

# Radiation from Brunel-induced Langmuir waves in ultra-relativistic laser–plasma interactions

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## Abstract

Intense laser light incident on solid targets has been shown to be a prolific source of harmonics. High harmonic intensities are characterized by a power-law spectrum  $P_m \sim m^{-p}$ , where  $p$  denotes a spectral decay index and  $m$  is the harmonic number. Across a wide range of light intensity and target plasma density, particle-in-cell (PIC) simulations have shown  $p = 8/3$ , a value supported by the observation. However, the claim that this decay is universal has been contested and shown to break down in simulations when the incident light is  $P$ -polarized. Here weaker decays with  $p = 5/3$  and at higher intensities  $p = 4/3$  were observed. The distinction between the two regimes was attributed to the contribution to the spectrum from emission at the plasma frequency and its harmonics from Langmuir waves excited by the Brunel electrons generated when the incident light is  $P$ -polarized. In the present work, a single-particle model has been devised to lend support to a wide range of PIC data. The model incorporates a Langmuir electric field and is hybrid in the sense that we have made use of PIC output to guide our choice of field amplitudes in the Langmuir source. We find that the spectrum computed in this way shows generally satisfactory agreement with the corresponding PIC spectra, both in decay coefficient and spectral cut-off. At the highest intensities considered the emission from the convective bunching of electrons in high-amplitude Langmuir waves is itself superseded by Brunel electron bremsstrahlung and we have identified the regions of parameter space over which each of these sources is dominant. Interestingly, the same distinction has been drawn in a very different plasma regime where comparably complex spectra have been predicted in gamma-ray burst spectra from Langmuir turbulence produced in astrophysical jets.

**Keywords:** High harmonic generation; Ultra-relativistic laser–plasma interactions

## 1. INTRODUCTION

High harmonic generation in relativistic interactions from femtosecond laser pulses on solid density targets has stimulated interest in a number of potential applications in plasma diagnostics and technological developments in ultrafast optics for studying dynamical processes in matter, and more recently as sources for generating ultrashort-intense pulses in the attosecond range (Teubner & Gibbon, 2009).

It has long been known from the experiments and particle simulations that laser-generated harmonics from the interaction of a highly intense light pulse with overcritical density plasmas are characterized by a power-law decay that can extend over hundreds or even thousands of harmonic orders. Across this range, the scaling of the intensity  $I$  with harmonic order  $m$  is

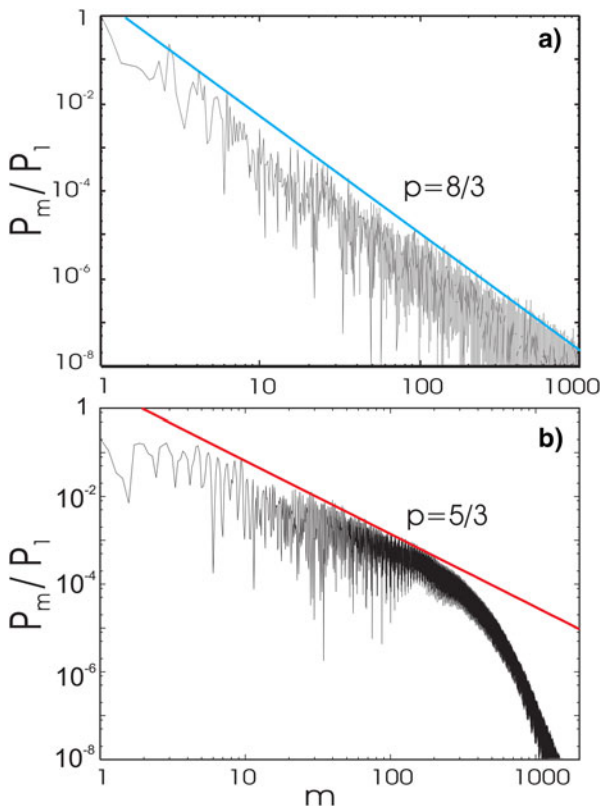
well represented by a power-law  $I \propto m^{-p}$ . Gibbon (1996) first proposed an empirical scaling on the basis of particle-in-cell (PIC) simulations, a scaling subsequently borne out by observation (Norreys *et al.*, 1996). These indicated that  $p$  lay in a range  $p = 5.5$  at  $a_0 \approx 0.6$ , decreasing to  $p \approx 3.3$  at  $a_0 = 2.9$ , where  $a_0 = 8.54(I_{20}\lambda_L^2)^{1/2}$  is the normalized quiver momentum, with  $I_{20}$  denoting the light intensity in units of  $10^{20}$  W cm<sup>-2</sup> and  $\lambda_L$  the laser wavelength measured in microns.

