

Kaiwen Chen¹ and Baofang Song¹,†

¹Center for Applied Mathematics, Tianjin University, Tianjin 300072, PR China

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We investigated the linear stability of pipe flow with anisotropic slip length at the wall by considering streamwise and azimuthal slip separately as the limiting cases. Our numerical analysis shows that streamwise slip renders the flow less stable but does not cause instability. The exponential decay rate of the least stable mode appears to be $\propto Re^{-1}$ when the Reynolds number is sufficiently large. Azimuthal slip can cause linear instability if the slip length is sufficiently large. The critical Reynolds number can be reduced to a few hundred given large slip lengths. In addition to numerical calculations, we present a proof of the linear stability of the flow to three-dimensional yet streamwise-independent disturbances for arbitrary Reynolds number and slip length, as an alternative to the usual energy analysis. Meanwhile we derived analytical solutions to the eigenvalue and eigenvector, and explained the structure of the spectrum and the dependence of the leading eigenvalue on the slip length. The scaling of the exponential decay rate of streamwise independent modes is shown to be rigorously $\propto Re^{-1}$. Our non-modal analysis shows that overall streamwise slip reduces the non-modal growth, and azimuthal slip has the opposite effect. Nevertheless, both slip cases still give the Re^2 -scaling of the maximum non-modal growth and the most amplified disturbances are still streamwise rolls, which are qualitatively the same as in the no-slip case.

Key words: instability

1. Introduction

The classic pipe flow with no-slip boundary condition has been proved linearly stable to axisymmetric perturbations (Herron 1991, 2017), and numerical studies suggest that the flow is linearly stable to any perturbations at arbitrary Reynolds numbers (Meseguer & Trefethen 2003). The recent work of Chen, Wei & Zhang (2019) presented a rigorous proof of the linear stability of the flow to general perturbations at high-Reynolds-number regime. Therefore, transition to turbulence in pipe flow is subcritical via finite-amplitude perturbations (see e.g. Eckhardt *et al.* 2007; Avila *et al.* 2011).

† Email address for correspondence: baofang_song@tju.edu.cn

However, velocity slip of viscous fluid can occur on superhydrophobic surfaces (Voronov, Papavassiliou & Lee 2008; Rothstein 2010), for which slip boundary condition instead of the classic no-slip condition should be adopted for the momentum equations, and the slip boundary condition can potentially influence the stability of the flow. A simplified and widely used slip boundary condition is the Navier slip boundary condition, which has been shown to apply to many flow problems and is frequently adopted for linear stability studies (Vinogradova 1999; Lauga & Cossu 2005; Min & Kim 2005; Gan & Wu 2006; Ren, Chen & Zhu 2008; Ghosh, Usha & Sahu 2014; Seo & Mani 2016; Chattopadhyay, Usha & Sahu 2017, to list a few). For pipe geometry, although many studies have investigated the linear stability of immiscible and miscible multifluid flows with either no-slip or Navier slip boundary condition (Hu & Joseph 1989; Joseph 1997; Li & Renardy 1999; Selvam et al. 2007; Sahu 2016; Chattopadhyay et al. 2017, etc.), far fewer studies were dedicated to the linear stability of single-phase pipe flow with slip boundary condition. Průša (2009) investigated this problem and showed that, subject to Navier slip boundary condition, pipe flow becomes less stable compared with the no-slip case, however, the destabilization effect is constrained to small Reynolds numbers and is not sufficient to render the flow linearly unstable. Their results indicated that the stability property of pipe flow is not qualitatively affected by the slip boundary condition, regardless of the slip length. For its counterpart in plane geometry, i.e. channel flow, on the contrary, Min & Kim (2005) and Lauga & Cossu (2005) reported a stabilizing effect of velocity slip on the linear stability.

Usually, slip length is assumed homogeneous and isotropic, i.e. independent of position and direction at the wall in stability analysis. However, anisotropy in the effective slip length can be incurred by anisotropy in the texture pattern on superhydrophobic surfaces, such as parallel periodic slats, grooves and grates (Lecoq et al. 2004; Bazant & Vinogradova 2008; Ng & Wang 2009; Belyaev & Vinogradova 2010; Asmolov & Vinogradova 2012; Pralits, Alinovi & Bottaro 2017). For example, Ng & Wang (2009) reported a ratio of down to approximately 0.25 between the transverse slip length (in the direction perpendicular to the slats) and longitudinal slip length (parallel to the slats). The linear stability of channel flow with anisotropic slip caused by parallel micrograves was analysed by Pralits et al. (2017) using the tensorial formulation of slip boundary condition proposed by Bazant & Vinogradova (2008). Their results showed possibilities of linear instability using special alignment of the micrograves. Recently, Chai & Song (2019) studied the linear stability of single-phase channel flow subject to anisotropy in slip length by considering streamwise and azimuthal slip separately as the limiting cases, which can potentially be realized or approximated by using specially designed surface texture, e.g. specially aligned microgrates/graves, according to Bazant & Vinogradova (2008). Their results showed that streamwise slip mainly stabilizes the flow (with increased critical Reynolds number), although it surprisingly destabilizes the flow slightly in a small Reynolds number range, and that azimuthal slip can greatly destabilize the flow and reduce the critical Reynolds number given sufficiently large slip length. The critical Reynolds number can be reduced to a few hundred with a dimensionless azimuthal slip length of O(0.1), in contrast to $Re_{cr} = 5772$ for the no-slip case. Their study also indicated that Squire's theorem (Squire 1933) ceases to apply when the wall-normal velocity and vorticity are coupled via the slip boundary condition, such that the leading instability becomes three-dimensional (3-D) rather than two-dimensional (2-D) when slip length is sufficiently large, in agreement with Pralits et al. (2017). The stability of 3-D perturbations was not considered by Min & Kim (2005) and Lauga & Cossu (2005) in which Squire's theorem was seemingly assumed.

Differing from channel flow, linear instability is absent at arbitrary Reynolds numbers in classic pipe flow. This raises the question of whether the anisotropy in slip length can also cause linear instability in pipe flow. To our knowledge, this problem has not been studied in pipe geometry. The pseudospectrum analysis of classic pipe flow of Schmid & Henningson (1994) and Meseguer & Trefethen (2003) suggests that, despite the linear stability, at sufficiently large Reynolds numbers, a small perturbation to the linear operator associated with the governing equation can possibly change the stability of the system. The slip boundary condition can be thought of as a perturbation to the linear operator with no-slip boundary condition. However, Průša (2009) showed that homogeneous and isotropic slip does not change the spectrum qualitatively no matter how large the slip length (i.e. operator perturbation) is, Following Chai & Song (2019), in this work, we still consider anisotropic slip length in the limiting cases and explore the possibility of linear instability for pipe flow. Aside from the critical Reynolds number as focused on by Chai & Song (2019), here we also investigate the effects of the slip on the spectrum and on the scaling of the leading eigenvalues with Reynolds number. In addition to numerical calculations, we also perform analytical studies on the eigenvalues and eigenvectors of the 3-D yet streamwise-independent modes, and discuss their structure as well as their dependence on the slip length on a theoretical basis, which to our knowledge has not been reported in the literature.

2. Numerical methods

The non-dimensional incompressible Navier-Stokes equations read

$$\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} = -\nabla p + \frac{1}{R_{\rho}} \nabla^2 \mathbf{u}, \quad \nabla \cdot \mathbf{u} = 0, \tag{2.1}$$

where u denotes velocity and p denotes pressure. For pipe geometry, cylindrical coordinates (r, θ, x) are considered, where r, θ and x denote the radial, azimuthal and streamwise coordinates, respectively. Velocity components u_r , u_θ and u_x are normalized by $2U_b$ where U_b is the bulk speed (the average of the streamwise velocity on the pipe cross-section), length by pipe radius R and time by R/U_b . The Reynolds number is defined as $Re = U_b R/v$ where v is the kinematic viscosity of the fluid. In order to eliminate the pressure and impose the incompressibility condition, we adopt the velocity–vorticity formulation of Schmid & Henningson (1994), with which the governing equations of disturbances reduce to only two equations about the wall-normal velocity u_r and wall-normal vorticity ω . With a Fourier-spectral-Chebyshev-collocation discretization, considering perturbations of the form of $\{u_r, \omega\} = \{\hat{u}_r(r), \hat{\omega}(r)\} \exp(-\mathrm{i}(\alpha x + n\theta))$, the governing equations in the Fourier spectral space read

$$Lq + \frac{\partial}{\partial \tau} Mq = 0, \tag{2.2}$$

where

$$L = \begin{pmatrix} \mathrm{i}\alpha ReU\Gamma + \mathrm{i}\frac{\alpha Re}{r}\left(\frac{U'}{k^2r}\right)' + \Gamma(k^2r^2\Gamma) & 2\alpha n^2Re\Gamma \\ -\frac{\mathrm{i}U'}{r} + \frac{2\alpha}{Re}\Gamma & \mathrm{i}\alpha Rek^2r^2U + \phi \end{pmatrix}, \tag{2.3}$$

$$M = \begin{pmatrix} \Gamma & 0 \\ 0 & k^2 r^2 \end{pmatrix},\tag{2.4}$$

where $\tau = t/Re$ is the scaled time and unknowns are

$$q = \begin{pmatrix} \hat{\phi} \\ \hat{\Omega} \end{pmatrix} = \begin{pmatrix} -ir\hat{u}_r \\ \frac{\alpha r\hat{u}_{\theta} - n\hat{u}_x}{nRek^2 r^2} \end{pmatrix} = \begin{pmatrix} -ir\hat{u}_r \\ \frac{\hat{\omega}}{inRek^2 r} \end{pmatrix}. \tag{2.5}$$

The real number α is the axial wavenumber and n, which is an integer, is the azimuthal wavenumber. The base flow is denoted as U, $k^2 = \alpha^2 + n^2/r^2$, $i = \sqrt{-1}$ and the prime denotes the derivative with respect to r. The operators Γ and ϕ are defined as

$$\Gamma = \frac{1}{r^2} - \frac{1}{r} \frac{\mathrm{d}}{\mathrm{d}r} \left(\frac{1}{k^2 r} \frac{\mathrm{d}}{\mathrm{d}r} \right) \tag{2.6}$$

and

$$\phi = k^4 r^2 - \frac{1}{r} \frac{\mathrm{d}}{\mathrm{d}r} \left(k^2 r^3 \frac{\mathrm{d}}{\mathrm{d}r} \right). \tag{2.7}$$

