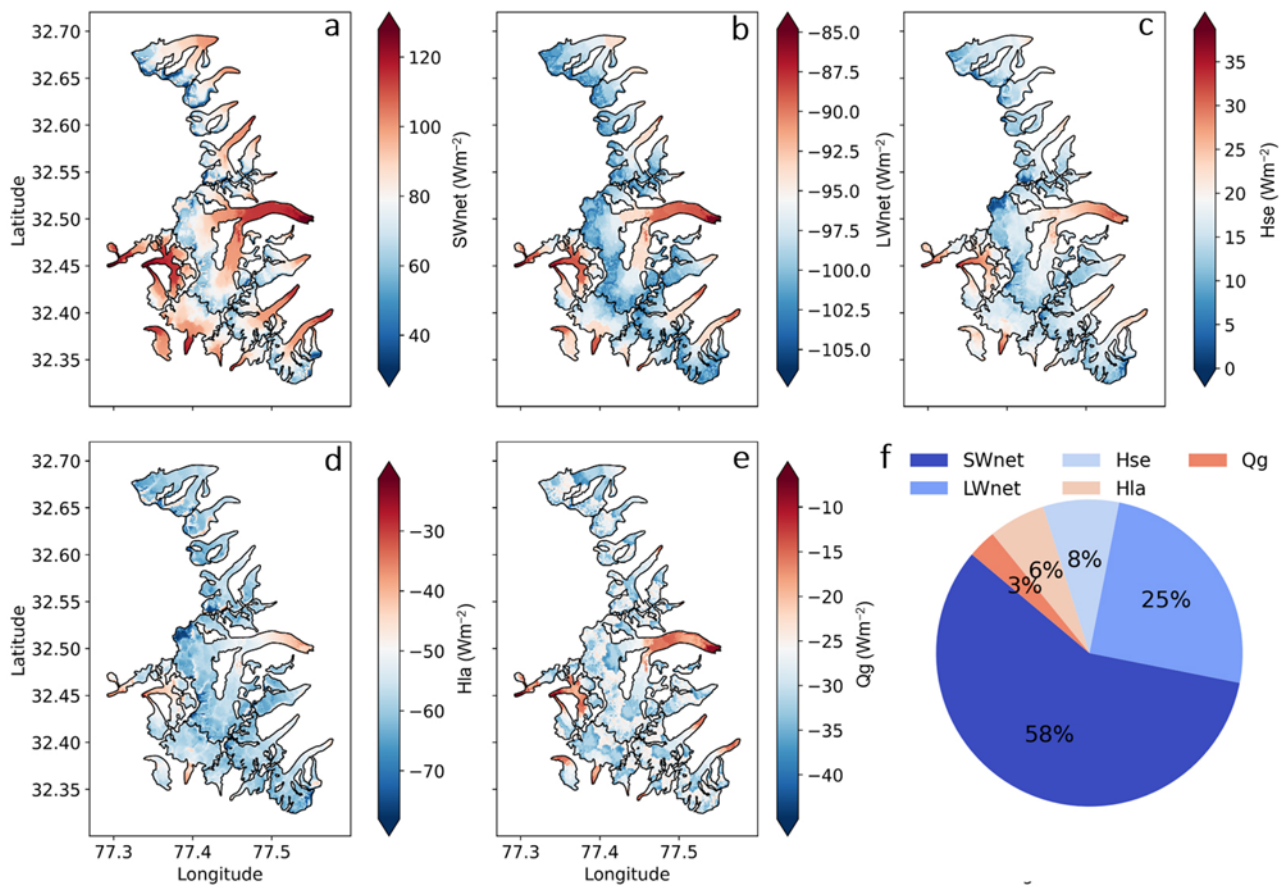
the Supplementary). This uncertainty range is determined by calculating the SD of the SMB results. The SD captures the spread in SMB estimates from the 500 simulations. With a normal distribution of results,  $\sim 68\%$  of simulations are within  $\pm 1\sigma$ . Therefore, the  $\pm 25\%$  uncertainty range represents  $\sim \pm 1\sigma$  of the SMB results from the simulations.