This trend persisted as  $I\lambda_L^2$  increased beyond these moderately relativistic levels into the ultra-relativistic (UR) range. Within this range, Baeva, Gordienko, and Pukhov (BGP; Baeva *et al.*, 2006) devised a similarity model in which the interaction physics remains invariant for constant values of a similarity parameter  $S = n_{e0}/n_c a_0$ , where  $n_{e0}/n_c$  is the ratio of the initial electron plasma density to the critical density. This model showed that  $p = 8/3$ , a decay claimed, at the time, to be universal. The BGP analysis led moreover, to a spectral cut-off at  $m_c = \sqrt{8\alpha}\gamma_s^3$ , where  $\gamma_s$  is the relativistic factor associated with the plasma surface, and  $\alpha$  is of order unity.

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This would indeed be a remarkable result were it truly universal. However, Boyd and Ondarza-Rovira (BOR) showed that while a  $p = 8/3$  decay is characteristic of harmonic spectra across the range of intensities considered in (Boyd & Ondarza-Rovira, 2010a, b), this was only true when the incident laser light was S-polarized. Figure 1a is representative of S-polarized spectra across a wide range of values of the parameters that define the BGP similarity index,  $S$ , viz.  $n_{e0}/n_c$  and  $a_0$ . This shows that to a good approximation,  $p = 8/3$  across a harmonic range  $10 \lesssim m \lesssim 1000$ . In this case,  $n_{e0}/n_c = 100$ ,  $a_0 = 20$ , and the pulse length  $t_p = 17$  fs. Tellingly, support for  $p = 8/3$  decay was soon forthcoming from measured spectra where Dromey *et al.* (2007) found  $p$  in a range 2.4–2.8 in experiments with  $a_0 = 13$ .

However, PIC simulations using P-polarized light tell a different story. Once again, the simulations encompassed a wide range of  $n_e/n_c$  and  $a_0$ , as well as different pulse lengths and pulse shapes; the spectra however, did not fit the so-called universal model. These results led BOR to contest the BGP claim that the index  $p = 8/3$  was truly universal. A typical spectrum in the case of a P-polarized incident light pulse is shown in Figure 1b for the parameters used in Figure 1a. In this case, the spectrum is best fitted by a decay index  $p = 5/3$  across the harmonic range of interest. BOR found that across a range of parameter combinations chosen such that  $1 \leq S \leq 10$ , with  $5 \leq a_0 \leq 30$ , the decay index was typically  $p = 5/3$ .



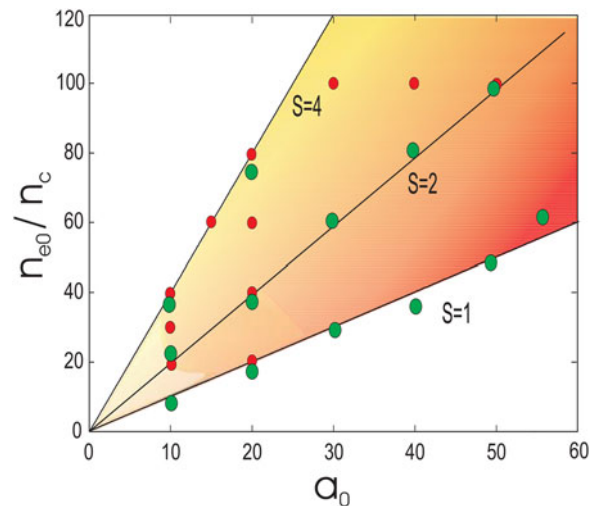
**Fig. 1.** Harmonic spectra from laser pulses incident obliquely on a plane target plasma for (a) S-polarization and (b) P-polarization of the incident light. Laser and plasma parameters:  $a_0 = 20$ ,  $n_e/n_c = 100$ , and  $t_p = 17$  fs.