The other two velocity components \hat{u}_x and \hat{u}_θ can be calculated as

$$\hat{u}_x = -\frac{\alpha}{k^2 r} \frac{\partial \hat{\Phi}}{\partial r} - n^2 r \hat{\Omega}, \quad \hat{u}_\theta = -\frac{n}{k^2 r^2} \frac{\partial \hat{\Phi}}{\partial r} + \alpha n r R e \hat{\Omega}.$$
 (2.8a,b)

We use the Robin-type Navier slip boundary condition at the pipe wall for streamwise and azimuthal velocities separately, i.e.

$$\left. \left(l_x \frac{\partial u_x}{\partial r} + u_x \right) \right|_{r=1} = 0, \quad \left(l_\theta \frac{\partial u_\theta}{\partial r} + u_\theta \right) \right|_{r=1} = 0, \quad (2.9a,b)$$

where $l_x \geqslant 0$ and $l_\theta \geqslant 0$ are streamwise and azimuthal slip lengths, respectively, and are independent of each other. In spectral space, these boundary conditions apply identically to \hat{u}_x and \hat{u}_θ given the homogeneity of the slip length. We use the no-penetration condition for the wall-normal velocity component at the wall, i.e. $u_r(1, \theta, x, t) = 0$. Lauga & Cossu (2005) and Chai & Song (2019) considered the same boundary conditions for slip channel flow. Note that in the isotropic slip case considered by Průša (2009), l_x and l_θ are related as $l_\theta = l_x/(1 + l_x)$, which gives $l_\theta \approx l_x$ for small slip lengths. With boundary condition (2.9a,b), given that we impose the same volume flux as in the no-slip case, i.e.

$$\int_0^1 U_x(r)r \, \mathrm{d}r = \frac{1}{4},\tag{2.10}$$

the velocity profile of the constant-volume-flux base flow reads

$$U(r) = \frac{1 - r^2 + 2l_x}{1 + 4l_x}\hat{\mathbf{x}},\tag{2.11}$$

where \hat{x} represents the unit vector in the streamwise direction. Note that the base flow is independent of l_{θ} . Converting to the $(\hat{\Omega}, \hat{\Phi})$ system, the boundary condition (2.9a,b) reads

$$\frac{\alpha}{k^2} \frac{\partial \hat{\Phi}}{\partial r} + n^2 Re \hat{\Omega} + l_x \left(n^2 Re \frac{\partial \hat{\Omega}}{\partial r} + \frac{\alpha}{k^2} \frac{\partial^2 \hat{\Phi}}{\partial r^2} + \alpha \frac{n^2 - \alpha^2}{(n^2 + \alpha^2)^2} \frac{\partial \hat{\Phi}}{\partial r} \right) = 0$$
 (2.12)

and

$$\alpha n Re \hat{\Omega} - \frac{n}{n^2 + \alpha^2} \frac{\partial \hat{\Phi}}{\partial r} + l_{\theta} \left(\alpha n Re \hat{\Omega} + \alpha n Re \frac{\partial \hat{\Omega}}{\partial r} - \frac{n}{n^2 + \alpha^2} \frac{\partial^2 \hat{\Phi}}{\partial r^2} + \frac{2n\alpha^2}{(n^2 + \alpha^2)^2} \frac{\partial \hat{\Phi}}{\partial r} \right) = 0.$$
(2.13)

It can be seen that $\hat{\Omega}$ and $\hat{\Phi}$, i.e. \hat{u}_r and $\hat{\omega}$, are coupled via the slip boundary condition.

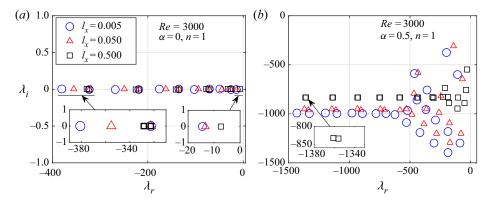


Figure 1. Spectrum of the flow at Re = 3000 with $l_x = 0.005$ (circles), 0.05 (triangles) and 0.5 (squares). (a) The mode $(\alpha, n) = (0, 1)$. (b) The mode $(\alpha, n) = (0.5, 1)$.

In order to avoid the singularity at the pipe centre, i.e. r=0, the domain [0,1] is extended to [-1,1] and an even number of Chebyshev grid points over [-1,1] are used such that there is no grid point at r=0. This extension also allows us to use the Chebyshev collocation method for the discretization in the radial direction and the resulted redundancy is circumvented by setting proper parity conditions on $\hat{\Phi}$ and $\hat{\Omega}$ with respect to r (Trefethen 2000; Meseguer & Trefethen 2003). In this way, no explicit boundary condition is imposed at the pipe centre.

To determine whether a mode (α, n) is linearly stable or not, one only needs to calculate the eigenvalues of the operator $-M^{-1}L$ and check if any eigenvalue has a positive real part, λ_r , which determines the asymptotic growth/decay rate of the corresponding eigenvector as $t \to \infty$.

3. Streamwise slip

We consider the case of $l_x \neq 0$ and $l_\theta = 0$ as the limiting case of streamwise slip being significant and azimuthal slip being negligible.

The effect of the slip on the spectrum is investigated for Re = 3000 and is shown in figure 1 for the modes $(\alpha, n) = (0, 1)$ and (0.5, 1). First, figure $1(\alpha)$ shows that the eigenvalues of the $(\alpha, n) = (0, 1)$ mode visually all fall on the $\lambda_i = 0$ line $(\lambda_i$ denotes the imaginary part of the eigenvalue) and in the left half-plane, which suggests that the eigenvalues are all real and negative. Meseguer & Trefethen (2003) reported the same finding for the no-slip case in a large Reynolds number range up to 10⁷. In fact, the eigenvalues being real and negative can be rigorously proved, see our proof in § 5.1. Second, as l_x increases, it can be observed that there are two groups of eigenvalue, one of which remains constant and the other of which shifts to the right, see the two insets in figure 1(a). Specifically, as l_x is increased to 0.5, the left eigenvalue in the left inset has moved from the circle to the triangle and finally to the square while the right eigenvalue remains constant. Nonetheless, the rightmost eigenvalue increases as l_x increases (see the right inset) which indicates that the flow becomes less stable. In § 5.2, we will show that the former group corresponds to disturbances with $\Phi \neq 0$, i.e. $u_r \neq 0$ and the latter group, on the contrary, is associated with disturbances with $\Phi \equiv 0$, i.e. $u_r \equiv 0$ and, the rightmost eigenvalue belongs to the latter group. Figure 1(b) shows the case for the mode $(\alpha, n) = (0.5, 1)$. The slip does not qualitatively change the shape of the spectrum. As l_x increases, the eigenvalues overall move to the right. In addition to a horizontal shift

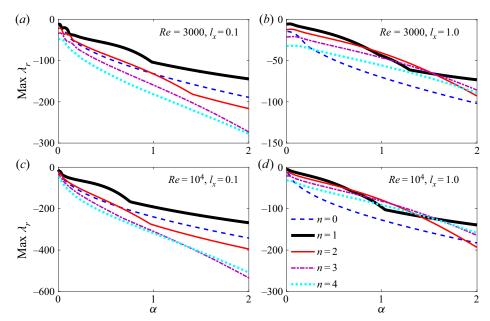


Figure 2. The maximum eigenvalue, max λ_r , as a function of α , for Re = 3000 (a,b) and 10^4 (c,d). For each Reynolds number, azimuthal wavenumbers n = 0, 1, 2, 3, 4 and slip lengths $l_x = 0.1$ and 1.0 are shown.

there is a shift in the vertical direction and, meanwhile, the spectrum is compressed in the vertical direction, see the comparison between the $l_x = 0.5$ and the other two cases. Using the term of Schmid & Henningson (1994) and Meseguer & Trefethen (2003), the horizontal branch of the spectrum (the part with $\lambda_r \lesssim -600$) corresponds to mean modes, the upper branch corresponds to wall modes and the lower branch to centre modes. Note that the speed of a wave is given by $(-\lambda_i)/\alpha Re$ in our formulation. It has been known that the wave speed of the mean modes follows the mean velocity of the '2-D' axial base flow, i.e. $\int_0^1 U_x(r) dr$ in pipe flow (see e.g. Drazin & Reid 1981), which gives 2/3 in the no-slip case (Schmid & Henningson 1994). In our case, the wave speed of the mean modes is decreased by the slip, reducing to 0.5559 for $l_x = 0.5$ (833.868/(0.5 × 3000), see the eigenvalue in table 1) which is very close to 5/9 given by $\int_0^1 U_x(r) dr$ with the base flow shown in (2.11). The wall modes, which are located close to the wall, move at lower speed than the centre modes, which are located close to the pipe centre and move at speeds close to the centreline velocity. Since we fix the volume flux of the flow while the slip length is varied, the speed of the base flow close to the wall increases as l_x increases, whereas the speed near the pipe centre decreases, i.e. the velocity profile becomes flatter, see the base flow given by (2.11). Therefore, it can be expected that as l_x increases, the speed of the wall modes increases and that of the centre modes decreases, and all three types of modes move at closer speeds. This is exactly what the compression in the vertical direction of the spectrum reveals. The other noticeable effect is that the slip brings the adjacent eigenvalues associated with the mean modes closer as the slip length increases, causing a seeming degeneracy of the spectrum, see figure 1(b).

Figure 2 shows the maximum of the real part of the eigenvalue, max λ_r , as a function of the streamwise wavenumber, α , for Re = 3000 and 10^4 . For each Re, slip lengths $l_x = 0.1$ and 1.0, and azimuthal wavenumbers n = 0, 1, 2, 3 and 4, are considered. The trend shown in the figure suggests that, for both Reynolds numbers, $\alpha = 0$ is nearly the least stable

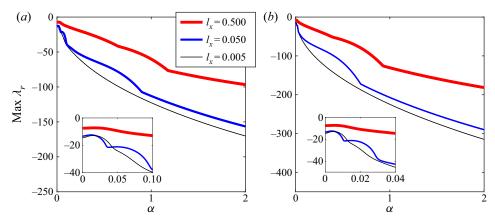


Figure 3. The influence of streamwise slip on max λ_r of n=1 modes for (a) Re=3000 and (b) $Re=10^4$. Slip lengths of $l_x=0.005$ (thin black), 0.05 (blue) and 0.5 (bold red) are shown. The insets show the close-up of the regions with very small α .

mode, i.e. the slowest decaying mode given that all max λ_r are negative, regardless of the slip length. At small α , where max λ_r is largest, the results suggest that n=1 is always the least stable one. At larger α , however, n=1 is still the least stable when l_x is small, see the case of $l_x=0.1$ in figure 2(a,c), but is not in a range of α around $\alpha=1$, see the case of $l_x=1.0$ in figure 2(b,d). Nevertheless, in this range, max λ_r is much smaller than that in the small α regime. Therefore, as we are most interested in the least stable mode, in the following, we will focus on the n=1 modes. In fact, for the $\alpha=0$ modes, we can rigorously prove that n=1 is the least stable azimuthal wavenumber, see appendix B.