## 4. Results and discussion

### 4.1. Climatic setting

Figure 2 shows the observed daily mean values of air temperature, surface temperature, relative humidity, wind speed, cloud cover,

air pressure, incoming shortwave and incoming longwave radiation over the upper Chandra Basin glaciers for the study period from October 2015 to September 2022. These meteorological variables show substantial temporal variation and distinct patterns over the study period. The daily mean air temperature ranges from  $-25.48^{\circ}C$  to  $15.67^{\circ}C$  with a mean annual value of  $1.67^{\circ}C$  (Fig. 2a). The daily mean relative humidity ranges from 16% to 99%, with a mean annual value of 60% (Fig. 2b). The daily mean wind speed ranges from 0 to  $9.84\ m\ s^{-1}$  with a mean annual value of  $4.57\ m\ s^{-1}$  (Fig. 2c). The daily mean incoming shortwave radiation ranges from 37 to  $400\ Wm^{-2}$  with a mean annual value of  $237\ Wm^{-2}$  (Fig. 2d). The incoming longwave radiation ranges from 139 to  $344\ Wm^{-2}$  with a mean annual value of  $243\ Wm^{-2}$  (Fig. 2e). The daily



**Figure 3.** Distributed mean surface energy fluxes for the glaciers of the upper Chandra Basin for the period from October 2015 to September 2022. (a)  $SW_{net}$  is net shortwave radiation ( $Wm^{-2}$ ), (b)  $LW_{net}$  is net longwave radiation ( $Wm^{-2}$ ), (c)  $H_{se}$  is sensible heat flux ( $Wm^{-2}$ ), (d)  $H_{la}$  is latent heat flux ( $Wm^{-2}$ ), (e)  $Q_g$  is ground heat flux ( $Wm^{-2}$ ) and (f) annual contribution of each energy flux in percentage.

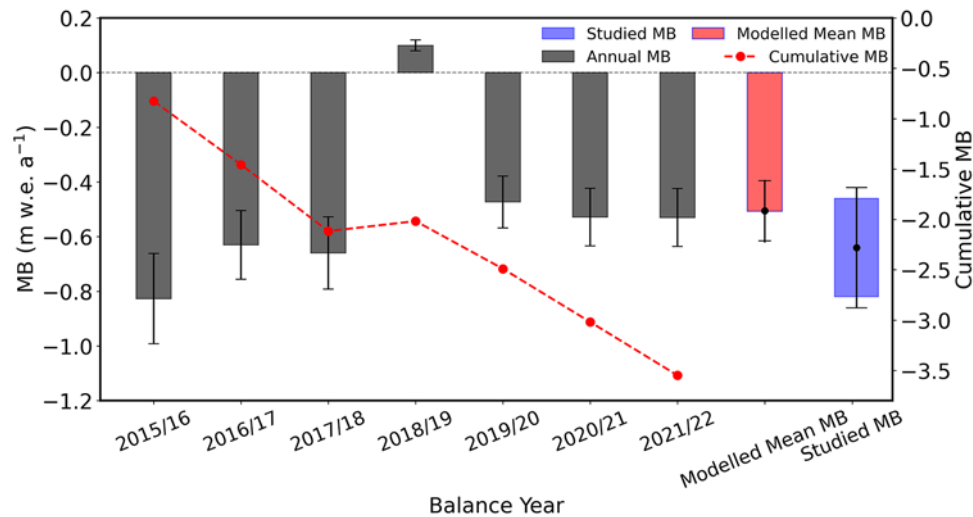
mean pressure ranges from 597 to 636 hPa (Fig. 2f). The cloud cover shows strong monthly variation (Fig. 2g). This significant variability in meteorological variables can be attributed to the semi-arid climate conditions prevailing in the area (Bookhagen and Burbank, 2010). During the summer, warm and humid air is brought into the region, causing temperatures to rise (Oulkar, 2022). Consequently, the higher humidity levels also contribute to the warmer climate during this period. In contrast, during the winter, the monsoon winds subside, and the WB becomes active, leading to a drop in temperatures and an increase in cloud cover (Fig. 2a, g). The combined influence of these factors results in the observed variations in the Chandra Basin climate. The variation of temperatures, radiations, wind and moisture throughout the year can affect the ablation and accumulation rates of the glaciers, leading to changes in their overall SMB.