However as the light intensity is increased beyond this, the simulations showed that the decay index relaxed further to  $p = 4/3$ . At the highest intensities used in the simulations ( $a_0 = 100$ ), it appeared lower still at  $p = 2/3$ , though at this end of the scale, noise levels degrade the spectral resolution. BOR characterized aspects of this emission, showing that the deviation seen in the spectra from a  $p = 8/3$  decay reflected details of the interaction physics in which emission at the plasma frequency  $\omega_p$  and its harmonics distort the spectrum of laser harmonics. Plasma waves are excited by bunches of UR electrons, generated by Brunel absorption (electrons that were first pulled out and then strongly accelerated back into the plasma by the laser electric field). These in turn couple to the radiation field in the presence of strong density gradients at the plasma surface.

Across a range  $1 \leq S \leq 4$ , BOR found that within the parameter space represented in Figure 2, the decay index was typically  $p = 5/3$ . As the light intensity increases, that is, as  $S$  decreases, the simulations show that the decay index relaxed further to  $p = 4/3$ . However, there appeared to be no marked delineation in this parameter space between the two decays, in contrast to the distinction between the two polarizations. Broadly speaking, the gradation in coloring shown in Figure 2 should be viewed more as an indication of the zones in parameter space across which either of these indices prevail, rather than any hard and fast delineation.

## 2. MODELING THE LANGMUIR SOURCE ELECTRIC FIELD

Modeling the spectrum of high harmonics along with the plasma emission is a challenge not yet met. Some light can



**Fig. 2.** Data points from PIC experiments and from the theoretical model introduced in Section 3 spanning the region of parameter space ( $n_{e0}/n_c$ ,  $a_0$ ) enclosed by the lines  $S = 1$ ,  $S = 4$ . Regions colored yellow (red) correspond to those over which  $p = 5/3$  ( $p = 4/3$ ) spectral decay prevails; the fading of one coloring into the other is intended to indicate that there is no clearly delineated region of the parameter space across which one or other index prevails. Green bullets correspond to PIC data and red bullets to the model described in this paper.

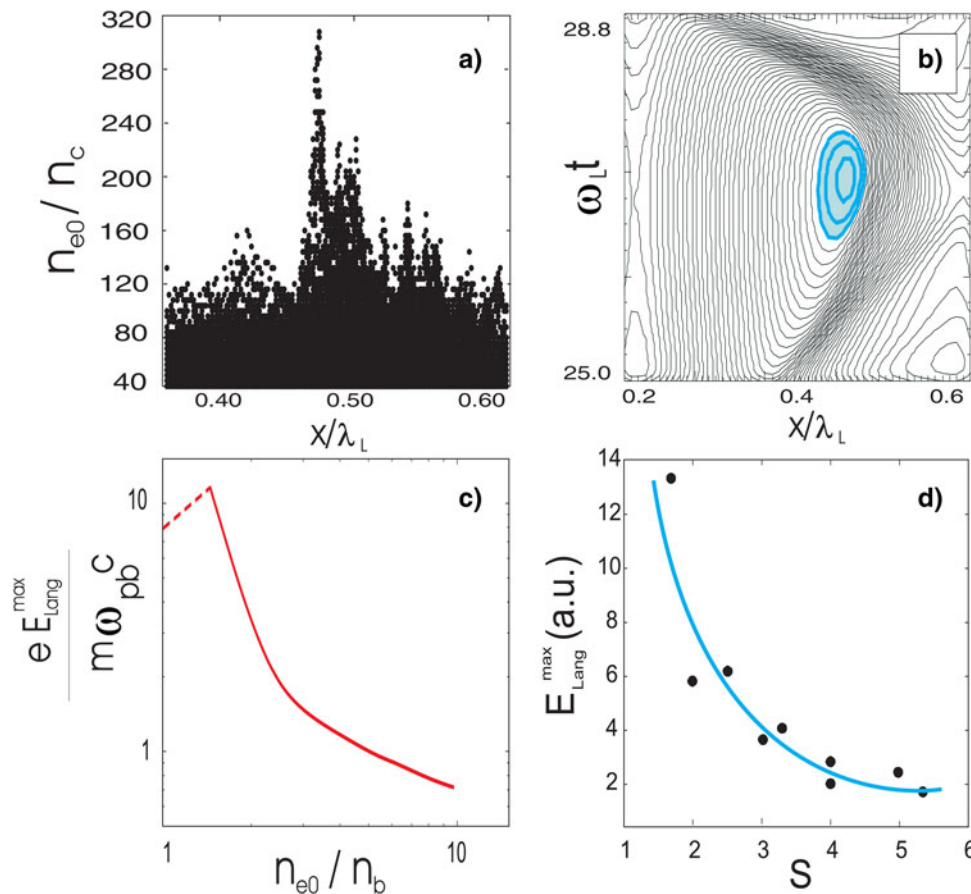
be shed on the problem, albeit piecemeal. The first step is to compare data from our PIC simulations that span a wide range of  $(n_{e0}/n_c, a_0)$  parameter space, with estimates of the Langmuir electric field computed from a cold beam–plasma model. This comparison in turn provides a basis for modeling the emission in phase 2, guiding our choice of parameters to optimize the Langmuir emission. This aspect will be outlined in Section 3.