Figure 3 shows $\max \lambda_r$ as a function of α of the n=1 modes for (a) Re=3000 and (b) $Re=10^4$. For each Re, overall $\max \lambda_r$ increases as l_x increases, i.e. the n=1 modes decay more slowly as l_x increases. The insets show the close-up of the small α region, in which the dependence of $\max \lambda_r$ on α is not monotonic, with the maximum appearing at some small but finite α instead of $\alpha=0$. Nevertheless, the difference between the peak value and the value for $\alpha=0$ is very small, i.e. $\alpha=0$ is nearly the least stable mode, as aforementioned. In fact, the dependence on l_x is not fully monotonic either, see the very small region around $\alpha=0.03$ for the $l_x=0.005$ (the thin black line) and $l_x=0.05$ (the blue line) cases as shown in the inset in figure 3(a) and around $\alpha=0.01$ in the inset in figure 3(b). However, for $\alpha=0$ and in most range of α , our results show a monotonic increase of $\max \lambda_r$ as l_x increases.

Figure 4 illustrates the dependence of max λ_r of the n=1 modes on l_x in a broader range of l_x . For each Re, $\alpha=0$, 0.1, 0.5, 1 and 2 are shown. The trend shows that as l_x keeps increasing, max λ_r seems to asymptotically approach a plateau with a negative value, i.e. all the modes shown in the figure appear to be linearly stable, for both Reynolds numbers.

The above results suggest that, with streamwise slip, the flow is linearly stable to any perturbations, regardless of the slip length. In order to show evidence in a broader parameter regime, we numerically searched for the global maximum of $\max \lambda_r$ over α and n and explored a wider range of l_x up to 10 and of Re up to 10^6 . Practically, based on our analysis, we only need to search in a small range of α immediately above zero (see the insets in figure 3) while setting n=1. Specifically, we search the range of [0, 1.2] at Re=100, and the range is decreased as Re^{-1} for higher Reynolds numbers. Then we plotted the global maximum of $\max \lambda_r$, still denoted as $\max \lambda_r$, as a function of l_x , for a few Reynolds numbers ranging from 10^2 to 10^6 in figure 5(a).

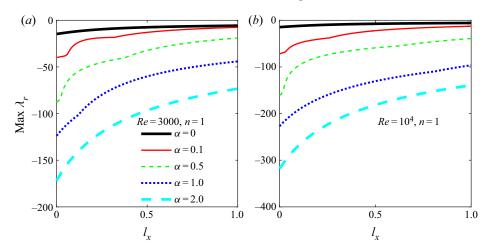


Figure 4. The max λ_r of n=1 modes with $\alpha=0,0.1,0.5,1.0$ and 2.0 as a function of l_x for (a) Re=3000 and (b) $Re=10^4$.

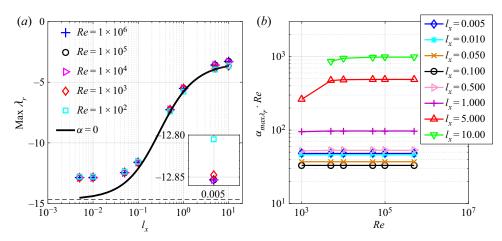


Figure 5. (a) The global maximum of $\max \lambda_r$, i.e. the maximum of $\max \lambda_r$ over α and n, for Re = 100, 1000, 1×10^4 , 1×10^5 and 1×10^6 (symbols). The bold black line shows the maximum $\max \lambda_r$ of the $\alpha = 0$ modes, which is associated with the $(\alpha, n) = (0, 1)$ mode and is independent of Re. The dashed line marks the value for the $(\alpha, n) = (0, 1)$ mode in the no-slip case (Meseguer & Trefethen 2003). The inset shows the close-up at $l_x = 0.005$. (b) The product of Re and $\alpha_{\max \lambda_r}$ (the α at which $\max \lambda_r$ takes the maximum) plotted against Re.

It is interesting to note that our data for high Reynolds numbers all collapse over the whole l_x range investigated, see the cases with Re above 1×10^4 in figure 5, suggesting that the maximum eigenvalue of the system is independent of Re. At lower Reynolds numbers, e.g. Re = 100 and 10^3 in the figure, there is almost a collapse for small $l_x (\leq 0.1)$ but a small deviation from the high Reynolds number cases can be seen, see the inset that shows the close-up at $l_x = 0.005$. As l_x increases further, the maximum eigenvalue for Re = 100 and 10^3 approaches that of $\alpha = 0$ modes, which is strictly Re-independent (see the proof in § 5.1). In addition, the figure also shows that the global maximum of $\max \lambda_r$ is slightly larger than the maximum of the $\alpha = 0$ modes over the whole l_x range and the difference is most significant at small l_x . We did not explore further larger l_x considering that the range we investigated is already much larger than the slip length that can be

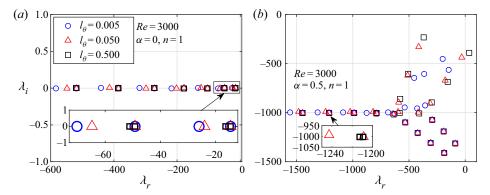


Figure 6. Spectrum of the flow at Re = 3000 with $l_{\theta} = 0.005$ (circles), 0.05 (triangles) and 0.5 (squares). (a) The mode $(\alpha, n) = (0, 1)$. (b) The mode $(\alpha, n) = (0.5, 1)$.

encountered in applications ($\lesssim 0.1$ in set-ups with characteristic length of one millimetre or larger, because so far the maximum slip length achieved in experiments is $O(100)~\mu$ m, see Lee, Choi & Jim (2008), Voronov et al. (2008) and Lee & Jim (2009)). Nevertheless, the S-shaped trend as l_x increases suggests that the flow remains stable no matter how large the slip length is. In fact, as $l_x \to \infty$, the full-slip boundary condition is recovered, and the velocity profile of the base flow will be completely flat and no mean shear exists, in which case linear stability can be expected for any perturbations. Figure 5(b) shows the product of Re and the α at which max λ_r maximizes globally, denoted as $\alpha_{max\lambda_r}$. Interestingly, it seems that this product is a constant when l_x is small ($\lesssim 0.1$) for all the Re investigated, and approaches a constant as Re is sufficiently high ($\gtrsim 10^4$) if $l_x \gtrsim 0.1$. This indicates that $\alpha_{max\lambda_r}$ scales as Re^{-1} for either not very large l_x or in high-Reynolds-number regime. It should be noted that we observed a non-monotonic dependence of $\alpha_{max\lambda_r} \cdot Re$ on the slip length, which minimizes at around $l_x = 0.1$.

That the global $\max \lambda_r$ is Re-independent, as our results suggest, indicates that the slowest exponential decay rate (referred to as decay rate for simplicity hereafter) of perturbations scales as Re^{-1} given that the scaled time $\tau = t/Re$ is used in our formulation, see (2.2). The same scaling was observed by the calculation of Meseguer & Trefethen (2003) for the $(\alpha, n) = (0, 1)$ mode of the classic pipe flow. Therefore, our results suggest that, as $Re \to \infty$, the decay rate of perturbations asymptotically approaches zero and remains negative, i.e. the flow is linearly stable at arbitrary Reynolds number. The Re^{-1} -scaling of the slowest decay rate can be rigorously proved for the $\alpha = 0$ modes, see § 5.1.

In a word, in the pure streamwise slip case, we did not observe any linear instability in the large ranges of l_x and Re that we considered, and based on the data shown in figure 5, we propose that streamwise slip destabilizes the flow but does not cause linear instability, regardless of the slip length and Reynolds number. A similar destabilizing effect was reported by Průša (2009) for the isotropic slip case.

4. Azimuthal slip

We consider the case of $l_{\theta} \neq 0$ and $l_{x} = 0$ as the limiting case of azimuthal slip being significant and streamwise slip being negligible.

The effect of azimuthal slip on the spectrum is investigated for Re = 3000 and is shown in figure 6 for the modes $(\alpha, n) = (0, 1)$ and (0.5, 1). Similar to the streamwise slip case,

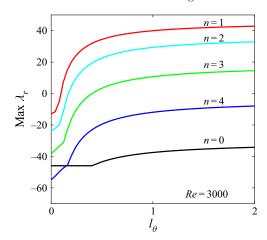


Figure 7. The max λ_r maximized over α , still denoted as max λ_r , as a function of l_{θ} . Modes with n = 0, 1, 2, 3 and 4 are shown for Re = 3000.

the eigenvalues of the $\alpha = 0$ mode also fall on the $\lambda_i = 0$ line and in the left half-plane, see figure 6(a). This suggests that the eigenvalues of streamwise-independent modes are all real and negative. We provide a rigorous proof of this observation in § 5.1. As l_{θ} increases, similar to the streamwise slip case, we also observed two groups of eigenvalues. One group remains constant as the azimuthal slip length changes and the other shifts to the right, see the inset in figure 6(a). As we theoretically show in §§ 5.2 and 5.3, the former group is associated with the disturbances with $\Phi \equiv 0$ and is independent of l_{θ} , and the latter group is associated with the disturbances with $\Phi \neq 0$. The rightmost eigenvalue belongs to the former group for $l_{\theta} < 1$ and can only be overtaken by the latter group if $l_{\theta} > 1$ (the two groups precisely overlap when $l_{\theta} = 1$), i.e. the rightmost eigenvalue can only increase with l_{θ} if $l_{\theta} > 1$. For the $\alpha = 0.5$ and n = 1 mode, the mean mode branch overall does not show either a vertical or horizontal shift, but adjacent eigenvalues are brought closer by the increasing slip length, and for $l_{\theta} = 0.5$ there is almost an eigenvalue degeneracy (see the inset in figure 6b). The centre mode branch is nearly unchanged as l_{θ} increases. However, the wall mode branch is significantly affected. As l_{θ} increases, the wall mode overtakes the centre mode and becomes the least stable perturbation, and as l_{θ} is sufficiently large, the rightmost eigenvalue appears in the right half-plane, indicating the onset of a linear instability. In contrast to the streamwise slip case, the wave speed of the mean modes does not change because the speed follows $\int_0^1 U_x(r) dr$, as aforementioned, and the base flow U(r) is not affected by the azimuthal slip. The speed of the centre modes is also not affected, whereas the wave speed of the wall modes is considerably decreased by the slip. This is reasonable because the slip boundary condition should mostly affect the flow close to the wall and should not affect significantly the flow far from the wall.