#### 4.2. Distributed surface energy fluxes

The distributed energy fluxes over 7 hydrological years (2015–22), including mean net shortwave ( $SW_{net}$ ) and longwave ( $LW_{net}$ ) radiation, mean turbulent heat fluxes ( $H_{se}$  and  $H_{la}$ ) and mean ground heat flux ( $Q_g$ ), are shown in Fig. 3. The mean  $SW_{net}$  ranges from 30 to 130  $Wm^{-2}$  (Fig. 3a) and varies with altitude. Specifically, the  $SW_{net}$  values decrease with increasing altitude, primarily due to the higher albedo over the accumulation zone. The  $SW_{net}$  is the

largest energy source, with higher values in the ablation zone compared to the accumulation zone. This reflects the lower albedo and greater absorbed solar radiation at lower elevations.  $SW_{net}$  shows strong seasonal variation, with peak value during the ablation season when incoming solar radiation is at a maximum. The mean  $LW_{net}$  varied from  $-106$  to  $-84$   $Wm^{-2}$  and decreased with the altitude (Fig. 3b). This variation can be attributed to the influence of temperature and relative humidity as a function of altitude (Oulkar, 2022). The  $LW_{net}$  is negative across the region, indicating radiative cooling. More negative values are found at higher altitudes, likely due to colder temperatures. The turbulent heat flux  $H_{se}$  ( $0$ – $38$   $Wm^{-2}$ ) and  $H_{la}$  ( $-78$  to  $-21$   $Wm^{-2}$ ) shows strong spatial variation (Fig. 3c, d). The turbulent heat fluxes of  $H_{se}$  and  $H_{la}$ , shows maximum values at lower elevations due to the presence of large gradients in surface temperature and water vapour pressure. As altitude increases, the turbulent heat fluxes decrease (Fig. 3c, d). The magnitude of  $Q_g$  heat flux is small compared to the other energy components and varied from  $-44$  to  $-7$   $Wm^{-2}$  (Fig. 3e).

The surface energy flux components show a strong seasonality, with net shortwave and longwave radiation being the dominant energy flux in the ablation season (May–September). In contrast, sensible and latent heat flux dominate in the accumulation season (October–April). The annual energy flux analysis showed that the net shortwave radiation contributed 58% to the total surface energy fluxes, followed by net longwave radiation at 25%, sensible heat at



**Figure 4.** Mean annual mass balance of upper Chandra Basin for 7 hydrological years from October 2015 to September 2022. The red bar is modelled mean mass balance for 7 years, the purple box is previously studied mean mass balance within Chandra Basin (Table 3), and the red line shows the cumulative mass balance.

8%, latent heat at 6% and ground heat flux at 3% (Fig. 3f). These observations are consistent with previous studies on Himalayan glaciers in this region (Azam, 2014b; Patel, 2021; Oulkar, 2022). Also, the altitude dependence and seasonal variations match previous observations on high mountain glaciers (Nair and others, 2023).