We have discussed elsewhere the bunching of the Brunel electrons that provides the driver for the Langmuir wave (Boyd & Ondarza-Rovira, 2010b). Figure 3a, 3b show the examples of phenomena that characterize the interaction in the case of *P*-polarized light (Boyd & Ondarza-Rovira, 2010b). Figure 3a shows clear electron bunching across a region approximately  $0.2\lambda_L$  in extent, with a peak density  $n_e/n_c \sim 230$  only 15 nm in width. The corresponding electric field contours in Figure 3b represents the Langmuir field increasing to a peak at  $x/\lambda_L \sim 0.4$  before dropping away sharply.

Figure 3c, 3d recapitulate results from earlier work by Boyd & Ondarza-Rovira (2010b) which adapted a one-dimensional (1D) cold electric beam model to describe the re-entrant Brunel electrons interacting with the plasma

to drive beam-resonant Langmuir waves of wave number  $k \sim \omega_p/v_b$ . This model predicts a maximum Langmuir field  $E_{\text{Lang}}^{\text{max}} \propto [n_b/(n_{e0} - 2n_b)]^{1/2}$ , where  $n_b$  denotes the beam density of the Brunel electrons. Approximating the Brunel beam density  $n_b \sim \gamma n_c \sim a_0 n_c / \sqrt{2}$  provides an estimate of the peak Langmuir field  $E \propto (S - \sqrt{2})^{-1/2}$ . In other words, optimal plasma emission is expected at low values of *S*. Evidence from our PIC simulations (Boyd & Ondarza-Rovira, 2010b) lends support to this prediction.

Figure 3c plots the maximum Langmuir field according to this simple beam model, showing its dependence on beam density. Figure 3d shows corresponding data points from a set of 1–1/2D PIC simulations in which the parameter range was set at  $10 \lesssim a_0 \lesssim 40$ ,  $30 \lesssim n_{e0}/n_c \lesssim 100$ . We see from this that the predictions from the model are broadly in line with the outcome of the PIC simulations and on this evidence, it appears that the range of *S* across which emission from a plasma source (or sources) is dominant, is  $1 \lesssim S \lesssim 5$ . Across this region we find the evidence of plasma effects present in the spectrum of harmonics. At frequencies above the plasma frequency, harmonic decay is generally well represented by a decay coefficient  $p = 5/3$ . What happens in the



**Fig. 3.** (a) Example of electron bunching from PIC simulation where  $a_0 = 30$ ,  $n_{e0}/n_c = 50$ ; (b) corresponding contours of the Langmuir electric field  $E_x$  (Boyd & Ondarza-Rovira, 2010b); (c) maximum Langmuir field excitation  $E \propto (S - \sqrt{2})^{-1/2}$  as a function of  $n_{e0}/n_b$  from cold electron beam model; and (d) measurements of maximum Langmuir field from PIC simulations across the ranges  $10 \lesssim a_0 \lesssim 40$ , and  $30 \lesssim n_{e0}/n_c \lesssim 100$ , such that  $1 \lesssim S \lesssim 6$  (Boyd & Ondarza-Rovira, 2010b).

region below  $S = \sqrt{2}$  is more problematic. Here our PIC data are relatively sparse and what there are, do not reflect the behavior of the cold beam–plasma model in this part of the range.