Figure 7 shows $\max \lambda_r$ maximized over α (over [0, 2] in practice), still denoted as $\max \lambda_r$, as a function of l_θ for n=0,1,2,3 and 4 at Re=3000. Overall, $\max \lambda_r$ increases monotonically as l_θ increases, while the n=0 case seems to remain constant until it starts to increase at around $l_\theta=0.4$. In the small l_θ regime, all modes are linearly stable. As l_θ is increased to around 0.1, $\max \lambda_r$ of the n=1 mode becomes positive, indicating a linear instability. As l_θ increases further, n=2 and 3 also become unstable. In the whole range of l_θ investigated, n=1 is the least stable/most unstable one, which is also the case for other Reynolds numbers we investigated. Therefore, in the following, we mainly discuss the n=1 modes.

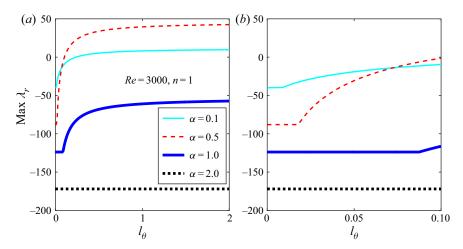


Figure 8. (a) The max λ_r of modes $\alpha = 0.1, 0.5, 1.0$ and 2.0 for Re = 3000 and n = 1 as a function of l_{θ} . (b) The details in the small l_{θ} regime.

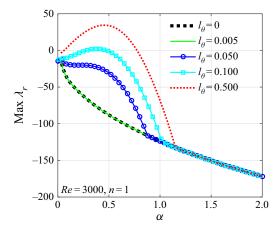


Figure 9. The max λ_r of the n=1 modes as a function of α for Re=3000. The data for $l_{\theta}=0$, 0.005, 0.05, 0.1 and 0.5 are plotted. Note that the curves for $l_{\theta}=0$ (the black bold dotted line) and $l_{\theta}=0.005$ (the green thin solid line) coincide.

Figure 8 shows $\max \lambda_r$ of modes $\alpha=0.1,\ 0.5,\ 1.0$ and 2.0 for Re=3000 and n=1 as a function of l_θ . The results show that when l_θ is small, overall $\max \lambda_r$ decreases as α increases. As l_θ is increased, some moderate α turns to be the least stable/most unstable mode, see the crossover of $\alpha=0.1$ (cyan thin line) and 0.5 (red dashed line) cases in the figure. Figure 8(b) shows the small l_θ range, in which it appears that $\max \lambda_r$ first remains nearly unchanged and then starts to increase, and the trend shows that the larger α , the later $\max \lambda_r$ starts to increase as l_θ is increased. The same behaviour is also observed for $\alpha=0$ modes and we will show a rigorous proof of this behaviour in § 5.3. Interestingly, the case of $\alpha=2$ seems to remain unchanged up to $l_\theta=2.0$.

The dependence of $\max \lambda_r$ on α is more comprehensively shown in figure 9. The smallest $l_\theta = 0.005$ shows a monotonic decrease with increasing α , which completely collapses onto the curve for $l_\theta = 0$, i.e. the classic no-slip case. However, as l_θ increases, $\max \lambda_r$ significantly increases in the region of $0 < \alpha \lesssim 1$ such that a bump appears in

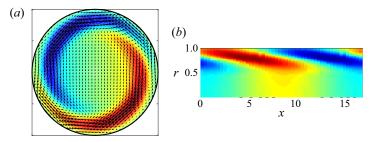


Figure 10. Visualization of the most unstable mode $(\alpha, n) = (0.383, 1)$ for Re = 3000 with $l_{\theta} = 0.1$: (a) the r- θ cross-section and (b) the x-r cross-section. In both panels, the streamwise velocity is plotted as the colourmap with red colour representing positive and blue representing negative values with respect to the base flow. In panel (a), the in-plane velocity field is plotted as arrows.

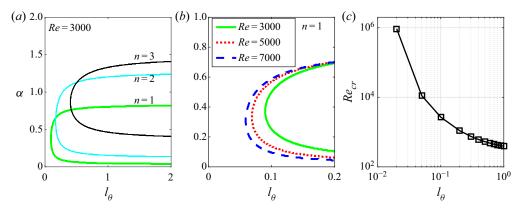


Figure 11. (a) The neutral stability curves for Re = 3000 and n = 1, 2 and 3 in the $l_{\theta} - \alpha$ plane. (a) The neutral stability curves for n = 1 and Re = 3000, 5000 and 7000. (b) The critical Reynolds number Re_{cr} as a function of l_{θ} .

the curves, see those for $l_{\theta}=0.05$, 0.1 and 0.5. At a certain point, the bump reaches $\max \lambda_r=0$ and the flow starts to become linearly unstable if l_{θ} increases further, see the cases of $l_{\theta}=0.1$ and 0.5. As observed in figure 8 for the $\alpha=2$ case, the results suggest that $\max \lambda_r$ of sufficiently large α seems unaffected by azimuthal slip in the l_{θ} range investigated, see the collapse of all curves above $\alpha\simeq 1.2$ in figure 9. It should be noted that as l_{θ} becomes larger, the bump widens up, i.e. $\max \lambda_r$ is affected by the slip in a wider range of α .

Figure 10 shows the velocity field of the most unstable perturbation of mode $(\alpha, n) = (0.383, 1)$ for Re = 3000 with $l_{\theta} = 0.1$. Figure 10(a) shows the in-plane velocity field in the $r - \theta$ pipe cross-section and figure 10(b) shows the pattern of the streamwise velocity in the x - r cross-section. The patterns shown suggest that the flow manifests with a pair of helical waves. The flow structures are mostly located near the wall $(r \gtrsim 0.5)$, indicating that the most unstable mode is a wall mode, which can also been seen in figure 6(b).

Obviously, azimuthal slip can cause linear instability given sufficiently large slip length. We can search in the l_{θ} - α plane to obtain the neutral stability curve for given Re and n. Figure 11(a) shows the neutral stability curves for Re = 3000 and n = 1, 2 and 3 (n = 4 and higher are all stable, see figure 7). It can be seen that, as n increases, the unstable region shifts to the right and upward. However, as the results in figure 7 show, n = 1 is the most unstable based on the eigenvalue maximized over α and, therefore, we only

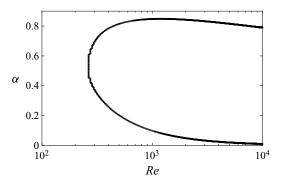


Figure 12. The neutral stability curve in the $Re - \alpha$ plane for $l_{\theta} = \infty$ and n = 1.

investigate the n=1 case in the following. Figure 11(b) shows the neutral stability curves for n=1 and Re=3000, 5000 and 7000. As Re increases, the neutral stability curve moves towards the smaller l_{θ} region, indicating that, for a given l_{θ} , the flow becomes more unstable as Re increases, as expected. The data show that the wavelength of the unstable modes is comparable or significantly larger than the pipe diameter, whereas very long waves ($\alpha \to 0$) and short waves ($\alpha \gg 1.0$) are generally stable. That the flow is always stable to perturbations with $\alpha=0$, regardless of the value of l_{θ} and Re, can be rigorously proved (see § 5.1).

Further, for each l_{θ} , a critical Reynolds number Re_{cr} can be determined by searching the first appearance of a positive $\max \lambda_r$ in the l_{θ} - α plane by varying Re. Figure 11(c) shows Re_{cr} as a function of l_{θ} . As shown, Re_{cr} is a few hundred if l_{θ} is large ($l_{\theta} \gtrsim 0.3$), but the trend suggests that it does not reduce to zero if $l_{\theta} \to \infty$. Since the classic pipe flow is linearly stable for arbitrary Reynolds number, there is an explosive increase in Re_{cr} as l_{θ} decreases, which can be expected because the classic pipe flow will be recovered if $l_{\theta} \to 0$. We also explored the limit of $l_{\theta} \to \infty$, in which case the boundary condition for the azimuthal velocity becomes the full slip condition of

$$\frac{\partial u_{\theta}}{\partial r} = 0. {(4.1)}$$

The neutral stability curve for n=1 in the $Re-\alpha$ plane is shown in figure 12, which shows that the unstable modes are still long waves with α approximately between 0 and 0.8. The critical Reynolds number (the nose of the curve) appears approximately at $Re_{cr}=260$.

5. Eigenvalues and eigenvectors of streamwise independent modes

We can rigorously prove the linear stability of the base flow to perturbations with $\alpha=0$. In the following, we do not consider the $(\alpha,n)=(0,0)$ mode, which should be strictly stable as it is purely dissipative and there can be no energy production mechanism associated with it. In fact, the stability of the classic pipe flow to streamwise independent perturbations has already been proved by Joseph & Hung (1971) using an energy analysis. Nevertheless, here we also account for the effect of the velocity slip and perform analytical studies on the eigenvalues and eigenvectors of $\alpha=0$ modes.

5.1. Proof of linear stability to $\alpha = 0$ modes

For $\alpha = 0$, the eigenvalue λ of the operator $-M^{-1}L$ satisfies

$$\Gamma(n^2\Gamma)\Phi + \lambda\Gamma\Phi = 0 \tag{5.1}$$

$$\frac{2i}{1+4l_x}\Phi + \phi\Omega + \lambda n^2\Omega = 0, (5.2)$$

where Φ and Ω compose the eigenvector q associated with λ (see the definition of q in (2.5)). The boundary conditions (2.12) and (2.13) reduce to

$$l_x \Omega' + \Omega = 0 (5.3)$$

and

$$l_{\theta}\Phi'' + \Phi' = 0. \tag{5.4}$$

It can be seen that for $\alpha=0$ modes, Ω and Φ are decoupled in the boundary conditions (5.3) and (5.4).