### 4.3. Mass balance

#### 4.3.1. Mean mass balance

Figure 4 shows the changes in the SMB of the upper Chandra Basin over 7 hydrological years from 2015 to 2022. The annual SMB ranges from  $-0.89$  to  $0.10$  m w.e., indicating an overall trend of negative SMB. The estimated mean SMB of upper Chandra Basin glaciers is  $-0.51 \pm 0.28$  m w.e.  $a^{-1}$  with a cumulative mass balance of  $-3.54$  m w.e. during the last 7 years. Overall, the modelled mean SMB for 7 hydrological years is consistent with the SMB reported by various studies within the Chandra Basin (Table 3). However, a positive SMB is observed for the hydrological year 2018/19. This could be attributed to higher accumulation driven by an extreme snowfall event in 2018 that covered most of the Lahaul–Spiti district (Pratap, 2023). However, this does not change the overall trend of glacier mass balance, driven by long-term climate change. The modelled result agrees with the general trend of decreasing SMB over time, indicating that the upper Chandra Basin is experiencing glacier mass loss.

#### 4.3.2. Distributed mass balance

Figure 5 highlights the variability of glacier SMB across upper Chandra Basin and shows the glaciers experiencing mass loss in the ablation zone. This study shows spatial variability in SMB, with lower values in the high-elevation areas of the glaciers and higher values in the lower elevation. This spatial variability can be attributed to several factors, including differences in glacier geometry, aspect and shading effects (Yu, 2013; Brun, 2019; Olson and Rupper, 2019; Kumar, 2021; Wang, 2022). Along with meteorological parameters, other variables like precipitation distribution, snowdrift effect, albedo differences and streamflow can also contribute to the spatial SMB variability (Yang, 2013; Brun, 2015).

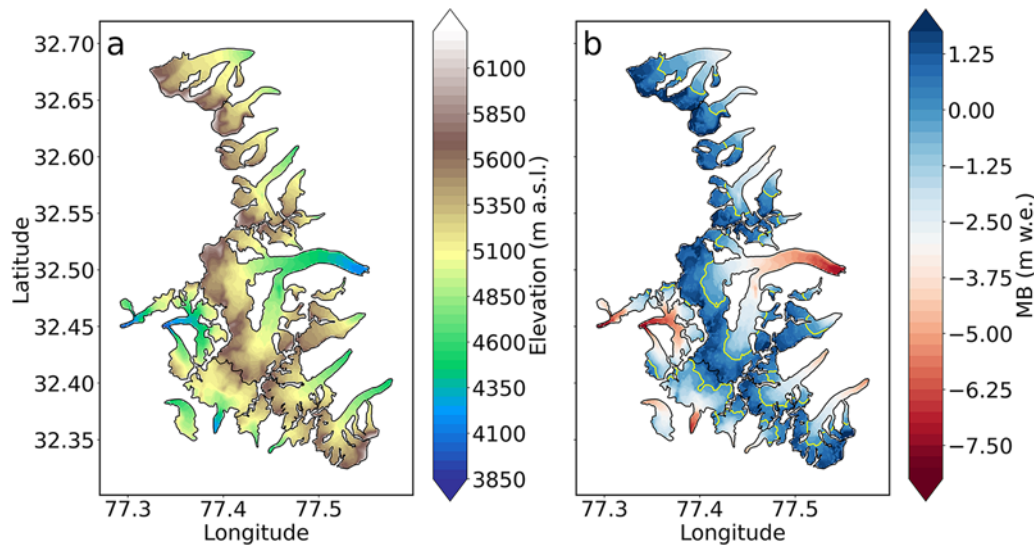
**Table 3.** Various methods of mean mass balance for Chandra Basin glaciers