### 3. SOLITON MODEL OF PLASMA EMISSION

With the Langmuir electric field characterized by means of our PIC simulations, we turn next to a representation of this field first proposed by Weatherall (1988) as a radiation source in relativistic beam devices as well as a model of galactic emission (Weatherall & Benford, 1991). This work drew attention to radiation from large-amplitude Langmuir waves excited by means of electron beam–plasma interactions, stressing the significance of electron bunching in generating high-amplitude waves. Weatherall modeled the Langmuir electric field as a dipole soliton

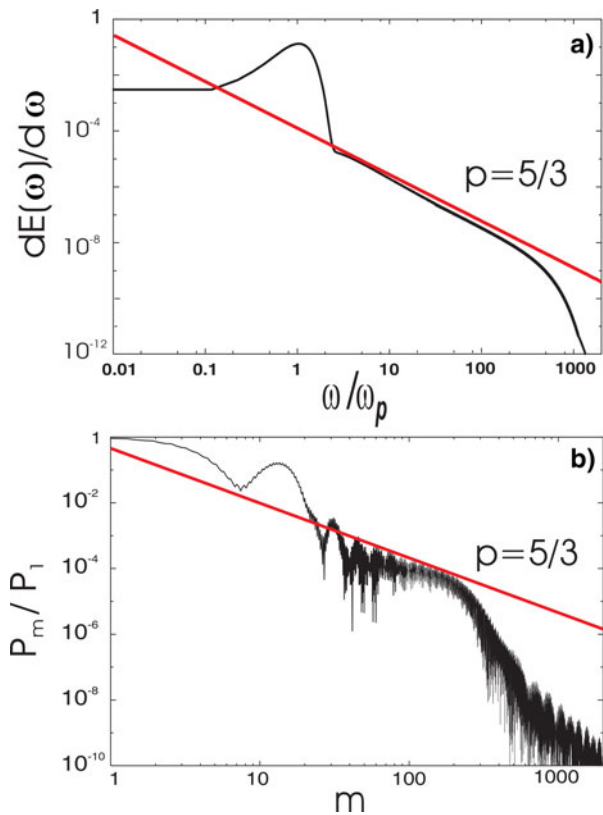
$$\mathbf{E}_{\text{Lang}} = -\frac{2}{3}\pi\rho_0 D \exp(-i\omega_p t)\hat{\mathbf{p}}, \quad (1)$$

where  $\rho_0$  is the amplitude of the electron charge density and  $D$  is the scale length of the condensate region, oscillating at the plasma frequency  $\omega_p$ , with a dipole moment oriented in the direction  $\hat{\mathbf{p}}$  (Weatherall, 1988).

In particular, Weatherall showed that for an appropriate choice of parameters, with beam bunching added to provide the degree of coherence needed, this model could account not only for strong emission at the plasma frequency and its harmonics (Weatherall & Hobbs, 1986), but additionally as a source of broadband radiation for ultra-relativistic electron beams. Figure 4a shows a power spectrum computed using Weatherall’s model. It shows, naturally, strong emission at the plasma frequency as well as a distinctive spectral cut-off at  $\omega_c = 2\gamma_b^2(c/D)$ , where  $\gamma_b$  is the relativistic gamma for the beam. Physically, this cut-off frequency corresponds to the highest frequency emitted by electrons scattered by solitons of scale length  $D$ . Over the range spanned by these frequencies, spectral decay is well represented by a power-law dependence with  $p = 5/3$ .

We have adapted the source in (1) and tailored it to our needs by choosing the parameters  $\rho_0$  and  $D$  to conform to the range of values measured in our PIC runs described in Section 2 (cf. Figure 3d). With these choices, we then integrated the relativistic Lorentz equation and applied the solution to determine the radiation field and emission spectrum. To distinguish our procedure from that of Weatherall, used to compute the spectrum in Figure 4a, while at the same time, acknowledging that we “borrowed” the Weatherall source in (1), we refer to this hybrid as the BORW model.

The BORW emission spectrum is represented in Figure 4b for a parameter choice  $a_0 = 40$  and  $n_e/n_c = 167$ . Like the spectrum in Figure 4a, this too shows not only strong emission at  $\omega_p$  but also a clear cut-off, at  $m^* \sim 300$ , while across the frequency range  $\omega_p/\omega_L \lesssim m \lesssim m^*$  the spectrum is well fitted by a power-law decay with  $p = 5/3$ .



**Fig. 4.** (a) Beam radiation power spectrum from density correlations in Weatherall’s model and (b) radiation from a laser-accelerated electron perturbed by a soliton field, radiation characterized by strong emission at the plasma frequency  $\omega_p$  and power decay index  $5/3$  up to  $m = 220$ , for  $a_0 = 40$ ,  $n_e/n_c = 167$ , and  $m = \omega_m/\omega_L$ .