We define a space $\Theta = \{f | f \in C^2[0, 1], f(0) = f(1) = 0\}$ and an inner product associated with this space

$$(f_1, f_2) = \int_0^1 r f_1 \bar{f}_2 \, \mathrm{d}r, \tag{5.5}$$

where the overbar represents the complex conjugate. Then the operator Γ has the following two properties:

(a)
$$(\Gamma f_1, f_2) = (f_1, \Gamma f_2), \quad \forall f_1, f_2 \in \Theta,$$
 (5.6)

Proof.

$$(\Gamma f_{1}, f_{2}) = \int_{0}^{1} r \left(\frac{f_{1}}{r^{2}} - \frac{1}{r} \frac{d}{dr} \left(\frac{r}{n^{2}} \frac{df_{1}}{dr} \right) \right) \bar{f}_{2} dr = \int_{0}^{1} \frac{f_{1} \bar{f}_{2}}{r} dr - \int_{0}^{1} \bar{f}_{2} d \left(\frac{r}{n^{2}} \frac{df_{1}}{dr} \right)$$

$$= \int_{0}^{1} \frac{f_{1} \bar{f}_{2}}{r} dr - \bar{f}_{2} \left(\frac{r}{n^{2}} \frac{df_{1}}{dr} \right) \Big|_{0}^{1} + \int_{0}^{1} \frac{r}{n^{2}} \frac{df_{1}}{dr} d\bar{f}_{2} = \int_{0}^{1} \frac{f_{1} \bar{f}_{2}}{r} dr + \int_{0}^{1} \frac{r}{n^{2}} \frac{d\bar{f}_{2}}{dr} df_{1}$$

$$(5.7)$$

and similarly, using integration by parts, it can be derived that

$$(f_{1}, \Gamma f_{2}) = \int_{0}^{1} r \left(\frac{\bar{f}_{2}}{r^{2}} - \frac{1}{r} \frac{d}{dr} \left(\frac{r}{n^{2}} \frac{d\bar{f}_{2}}{dr} \right) \right) f_{1} dr = \int_{0}^{1} \frac{f_{1}\bar{f}_{2}}{r} dr + \int_{0}^{1} \frac{r}{n^{2}} \frac{d\bar{f}_{2}}{dr} df_{1} = (\Gamma f_{1}, f_{2});$$

$$(5.8)$$

(b)
$$(\Gamma f, f) \geqslant 0, \quad \forall f \in \Theta, \tag{5.9}$$

Proof. taking $f = f_1 = f_2$ in proof 5.1,

$$(\Gamma f, f) = \int_0^1 \frac{f\bar{f}}{r} dr + \int_0^1 \frac{r}{n^2} \frac{d\bar{f}}{dr} df = \int_0^1 \frac{f\bar{f}}{r} dr + \int_0^1 \frac{r}{n^2} \frac{d\bar{f}}{dr} \frac{df}{dr} dr \geqslant 0.$$
 (5.10)

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Note that property (5.9) still holds for those f with $f(1) \neq 0$ but satisfies f(1) + bf'(1) = 0, where b > 0 is a constant, because

$$(\Gamma f, f) = \int_{0}^{1} \frac{f\bar{f}}{r} dr - f \left(\frac{r}{n^{2}} \frac{d\bar{f}}{dr} \right) \Big|_{0}^{1} + \int_{0}^{1} \frac{r}{n^{2}} \frac{d\bar{f}}{dr} df$$

$$= \int_{0}^{1} \frac{f\bar{f}}{r} dr + \int_{0}^{1} \frac{r}{n^{2}} \frac{d\bar{f}}{dr} dr + \frac{1}{bn^{2}} f(1)\bar{f}(1) \geqslant 0.$$
 (5.11)

First, for the case of $\Phi \equiv 0$ (i.e. the wall-normal velocity component $u_r \equiv 0$) and $\Omega \not\equiv 0$, (5.2) becomes

$$\phi \Omega + \lambda n^2 \Omega = 0 \tag{5.12}$$

and the operators ϕ and Γ are related as

$$\phi = \frac{n^4}{r^2} - \frac{1}{r} \frac{\mathrm{d}}{\mathrm{d}r} \left(n^2 r \frac{\mathrm{d}}{\mathrm{d}r} \right) = n^4 \Gamma. \tag{5.13}$$

Therefore, (5.12) becomes

$$n^4 \Gamma \Omega + \lambda n^2 \Omega = 0. \tag{5.14}$$

Taking the inner product of (5.14) with Ω , we have

$$(n^4 \Gamma \Omega, \Omega) + (\lambda n^2 \Omega, \Omega) = 0. \tag{5.15}$$

According to property (5.9), $(n^4 \Gamma \Omega, \Omega) \ge 0$ given $\Omega(1) = 0$ (without streamwise slip) or $\Omega(1) + l_x \Omega'(1) = 0$ (with streamwise slip), which leads to $\lambda < 0$, i.e. the eigenvalue is real and negative.

Second, we discuss the $\Phi \neq 0$ case, i.e. the wall-normal velocity component $u_r \neq 0$. From (5.1), by denoting $g = n^2 \Gamma \Phi + \lambda \Phi$, we have $\Gamma g = 0$, i.e.

$$n^2 g = r(rg')', (5.16)$$

from which it can be obtained that

$$r^n g = Cr^{2n} + C_1, (5.17)$$

where C and C_1 are constants. Note that for n=2, $r^2g=n^2r^2\Gamma\Phi+\lambda r^2\Phi=n^2\Phi-nr(r\Phi')'$ has to vanish at r=0, because Φ vanishes, and Φ' and Φ'' are finite at r=0. The same applies to n>2. If n=1, $rg=\Phi/r-(r\Phi')'+\lambda r\Phi=\Phi/r-\Phi'-r\Phi''+\lambda r\Phi$, which also vanishes when $r\to 0$ (using L'Hôpital's rule). Therefore, $C_1\equiv 0$ and $r^ng=Cr^{2n}$, i.e.

$$n^2 \Gamma \Phi + \lambda \Phi = Cr^n. \tag{5.18}$$

5.1.1. The case without azimuthal slip, i.e. $l_{\theta} = 0$

In case of $l_{\theta}=0$, the boundary condition (2.13) or (5.4) becomes $\Phi'=0$. Taking the inner product (5.5) of (5.18) and $\Gamma\Phi$, we have

$$n^{2}(\Gamma\Phi, \Gamma\Phi) + \lambda(\Phi, \Gamma\Phi) = C(r^{n}, \Gamma\Phi) = C(\Gamma r^{n}, \Phi) = C(0, \Phi) = 0.$$
 (5.19)

The second equality in (5.19) holds in spite of the fact that $r^n \notin \Theta$ and thus, property (5.6) cannot be applied directly. Nevertheless, as $\Phi = \Phi' = 0$ at r = 1 in the case of $l_{\theta} = 0$, property (5.6) still holds (this can be seen by taking Φ as f_2 and r^n as f_1 in proof 5.1). What follows is that the eigenvalue λ is real and $\lambda < 0$ because $(\Gamma \Phi, \Gamma \Phi) > 0$ and $(\Phi, \Gamma \Phi) > 0$.

5.1.2. The case with azimuthal slip, i.e. $l_{\theta} \neq 0$

In the case of $l_{\theta} \neq 0$, (5.19) does not hold, except for C = 0, because $\Phi' = 0$ at r = 1 does not necessarily hold and therefore the second equality in (5.19) does not hold either. For $C \neq 0$, consider the special case of C = 1 (if $C \neq 1$, a rescaling of $\tilde{\Phi} = \Phi/C$ can easily convert to this special case), (5.18) can be written as

$$(n^2 + \lambda r^2)\Phi - r(r\Phi')' = (n^2 + \lambda r^2)\Phi - r(\Phi' + r\Phi'') = r^{n+2}.$$
 (5.20)

As $r \rightarrow 1$, (5.20) turns to

$$-(\Phi' + \Phi'') = 1. (5.21)$$

Further, the azimuthal slip requires

$$\Phi'(1) + l_{\theta}\Phi''(1) = 0, \quad l_{\theta} \in (0, +\infty).$$
 (5.22)

It follows that, for $l_{\theta}=1$, C has to be zero, otherwise (5.21) and (5.22) would conflict with each other. That C=0 leads to $\lambda<0$, see (5.19). For $l_{\theta}\neq 1$, one can solve for $\Phi'(1)$ from (5.21) and (5.22) as

$$\Phi'(1) = \frac{l_{\theta}}{1 - l_{\theta}},\tag{5.23}$$

which indicates that $\Phi'(1)$ is real and $\Phi'(1) \in (-\infty, -1) \cup (0, +\infty)$.