Mass Balance Method	Mean Mass Balance (m w.e. $a^{-1}$ )	Year	Reference
Geodetic*	$-0.47 \pm 0.50$	1989–2020	Chandrasekharan and Ramsankaran (2023)
VIC Model*	$-0.18 \pm 0.14$	1980–2018	Laha and others (2023)
Energy balance <sup>+</sup>	$-0.82 \pm 0.21$	2015–17	Oulkar (2022)
Energy balance <sup>+</sup>	$-0.59 \pm 0.12$	2013–18	Patel (2021)
Glaciological <sup>@</sup>	$-0.57 \pm 0.12$	2013–17	Sharma and others (2020)
Glaciological <sup>§</sup>	$-0.46 \pm 0.40$	2002–19	Mandal (2020)
Geodetic*	$-0.31 \pm 0.08$	2000–16	Shean (2020)
Geodetic*	$-0.13 \pm 0.14$	1975–2000	Maurer and others (2019)
Geodetic*	$-0.48 \pm 0.15$	2001–16	Maurer and others (2019)
Geodetic*	$-0.30 \pm 0.10$	2000–15	Mukherjee and others (2018)
Geodetic*	$-0.37 \pm 0.09$	2000–16	Brun and others (2017)
Geodetic*	$-0.61 \pm 0.46$	1984–2012	Tawde and others (2017)
Geodetic*	$-0.37 \pm 0.09$	2000–16	Brun and others (2017)
Glaciological <sup>§</sup>	$-0.56 \pm 0.40$	2002–14	Azam (2016)
Geodetic*	$-0.52 \pm 0.32$	2000–12	Vijay and Braun (2016)
Geodetic <sup>^</sup>	$-1.44 \pm 0.69$	2001–12	Mishra and others (2014)
Geodetic*	$-0.68 \pm 0.15$	1999–2011	Gardelle and others (2013)
Geodetic*	$-0.44 \pm 0.09$	1999–2011	Vincent (2013)

Region: \*Chandra Basin, <sup>+</sup>Sutri Dhaka, <sup>@</sup>Eight glaciers, <sup>@</sup>Five glaciers, <sup>§</sup>Chhota Shigri, <sup>^</sup>Hampta.

### 4.4. Non-climatic parameters and mass-balance spatial variability

The upper Chandra Basin has varying topography, including differences in elevation, size, aspect, orientation and slope, as detailed in Table S2. Higher elevations ( $>5500$  m a.s.l.) have lower temperatures, humidity, air pressure and different precipitation patterns (Bhattacharya, 2023). This leads to more positive or balanced SMB at and above median elevations but increasing negative SMB at lower elevations (Fig. 5a and Table S2 in the Supplementary). For example, glaciers in the study area lost 31% of its SMB within the 4700–5400 m a.s.l. elevation range and remaining SMB loss over lower elevation. Furthermore, higher elevations receive more solid precipitation due to colder temperatures,



**Figure 5.** (a) Digital elevation model (DEM) and (b) distributed mean surface mass balance for the glaciers of the upper Chandra Basin for the period from October 2015 to September 2022. The yellow line indicates the equilibrium line altitude (ELA).

resulting in more accumulation (Bhattacharya, 2023). However, meteorological conditions can vary at different elevations at particular seasons, such as the onset or end of summer when snowfall may occur in upper accumulation zones while melt continues in lower ablation zones (Cuffey and Paterson, 2010). The size and shape of a glacier influence its SMB by affecting accumulation and ablation patterns. For example, the larger Samudra Tapu Glacier had a more negative SMB compared to a smaller glacier (Fig. 5b). Figure 5 depicts that the glaciers with more extensive area coverage at higher elevations are likely to accumulate more mass from solid precipitation, while glaciers with greater area at lower elevations may experience higher ablation rates due to warmer temperatures (Racoviteanu and others, 2015).

Furthermore, our results indicate that southeast, south, southwest and west-facing aspects correspond to the zones with the highest average rates of ablation in the study area (Figs. 5b and 6a). This is likely due to the effects of topographic shading on spatial variability in SMB. Glaciers on northeast-facing slopes receive less direct solar radiation, and experience reduced ablation compared to southwest-facing glaciers (Figs. 5b and 6a), which is consistent with findings by Olson and Rupper (2019). Mountain shadows can partially cover glaciers, limiting sunlight exposure and reducing melting, which is in agreement with previous studies (Klok and Oerlemans, 2002; Wang, 2022). Topographic shading directly alters the radiation budget at the glacier surface by obstructing incoming solar radiation (Olson and Rupper, 2019; Wang, 2022; Zhang and others, 2024). Our results indicate that topographic shading plays a crucial role in controlling SEB and SMB estimates.