Notwithstanding the similarity between the spectra in Figure 4, other than the emission peak at the plasma frequency, the two are quite distinct in their spectral detail. As already pointed out, the  $5/3$  decay seen in Figure 4a is an artifact of the assumption made about the fluctuation spectrum of beam electrons in Weatherall’s model, whereas that in Figure 4b is intrinsic in the BORW model and correlates with the spectral decay observed in our PIC experiments.

### 4. EMISSION SPECTRA FROM BORW MODEL

Figure 5 shows a comparison of output from the BORW model and a PIC simulation for a particular choice of laser–plasma parameters,  $a_0 = 60$ ,  $n_{e0}/n_c = 95$ , that is,  $S = 1.6$ , so that according to the scaling in Figure 2 we might expect a spectrum with decay  $p = 5/3$ . As it happens, our choice of parameters was prompted by the fact that this particular spectrum had been considered independently by an der Brügge and Pukhov (an der Brügge & Pukhov, 2010), in support of their claim that this spectrum derived from a synchrotron source. In the present context, this spectrum provided a ready-made comparison for our own spectrum, shown in Figure 5a. The two spectra match closely, both in decay constant  $p = 5/3$  and with a cut-off at  $m^* = 110$  followed

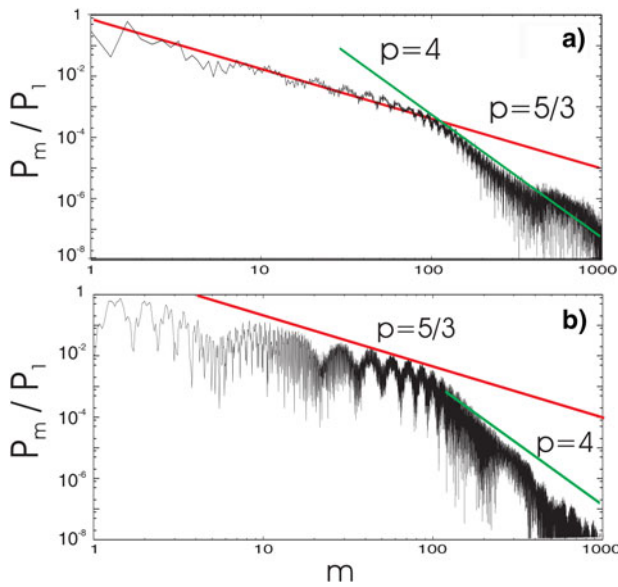


Fig. 5. PIC (a) and BORW model (b) for  $a_0 = 60$  and  $n_e/n_c = 95$  showing the transition from a region characterized by  $p = 5/3$  decay to one, where  $p = 4$ .

by a range over which spectral intensities decay as  $p = 4$ . Moreover, by applying details of the Langmuir field measured from our simulation in Figure 5a, to determine the Weatherall source (1), we then computed the BORW emission spectrum, reproduced in Figure 5b. The agreement between the two spectra is generally satisfactory, given the limitations of the BORW model.

A second comparison is shown in Figure 6 for the combination of parameters,  $a_0 = 50$ ,  $n_{e0}/n_c = 100$ , such that  $S = 2$ ;  $t_p = 10$  fs. The PIC spectrum for this choice, Figure 6a, shows some difference from that in Figure 5. Now spectral decay governed by a  $p = 4/3$  decay coefficient extends only to a cut-off  $m^* = 30$ , before a region of steep  $p = 4$  decay, akin to that found in Figure 4. At frequencies  $m \geq 100$  there is a further transition, though one not well characterized by a unique decay coefficient. Interestingly, the BORW spectrum shown in Figure 6b, likewise exhibits three more or less distinct regions: the zone governed by  $p = 4/3$  decay here extends as in Figure 5 to a  $m^* \sim 100$  cut-off before a  $p = 4$  transition zone. This in turn is truncated in a sharp reversal to a region of spectrum characterized by  $p = 2/3$  across  $200 \lesssim m \lesssim 600$ . With the hindsight afforded by the BORW spectrum, one might identify the high-frequency region in the PIC spectrum with this new feature, with a decay constant lower by far than any observed hitherto.