It can be verified that (5.20) has a special solution

$$\Phi = \frac{r^n}{\lambda} \tag{5.24}$$

and its corresponding homogeneous differential equation is

$$r^{2}\Phi'' + r\Phi' - (n^{2} + \lambda r^{2})\Phi = 0.$$
 (5.25)

From the theory of ordinary differential equations, this equation has two linearly independent solutions in (0, 1]. One of the two solutions can be represented as a generalized power series

$$\Phi_1 = \sum_{m=0}^{\infty} B_m r^{m+\rho} \quad (B_0 \neq 0), \tag{5.26}$$

in which it can be obtained that $\rho = n$, $B_{2k+1} = 0$ and $B_{2k} = (\lambda/4)^k (B_0/k!(n+k)!)$ using the standard method of undetermined coefficients. Denoting $a_n = B_0$, (5.26) can be written as

$$\Phi_1 = a_n r^n \sum_{k=0}^{\infty} \left(\frac{\lambda}{4}\right)^k \frac{r^{2k}}{k!(n+k)!}.$$
 (5.27)

The other solution of (5.25) has the following form:

$$\Phi_2 = \Phi_1 \int_1^r \frac{1}{\Phi_1^2(s)} \exp\left(-\int_1^s \frac{1}{t} dt\right) ds = \Phi_1 \int_1^r \frac{1}{s\Phi_1^2(s)} ds.$$
 (5.28)

However, by L'Hôpital's rule,

$$\lim_{r \to 0} \Phi_2(r) = \lim_{r \to 0} \frac{\int_1^r \frac{1}{s\Phi_1^2(s)} \, \mathrm{d}s}{\frac{1}{\Phi_1(r)}} = \lim_{r \to 0} \frac{\frac{1}{r\Phi_1^2(r)}}{\frac{\Phi_1'(r)}{\Phi_2^2(r)}} = \infty, \tag{5.29}$$

which is unphysical, and therefore Φ_2 should not appear in the general solution of (5.20), i.e. the general solution of (5.20) can be solved as

$$\Phi = \frac{1}{\lambda}r^n + a_n r^n \sum_{k=0}^{\infty} \left(\frac{\lambda}{4}\right)^k \frac{r^{2k}}{k!(n+k)!}.$$
 (5.30)

For simplicity, denoting $\mu = \lambda/4$ and using the boundary condition $\Phi(1) = 0$, one can solve for a_n as

$$a_n = -\frac{1}{4\mu} \left(\sum_{k=0}^{\infty} \frac{\mu^k}{k!(n+k)!} \right)^{-1},\tag{5.31}$$

consequently,

$$\Phi'(1) = \frac{n}{4\mu} + a_n \sum_{k=0}^{\infty} \frac{\mu^k (n+2k)}{k!(n+k)!} = -\frac{1}{4\mu} \left(\sum_{k=0}^{\infty} \frac{\mu^k}{k!(n+k)!} \right)^{-1} \sum_{k=0}^{\infty} \frac{2k\mu^k}{k!(n+k)!}, \quad (5.32)$$

i.e. μ satisfies

$$\sum_{k=1}^{\infty} \frac{k\mu^{k-1}}{k!(n+k)!} + 2\Phi'(1) \sum_{k=0}^{\infty} \frac{\mu^k}{k!(n+k)!} = 0.$$
 (5.33)

In the following, we prove that μ has to be real and $\mu < 0$ given (5.33). For simplicity, let $s = \Phi'(1)$ and define f(z) as

$$f(z) = \sum_{k=0}^{\infty} \frac{z^k}{k!(n+k)!},$$
(5.34)

where z is complex. Then, (5.33) states that μ is a root of the equation f'(z) + 2sf(z) = 0. Note that

$$\left(z^{n+1}f'(z)\right)' = \left(\sum_{k=1}^{\infty} \frac{kz^{n+k}}{k!(n+k)!}\right)' = z^n \sum_{k=1}^{\infty} \frac{z^{k-1}}{(k-1)!(n+k-1)!}$$
$$= z^n \sum_{k=0}^{\infty} \frac{z^k}{k!(n+k)!} = z^n f(z). \tag{5.35}$$

Then, defining $f_{\mu}(z) = f(\mu z)$, it can be verified that

$$(z^{n+1}f'_{\mu}(z))' = \mu z^n f_{\mu}(z), \tag{5.36}$$

in which the prime denotes the derivative with respect to z. Further, note that $\bar{\mu}$ is also a root of (5.33), because the coefficients are all real. That is to say,

$$(z^{n+1}f'_{\bar{\mu}}(z))' = \bar{\mu}z^n f_{\bar{\mu}}(z), \tag{5.37}$$

where $f_{\bar{\mu}} = f(\bar{\mu}z)$. Then, the difference between (5.36) multiplied by $f_{\bar{\mu}}(z)$ and (5.37) multiplied by $f_{\mu}(z)$, integrated along the real axis from 0 to 1, gives that

$$\int_{0}^{1} (z^{n+1} f'_{\mu}(z))' f_{\bar{\mu}}(z) dz - \int_{0}^{1} (z^{n+1} f'_{\bar{\mu}}(z))' f'_{\mu}(z) dz$$

$$= z^{n+1} f'_{\mu}(z) f_{\bar{\mu}}(z) \Big|_{0}^{1} - \int_{0}^{1} z^{n+1} f'_{\mu}(z) df_{\bar{\mu}}(z) - z^{n+1} f'_{\bar{\mu}}(z) f_{\mu}(z) \Big|_{0}^{1} + \int_{0}^{1} z^{n+1} f'_{\bar{\mu}}(z) df_{\mu}(z)$$

$$= \int_{0}^{1} z^{n+1} |f'_{\mu}(z)|^{2} dz - \int_{0}^{1} z^{n+1} |f'_{\mu}(z)|^{2} dz + f'_{\mu}(1) f_{\bar{\mu}}(1) - f_{\mu}(1) f'_{\bar{\mu}}(1)$$

$$= 2s(\bar{\mu} - \mu) |f_{\mu}(1)|^{2} = (\mu - \bar{\mu}) \int_{0}^{1} z^{n} |f_{\mu}(z)|^{2} dz, \qquad (5.38)$$

where the condition f'(z) + 2sf(z) = 0 is used to derive $f'_{\mu}(z) + 2\mu s f_{\mu}(z) = 0$ and $f'_{\bar{\mu}}(z) + 2\bar{\mu}s f_{\bar{\mu}}(z) = 0$. Then we have

$$(\mu - \bar{\mu}) \left(\int_0^1 z^n |f_{\mu}(z)|^2 dz + 2s |f_{\mu}(1)|^2 \right) = 0.$$
 (5.39)

Similarly, the sum of (5.36) multiplied by $f_{\bar{\mu}}(z)$ and (5.37) multiplied by $f_{\mu}(z)$, integrated along the real axis from 0 to 1, gives that

$$(\mu + \bar{\mu}) \left(\int_0^1 z^n |f_{\mu}(z)|^2 dz + 2s |f_{\mu}(1)|^2 \right) = -2 \int_0^1 z^{n+1} |f'_{\mu}(z)|^2 dz, \tag{5.40}$$

which indicates $\int_0^1 z^n |f_{\mu}(z)|^2 dz + 2s|f_{\mu}(1)|^2 \neq 0$ because the right-hand side is non-zero. Consequently, (5.39) indicates that $\mu - \bar{\mu} = 0$, i.e. μ is real.

Subsequently, we can deduce that $\mu < 0$ if $s \in (0, \infty)$ because the term in the parentheses and the integral on the right-hand side of (5.40) are all positive. In case of $s \in (-\infty, -1)$, if μ were positive, one would obtain

$$-2s\sum_{k=0}^{\infty} \frac{\mu^k}{k!(n+k)!} = \sum_{k=1}^{\infty} \frac{k\mu^{k-1}}{k!(n+k)!} = \sum_{k=0}^{\infty} \frac{\mu^k}{k!(n+k+1)!} \leqslant \sum_{k=0}^{\infty} \frac{\mu^k}{k!(n+k)!}$$
(5.41)

and consequently $-2s \leqslant 1$, which would conflict with $s \in (-\infty, -1)$. Therefore, $\mu < 0$. Finally, we obtain that $\mu < 0$ for $s \in (-\infty, -1) \cup (0, +\infty)$, i.e. for any value of $l_{\theta} \neq 1$. Since we have shown before that $\lambda < 0$ for $l_{\theta} = 1$ and for $l_{\theta} = 0$, now we reach the conclusion that λ is real and $\lambda < 0$ for any $l_{\theta} \in [0, +\infty)$, regardless of l_x , i.e. the flow is rigorously linearly stable to perturbations with $\alpha = 0$ with or without velocity slip.

5.2. Analytical solution of the eigenvalue and eigenvector for $\alpha = 0$ modes We consider the general case with both streamwise and azimuthal slip. For $\Phi \not\equiv 0$, if $C \not\equiv 0$, we obtain from (5.33) that $\mu = \lambda/4$ satisfies

$$(1 - l_{\theta}) \sum_{k=1}^{\infty} \frac{k\mu^{k-1}}{k!(n+k)!} + 2l_{\theta} \sum_{k=0}^{\infty} \frac{\mu^{k}}{k!(n+k)!} = 0,$$
 (5.42)

where (5.23) is used. The Bessel functions of integer order n and n + 1 read

$$J_n(z) = \sum_{k=0}^{\infty} \frac{(-1)^k}{k!(n+k)!} \left(\frac{z}{2}\right)^{2k+n} = \left(\frac{z}{2}\right)^n \sum_{k=0}^{\infty} \frac{1}{k!(n+k)!} \left(-\frac{z^2}{4}\right)^k, \tag{5.43}$$

$$J_{n+1}(z) = \sum_{k=0}^{\infty} \frac{(-1)^k}{k!(n+1+k)!} \left(\frac{z}{2}\right)^{2k+n+1} = \left(\frac{z}{2}\right)^{n+1} \sum_{k=0}^{\infty} \frac{1}{k!(n+k+1)!} \left(-\frac{z^2}{4}\right)^k.$$
(5.44)

Denoting $\mu = -(\eta^2/4)$, i.e. the eigenvalue $\lambda = 4\mu = -\eta^2$, it can be observed that η is a root of the equation

$$(1 - l_{\theta})J_{n+1}(z) + l_{\theta}zJ_n(z) = 0.$$
 (5.45)

Next, we show that $C \neq 0$ if $l_{\theta} \neq 1$. Assuming C = 0 and $l_{\theta} \neq 1$, $\Phi'(1) + \Phi''(1) = 0$ and the boundary condition $\Phi'(1) + l_{\theta}\Phi''(1) = 0$ would give $\Phi'(1) = \Phi''(1) = 0$. Recall that the solution to the homogeneous equation (5.25) is

$$\Phi_1 = a_n r^n \sum_{k=0}^{\infty} \left(\frac{\lambda}{4}\right)^k \frac{r^{2k}}{k!(n+k)!},\tag{5.46}$$

where a_n is a constant. Using the notation of (5.34),

$$f(\mu) = \sum_{k=0}^{\infty} \mu^k \frac{1}{k!(n+k)!} = 0$$
 (5.47)

results from $\Phi_1(1) = 0$ (note that $\Phi = \Phi_1$ if C = 0), which would indicate that the corresponding η satisfies $J_n(z) = 0$. Further, $\Phi'(1) = 0$ gives

$$\sum_{k=0}^{\infty} \mu^k \frac{n+2k}{k!(n+k)!} = 0. {(5.48)}$$

In combination with $f(\mu) = 0$, we would obtain that

$$\sum_{k=1}^{\infty} \mu^k \frac{1}{(k-1)!(n+k)!} = \sum_{k=0}^{\infty} \mu^k \frac{1}{k!(n+k+1)!} = 0,$$
 (5.49)

which means that η would also be a zero of $J_{n+1}(z)$, i.e. η would be a zero of both $J_n(z)$ and $J_{n+1}(z)$. This would conflict with the fact that there exists no common zero of $J_n(z)$ and $J_{n+1}(z)$. Therefore, $C \neq 0$ and η is a root of (5.45) if $l_{\theta} \neq 1$.