The slope of selected glaciers in the upper Chandra Basin ranges from  $1^\circ$  to  $57^\circ$  with a mean of  $15^\circ$  (Fig. 6b and Table S2 in the Supplementary). However, the majority of the glacier area in the study region has a gentle slope of  $12\text{--}20^\circ$  (Fig. 6b and Table S2 in the Supplementary). These moderate-slope zones, comprising most of the glacier area, experience the highest melting rates, agreeing with other findings (Fischer and others, 2015; Rabatel and others, 2016; Kumar, 2021). In the lower ablation zone of a glacier, the slope is gentle, resulting in a low-temperature gradient in most parts of the Himalaya (Zhang and others, 2022). Consequently, this area experiences higher rates of melting. In contrast, the accumulation zone

of the glacier has a steeper slope, leading to a higher temperature gradient and lower rates of melting across the High Mountain Asia glaciers (Zhang and others, 2022). The steep slopes in upper accumulation zones are less susceptible to regional mass loss (Fischer and others, 2015; Rabatel and others, 2016; Davaze and others, 2020; Kumar, 2021). More gentle slopes correlate with more negative SMB over lower ablation zones, as also reported by Davaze and others (2020) for glaciers in the European Alps.

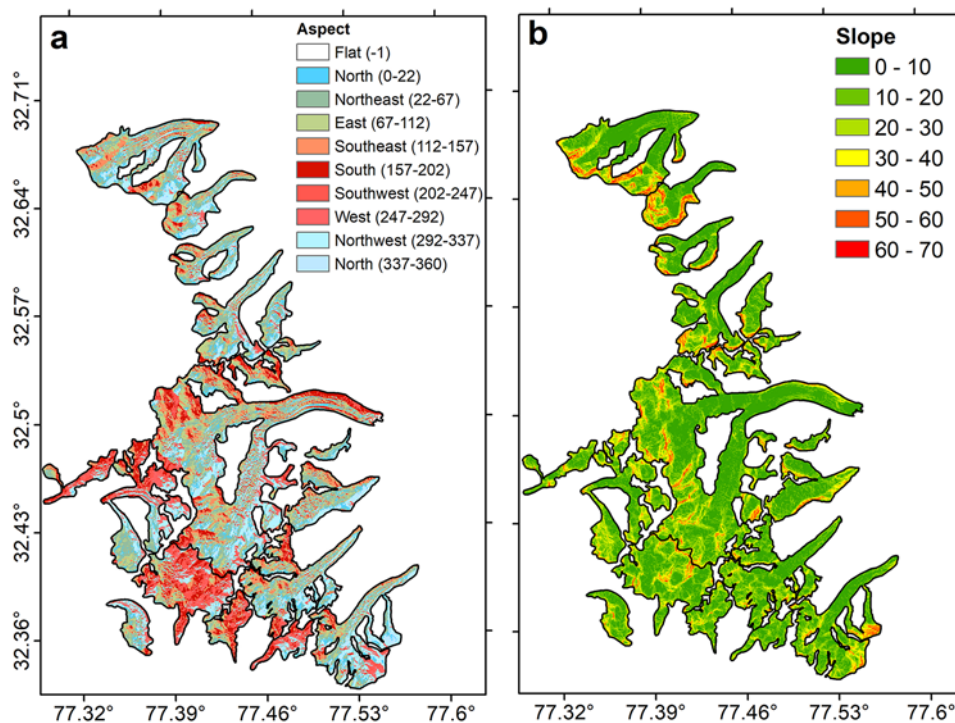
We find that the aspect, slope, size and elevation of the glacier contribute towards the spatial variability of SMB within the study region, consistent with previous studies that have examined the complex relationship between topography and glacier melt dynamics.