From PIC output we found that  $p = 2/3$  decays corresponded to the most ultra-relativistic interactions within our range, showing the highest Langmuir electric fields. In Figure 7, we present a correlation between the Langmuir fields excited over an ensemble of interactions and the power decay indices found in the corresponding spectra. It was found that the highest Langmuir fields for plasma

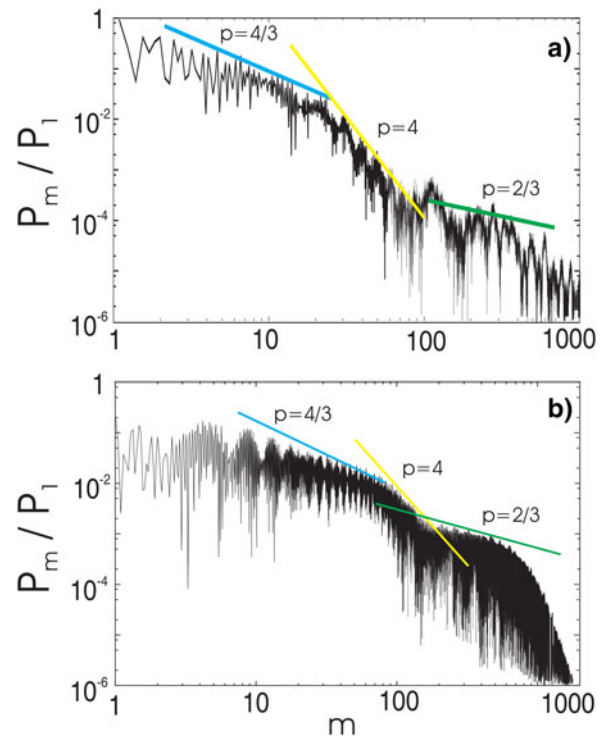


Fig. 6. (a) PIC and (b) BORW harmonic spectra for  $a_0 = 50$  and  $n_e/n_c = 100$ . The spectra show harmonic transition decays from  $p = 4/3$  to  $2/3$ .

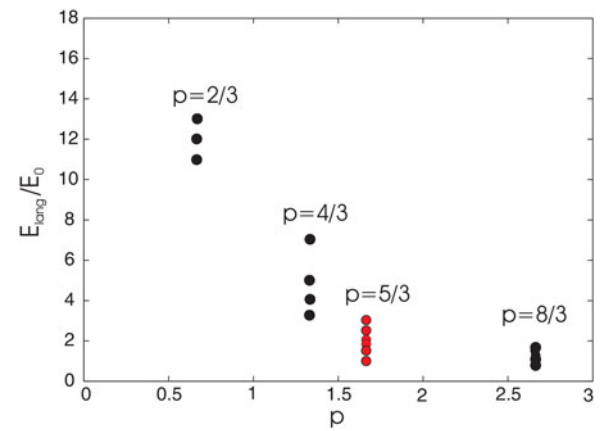


Fig. 7. Correlation between the Langmuir field  $E_{Lang}$  and the spectral decay index  $p$ .

emission corresponded to the lowest  $2/3$  indices, and that the plasma emission amplitudes decreased as the spectral decay index approached the value  $8/3$ .

### 5. SUMMARY

A particle model has been applied to determine the radiation spectrum from a target plasma irradiated by high-intensity laser light. This model is hybrid in the sense that PIC output has been used to characterize the Langmuir electric field source in the relativistic Lorentz equation. For a wide

range of combinations of the key parameters that characterize the interaction,  $n_{e0}/n_c$  and  $a_0$ , chosen, so that the similarity index  $S$  lies in the range  $1 \lesssim S \lesssim 5$ , we found that over a significant frequency range, the spectral power decay index was typically  $p = 5/3$ , rising to  $p = 4/3$  at higher intensities.

Moreover, at the highest intensities considered, there were clear indications that bremsstrahlung from the UR Brunel electrons was contributing to the spectrum, indicated by a spectral region with decay constant  $p = 2/3$ . To date there has been no experimental evidence from UR laser–plasma interactions, for either spectral decay with  $p = 5/3$ ,  $4/3$ , or  $p = 2/3$  despite some pointers from relativistic electron beam devices and from galactic emission spectra (Weatherall & Benford, 1991; Fleishman & Toptygin, 2008). With predictions from the model described in this work supporting our earlier reports of  $p = 5/3$  and  $4/3$  decays, this seems an opportune moment to appeal to experimentalists to test our predictions by looking for the effects on the spectrum presented in this contribution to ECLIM 33.

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