We have proved before that C=0 if $l_{\theta}=1$, which gives $\Phi=\Phi_1$. Consequently, $\Phi(1)=\Phi_1(1)=0$ gives $f(\mu)=0$, which means η is a root of $J_n(z)=0$, i.e. η is also a root of (5.45) (note that the first term disappears if $l_{\theta}=1$). Therefore, η is a root of (5.45) for any $l_{\theta} \ge 0$.

For the case of $\Phi \equiv 0$, (5.14) and the corresponding boundary condition $\Omega(1) + l_x \Omega'(1) = 0$ (see (2.13) and (5.3)) imply that the eigenvector Ω has the same form as

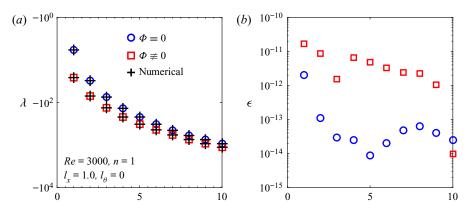


Figure 13. Validation of the analytical eigenvalues against the numerical calculation for the case of Re = 3000, n = 1, $l_x = 1.0$ and $l_\theta = 0$. In panel (a) the first 20 eigenvalues are shown as squares and circles, and the numerical results are shown as crosses. The circles are the first 10 eigenvalues (in descending order) for the cases with $\Phi \equiv 0$ and the squares are the first 10 eigenvalues (in descending order) for the $\Phi \not\equiv 0$ case. The abscissa shows the indices of the eigenvalues. (b) The relative error ϵ between the analytical and numerical ones.

the solution Φ_1 , and we can deduce that $\lambda = -\gamma^2$, in which γ is a root of

$$(1 + nl_x)J_n(z) - l_x z J_{n+1}(z) = 0. (5.50)$$

For an eigenvalue $\lambda = -\eta^2$, the corresponding eigenvector can be solved as

$$\Phi = J_n(\eta r) - J_n(\eta)r^n, \tag{5.51}$$

$$\Omega = b_n J_n(\eta r) + \frac{2i}{(1 + 4l_x)n^2} \left(\frac{r}{2\eta} J_{n+1}(\eta r) - \frac{J_n(\eta)}{\eta^2} r^n \right), \tag{5.52}$$

where b_n is a constant and should be determined by the boundary condition (5.3). For an eigenvalue $\lambda = -\gamma^2$, the eigenvector can be solved as

$$\Phi \equiv 0, \quad \Omega = J_n(\gamma r).$$
 (5.53*a*,*b*)

To sum up, there are always two groups of eigenvalues, corresponding to $\Phi \equiv 0$ (given by (5.50)) and $\Phi \not\equiv 0$ (given by (5.45)), respectively. Particularly, for the no-slip case, (5.45) reduces to $J_{n+1}(z) = 0$ and (5.50) reduces to $J_n(z) = 0$, and it is known that the zeros of $J_n(z)$ and $J_{n+1}(z)$ distribute alternately. Therefore, in the no-slip case, these two groups of eigenvalues distribute alternately. For the streamwise slip case, (5.45) still reduces to $J_{n+1}(z) = 0$, i.e. the $\Phi \not\equiv 0$ eigenvalues do not change with l_x , whereas the $\Phi \equiv 0$ eigenvalues will change with l_x . However, there cannot be common roots between $J_{n+1}(z) = 0$ and (5.50), otherwise there would be common zeros between $J_n(z)$ and $J_{n+1}(z)$, which conflicts with the fact that there are no common zeros between the two. Therefore, as l_x changes, the two groups of eigenvalues distribute in the same alternating pattern as in the no-slip case and there is no overtaking between the two groups, see figures 1(a) and 13(a). However, this behaviour is not guaranteed in the azimuthal slip case as there can be common roots between (5.45) and (5.50) given $l_x = 0$ and $l_\theta = 1.0$, i.e. the roots of $J_n(z) = 0$. Nonetheless, it should be noted that the common roots can only exist at $l_{\theta} = 1.0$. This implies that, when eigenvalues change with l_{θ} , an overtaking between the two groups may occur at precisely $l_{\theta} = 1.0$.

Figure 13 shows the comparison between our analytical solution of the two groups of eigenvalues and numerical calculation for the streamwise slip case of Re = 3000,

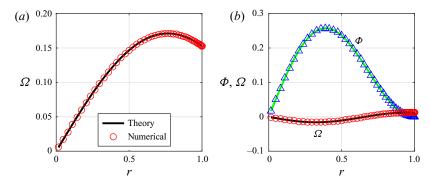


Figure 14. Validation of the analytical eigenvectors against the numerical calculation for the case of Re = 3000, n = 1, $l_x = 1.0$ and $l_\theta = 0$. (a) The Ω component of the eigenvector associated with the leading eigenvalue (the leftmost blue circle in figure 13a). The Φ component is zero and is not shown. (b) The Ω and Φ component of the eigenvector associated with the second largest eigenvalue (the leftmost red square in figure 13a).

n=1, $l_x=1.0$ and $l_\theta=0$. In figure 13(a), blue circles are analytical solutions of the first 10 largest eigenvalues given by (5.50), i.e. the corresponding eigenvectors all have $\Phi \equiv 0$, and red squares represent the first 10 largest eigenvalues given by (5.45), i.e. the corresponding eigenvectors all have $\Phi \equiv 0$. Clearly, the leading eigenvalue is and will always be associated with $\Phi \equiv 0$ disturbances because no overtaking between the two groups of eigenvalues can occur as l_x varies, as we concluded in § 5.2. These analytical solutions agree very well with the numerical calculations (the crosses) with relative errors of $O(10^{-11})$ or lower, see figure 13(b). The eigenvector associated with the leading eigenvalue (the leftmost circle in figure 13a) is plotted in figure 14(a). The black line shows the analytical solution given by (5.53a,b) and the circles show the numerical calculation. The Φ part of the eigenvector is not shown because $\Phi \equiv 0$. Figure 14(b) shows the eigenvector associated with the second largest eigenvalue (the leftmost red square in figure 13a), which has a non-zero Φ part. The figure shows that, for both Φ and Ω , our analytical solutions (lines) agree very well with the numerical calculations (symbols). This comparison validates our theory about the eigenvalue and eigenvector.

The two groups of eigenvalues of the azimuthal slip cases of $l_{\theta} = 0.05$ and 2.0 for Re = 3000, n = 1 and $l_x = 0$ are also shown in figure 15. Again, perfect agreement between the analytical and numerical ones is observed. We can see that, for $l_{\theta} = 0.05$, the $\Phi \neq 0$ group is entirely below the $\Phi \equiv 0$ group, which is independent of l_{θ} , whereas it is entirely above the $\Phi \equiv 0$ group for $l_{\theta} = 2.0$, indicating that an overtaking indeed occurs between the two groups as l_{θ} increases. Therefore, for $l_{\theta} < 1.0$, the leading eigenvalue is associated with $\Phi \equiv 0$ disturbances and does not change with l_{θ} (see also figure 6a), whereas it is associated with $\Phi \not\equiv 0$ disturbances and increases with l_{θ} for $l_{\theta} > 1.0$.

5.3. The dependence of the leading eigenvalue on slip length for $\alpha = 0$ modes Denoting $F(z, l_{\theta}) = (1 - l_{\theta})J_{n+1}(z) + l_{\theta}zJ_n(z)$, it can be obtained that, as $z \to 0$,

$$J_n(z) \sim \frac{z^n}{2^n n!}, \quad F(z, l_\theta) \sim \left(\frac{1 - l_\theta}{2^{n+1}(n+1)!} + \frac{l_\theta}{2^n n!}\right) z^{n+1}.$$
 (5.54a,b)

It can be seen that

$$\frac{1 - l_{\theta}}{2^{n+1}(n+1)!} + \frac{l_{\theta}}{2^n n!} > 0 \tag{5.55}$$

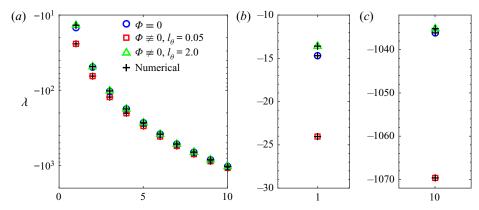


Figure 15. Eigenvalues for Re = 3000, n = 1 and $l_x = 0$ with $l_\theta = 0.05$ and 2.0. The abscissa shows the indices of the eigenvalues in descending order. In panel (a) analytical eigenvalues are shown as circles ($\Phi \equiv 0$, for which eigenvalues are independent of l_θ), squares ($\Phi \neq 0$ and $l_\theta = 0.05$) and triangles ($\Phi \neq 0$ and $l_\theta = 2.0$), and the numerical calculations are shown as crosses. Panels (b,c) show the close-up of the two ends of the spectrum shown in panel (a).

for $l_{\theta} \ge 0$, therefore, $F(z, l_{\theta})$ is positive for sufficiently small z. Let z_1 be the minimum root of $F(z, l_{\theta 1}) = 0$ and z_2 be the minimum root of $F(z, l_{\theta 2}) = 0$. If $l_{\theta 1} < l_{\theta 2}$, it can be derived that

$$F(z_1, l_{\theta 2}) = (1 - l_{\theta 2})J_{n+1}(z_1) + l_{\theta 2}z_1J_n(z_1) = (1 - l_{\theta 2})J_{n+1}(z_1) - \frac{1 - l_{\theta 1}}{l_{\theta 1}}l_{\theta 2}J_{n+1}(z_1)$$

$$= \frac{l_{\theta 1} - l_{\theta 2}}{l_{\theta 1}}J_{n+1}(z_1) < 0.$$
(5.56)