#### 4.5. Glaciers mass-balance sensitivity to climatic conditions

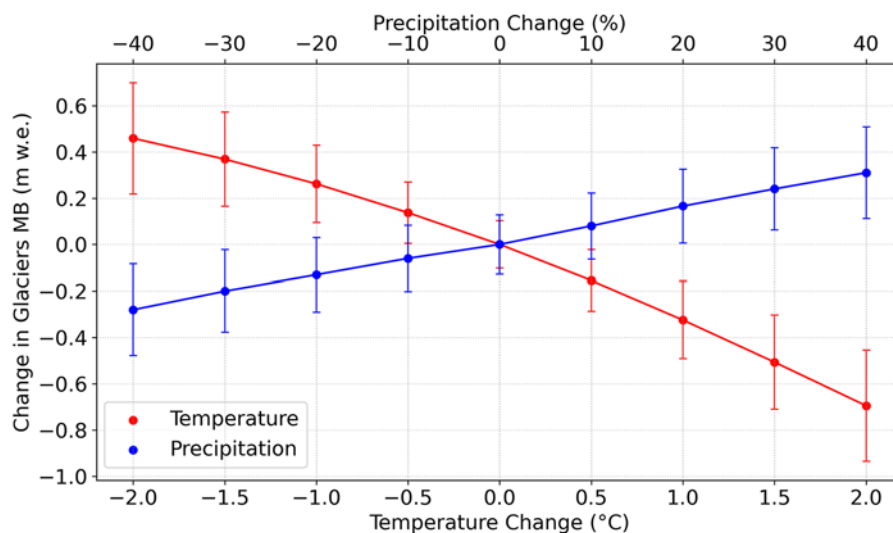
Variations in air temperature and precipitation play pivotal roles in driving glacier SMB changes, as explored through simulations conducted to investigate their impact on the SMB of the Sutri Dhaka Glacier (Oulkar, 2022). Therefore, to evaluate the climate sensitivity of the upper Chandra Basin glacier SMB, perturbation experiments of air temperature and precipitation are performed. During these simulations, other pertinent variables (incoming shortwave radiation, incoming longwave radiation, relative humidity, surface pressure, wind speed and cloud cover) remained constant. Additionally, a coupled parameter perturbation approach is employed, wherein alterations in both air temperature and precipitation are simultaneously introduced as forcings to the model. This approach mainly sought to address the extent to which changes in precipitation could compensate for the additional loss of glacier mass attributed to a  $1^\circ\text{C}$  increase in air temperature.

To simulate temperature variations, we changed the air temperature by a step of  $0.5^\circ\text{C}$  from  $-2^\circ\text{C}$  to  $+2^\circ\text{C}$  while keeping other input parameters constant (Fig. 7). The results show that a temperature increase of  $1^\circ\text{C}$  led to a 64% reduction in the total SMB (an increase in mass loss from mean SMB). In contrast, a temperature decrease of  $1^\circ\text{C}$  resulted in a 51% increase in the SMB (a decrease in mass loss) (Fig. 7). Moreover, when the temperature is increased





**Figure 6.** The upper Chandra Basin map showing (a) aspect, (b) slope.



**Figure 7.** Sensitivity of mass balance in the upper Chandra Basin glaciers to air temperature and precipitation changes.

by 2°C, the total SMB decreases by 136%. In contrast, a temperature decrease of 2°C led to a 90% increase in the total SMB. Similarly, we conducted the simulations to explore the effects of precipitation changes using eight scenarios, where the precipitation is adjusted in 10% increments from -40% to +40%. Increasing the precipitation by 20% resulted in a 32% increase in the total SMB, while a 20% decrease led to a 25% reduction in the total SMB. Notably, at a 40% increase in precipitation, the total SMB exhibited a slightly less negative equivalent to a 60% increase in glacier SMB.

Further, the temperature and precipitation have different elevation gradient, indicating that temperature and precipitation have different effects on glacier SMB (Fig. 7). Notably, augmenting precipitation yields a slightly positive effect compared to reducing precipitation (Fig. 7). Across the upper Chandra Basin, the

mass-balance sensitivity for air temperature and precipitation change is  $-0.29 \text{ m w.e. a}^{-1} \text{ }^{\circ}\text{C}^{-1}$  and  $0.14 \text{ m w.e. a}^{-1} (10\%)^{-1}$ , respectively. However, the impact of temperature changes on glacier SMB is more pronounced than that of precipitation changes, as the temperature shows a steeper gradient in comparison to the precipitation (Fig. 7). This is consistent with the findings of previous studies that have shown that temperature is the dominant factor controlling glacier SMB globally (Azam, 2018; Singh and others, 2018; Fugger, 2022). However, precipitation can still influence glacier SMB by affecting the amount and type of precipitation, snowpack density, albedo and run-off. Furthermore, it is observed that higher air temperatures accelerated the rate of glacier mass loss, while the compensatory effect of increased precipitation for change in glacier SMB is gradually limited (Fig.

7). Our findings reveal that a 42% increase in precipitation is necessary to counteract the additional mass loss resulting from a 1°C increase in air temperature for the upper Chandra Basin glaciers.

## 5. Conclusions

We investigated the distributed SEB and SMB of glaciers in the upper Chandra Basin of the western Himalaya, using a physically based COSIPY model, to estimate energy fluxes and SMB for the Chandra Basin glaciers for 7 hydrological years from 2015 to 2022. Meteorological data from HBC AWS and bias-corrected ERA5 data are employed as input for the model, and the parameters are calibrated using in situ observations. The study addressed uncertainties inherent in the modelling processes through a Monte Carlo simulation approach. The study highlighted the strong seasonality of energy fluxes, with net radiation being the dominant energy flux during summer months, while sensible and latent heat fluxes dominated in the winter. The results also revealed the spatial variability in energy fluxes and SMB across the glaciers within the upper Chandra Basin, highlighting the additional influence of factors like glacier geometry, shading effects, local topography and orientations/aspects. The estimated SMB of the upper Chandra Basin glaciers indicated an overall negative mean annual SMB of  $-0.51 \pm 0.28$  m w.e.  $a^{-1}$  over the 7 year period. This underscored the impact of climate change on glacier mass loss in the region. Our study demonstrated that the glacier SMB is highly sensitive to changes in air temperature and precipitation, with even small temperature variations causing significant shifts in the SMB. The findings reveal that a 42% increase in precipitation is necessary to counteract the additional mass loss resulting from a 1°C increase in air temperature in the region. The comparison of model results with other SMB studies indicated consistency with the model output and validated the accuracy of the model employed. Our study offers a well-constrained distributed energy and mass balance of glaciers in the upper Chandra Basin and can be used to better understand the impacts of climate variability on the SMB of Himalayan glaciers.

**Supplementary material.** The supplementary material for this article can be found at [10.1017/aog.2024.46](https://doi.org/10.1017/aog.2024.46).

**Data availability statement.** The data used in this analysis are freely available at the National Centre for Polar and Ocean Research (NCPOR) data repository. Himansh Base Camp (HBC) AWS data is available at [https://data.ncpor.res.in/static/datasets/him/1\\_himansh\\_station\\_aws\\_4052.xlsx](https://data.ncpor.res.in/static/datasets/him/1_himansh_station_aws_4052.xlsx) and point mass balance for Sutri Dhaka and Samudra Tapu Glacier is available at [https://data.ncpor.res.in/static/datasets/him/2\\_Glaciological\\_Mass\\_Balance\\_2015\\_22.xlsx](https://data.ncpor.res.in/static/datasets/him/2_Glaciological_Mass_Balance_2015_22.xlsx).

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**Competing interests.** The authors declare no competing interests.

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