In (5.56), $J_{n+1}(z) > 0$ follows from that, at the minimum positive zero of $J_{n+1}(z)$, denoted as z_0 , we have $F(z_0, l_{\theta 1}) < 0$ because $J_n(z_0) < 0$. We showed before that $F(z, l_{\theta 1}) > 0$ at sufficiently small z, therefore, the minimum positive zero of $F(z, l_{\theta 1}), z_1$, should be smaller than z_0 given that $F(z, l_{\theta 1})$ is continuous with respect to z, i.e. $z_1 < z_0$, and therefore $J_{n+1}(z_1) > 0$. Consequently, given $F(z_1, l_{\theta 2}) < 0$, there must be a zero in $(0, z_1)$, i.e. $z_2 < 0$ z_1 because the function $F(z, l_{\theta 2})$ is continuous with respect to z. This states that, for the case of $\Phi \not\equiv 0$, the maximum eigenvalue λ , denoted as λ_1 in the following, increases as l_{θ} increases and is independent of l_{x} . Similarly, one can deduce that the minimum root of (5.50) decreases as l_x increases, consequently, the maximum eigenvalue for the $\Phi \equiv 0$ case, denoted as λ_2 , increases as l_x increases and is independent of l_θ . For the special case of $l_x = 0$, (5.50) becomes $J_n(z) = 0$ and for the case of $l_\theta = 1$, (5.45) turns into $zJ_n(z)=0$. Clearly, these two cases share the non-zero roots, i.e. $\lambda_1=\lambda_2$. Therefore, the minimum root of (5.50) is always greater than that of (5.45), i.e. $\lambda_1 > \lambda_2$, when $l_{\theta} < 1$. This explains why, for a given $l_{\theta} \leq 1$, max λ increases monotonically as l_x increases from zero, whereas for a given $l_{\theta} > 1$, max λ first remains constant and only starts to increase until l_x is increased above a threshold, see figure 16(a). If only azimuthal slip is present, i.e. $l_x = 0$, max λ first remains constant and only starts to increase precisely at $l_\theta = 1$, see figure 16(b). The data shown in the inset of figure 6(a) also support this conclusion, see that max λ for $l_{\theta} = 0.005$, 0.05 and 0.5 are identical. It can also be inferred that, given a fixed $l_x > 0$ and that λ_2 increases with l_x , max λ can only start to increase as l_θ increases at some $l_{\theta} > 1$.

In summary, the maximum eigenvalue of $\alpha = 0$ modes is an increasing function of l_{θ} or l_{x} (may not be strictly increasing, depending on the slip length setting, as figure 16 shows)

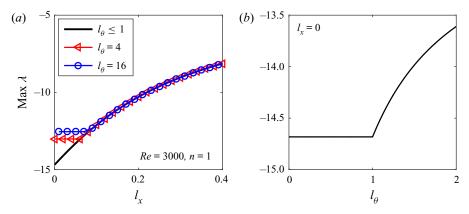


Figure 16. The dependence of $\max \lambda$ on slip length for Re = 3000 and n = 1. (a) Both streamwise and azimuthal slip are present. The black line shows the dependence on l_x given $l_\theta \le 1$. Symbol lines show two cases for $l_\theta > 1$. (b) The dependence on l_θ in case of $l_x = 0$.

and is independent of the Reynolds number, which is obvious as Re does not appear in (5.45) and (5.50). Nevertheless, the eigenvalues remain negative.

6. Non-modal stability

It has been known that in many shear flows (e.g. pipe, channel and plane-Couette flows), small disturbances can be transiently amplified due to the non-normality of the linearized equations, despite their asymptotic linear stability (Schmid & Henningson 1994; Meseguer & Trefethen 2003; Schmid 2007). This transient amplification is believed to play an important role in the subcritical transition in shear flows. Here, we also investigated the effects of the anisotropic slip on the non-modal stability of the flow. The same method for calculating the transient growth described by Schmid & Henningson (1994) is adopted. The transient growth at time t for a mode (α, n) is defined as

$$G(t; \alpha, n) = \max_{E(0) \neq 0} \frac{E(t)}{E(0)},$$
(6.1)

where $E(t) = \int_V u(t)^2 dV$ is the kinetic energy of the perturbation u integrated in the whole flow domain at time t. For linearly stable flow, G will reach the maximum, G_{max} , at a certain time and monotonically decay at larger times. For linearly unstable flow, G can be either non-monotonic or monotonic at early stages, depending on the competition between the modal and non-modal growth, and will be dominated by the exponential growth of the most unstable disturbance at large times.

Figure 17 shows G_{max} at Re=3000 in the l_x - α plane ($l_x=0$) for n=1 and n=2 (low azimuthal wavenumbers are generally most amplified by non-normality). From the contour lines we can see that streamwise slip reduces G_{max} and the decrease is monotonic as l_x increases. Intuitively, streamwise slip reduces the background shear such that the lift-up mechanism (Brandt 2014) is subdued. Therefore, the transient growth should be reduced as our results show. The most amplified mode is still the (α , n) = (0, 1) mode (streamwise rolls) as in the no-slip case (Schmid & Henningson 1994; Meseguer & Trefethen 2003).

Figure 18 shows G_{max} for the azimuthal slip case at Re = 3000 in the $l_{\theta} - \alpha$ plane. Azimuthal wavenumbers n = 1 and n = 2 are considered. From the orientation of the contour lines we can see that azimuthal slip increases G_{max} . Presumably, azimuthal

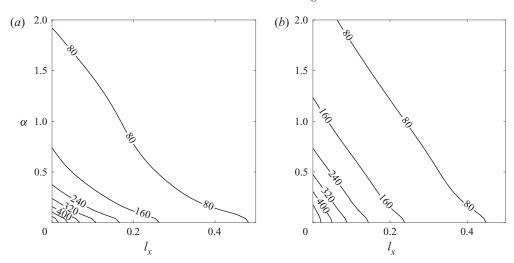


Figure 17. The maximum transient growth, G_{max} , at Re = 3000 plotted in the l_x - α plane for n = 1 (a) and n = 2 (b). Azimuthal slip length l_θ =0. The contour level step is 80 in both panels.

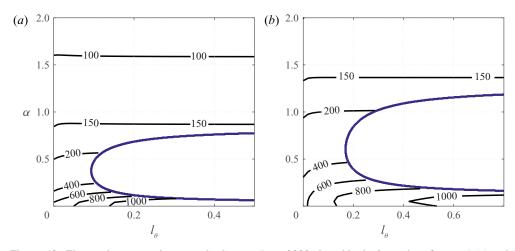


Figure 18. The maximum transient growth, G_{max} , at Re = 3000 plotted in the l_{θ} - α plane for n = 1 (a) and n = 2 (b). Streamwise slip length l_x =0. The bold lines enclose the linearly unstable regions.

slip can enhance streamwise vortices because it reduces wall friction and allows finite azimuthal velocity at the wall, and therefore, the lift-up mechanism can be enhanced exhibiting increased transient growth as our results show. For n = 1, in the l_{θ} range investigated, the most amplified mode is still the $(\alpha, n) = (0, 1)$ mode as in the no-slip case. However, for n = 2, as l_{θ} increases, the most amplified mode is no longer the streamwise independent one but one with a small finite streamwise wavenumber (long wavelength), see figure 18(b). For example, the most amplified mode is approximately $\alpha = 0.05$ for l_{θ} above approximately 0.1. Similar behaviour has also been reported for channel flow (Chai & Song 2019). The two bold lines in the figure enclose the linearly unstable regions in which G_{max} is theoretically infinite, therefore, the linearly unstable region is left blank. Unlike the streamwise slip case where the slip length significantly reduces the transient growth throughout the α and l_x ranges investigated (see figure 17),

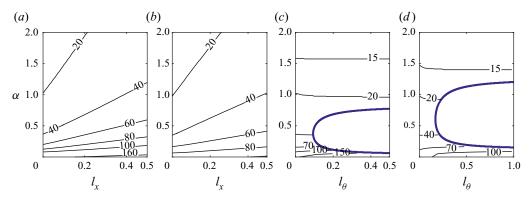


Figure 19. The time instant when G_{max} is reached, t_{max} , in the l_x - α plane for n=1 (a) and n=2 (b) and in the l_{θ} - α plane for n=1 (c) and n=2 (d).

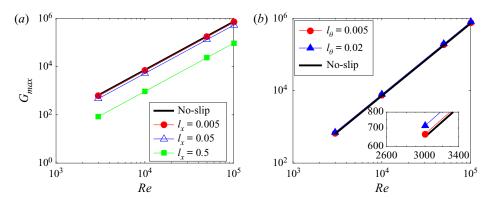


Figure 20. The scaling of global G_{max} with Re. (a) Streamwise slip $l_x = 0.005$, 0.05 and 0.5. (b) Azimuthal slip $l_\theta = 0.005$ and 0.02. In both panels, the bold black line shows the scaling for the no-slip case. The inset in (b) is a close-up around Re = 3000.

azimuthal slip only significantly affects the transient growth for small α and nearly does not affect that of larger α , see the nearly horizontal contour lines for relatively large α . In addition, even for small α , G_{max} quickly saturates as l_{θ} increases. To sum up, azimuthal slip only affects the transient growth of the modes with small streamwise wavenumbers and the effect saturates as the slip length increases.

Figure 19 shows the time instant when G_{max} is reached, t_{max} , for the cases shown in figures 17 and 18. In the streamwise slip case (figure 19a,b), for both n=1 and 2, the slip increases t_{max} and the effect is more significant for larger α . In the azimuthal slip case (figure 19c,d), for small $\alpha (\leq 0.2)$, t_{max} is also slightly increased by the slip, whereas for larger α , t_{max} is slightly decreased by the slip, in contrast to the streamwise slip case. Overall, for the modes with small α , i.e. most amplified modes due to non-normality, both streamwise and azimuthal slip only mildly increase t_{max} .

In the no-slip case, the global G_{max} (maximized over α and n) is known to scale as Re^2 when Re is large (Meseguer & Trefethen 2003). In the presence of streamwise and azimuthal slip, the scaling is also investigated and shown in figure 20. In the streamwise slip case (figure 20a), the lines for $l_x = 0.005$, 0.05 and 0.5 appear to be parallel to the no-slip case with a downward vertical shift, suggesting a Re^2 -scaling for any streamwise slip length. Similarly, for azimuthal slip, the scaling also seems to be Re^2 . For this case,

we only investigated $l_{\theta} = 0.005$ and 0.02 in the linearly stable regime because larger l_{θ} will cause linear instability.

7. Conclusions and discussions

It has been well established that the classic pipe flow is (asymptotically) linearly stable. In this paper, we studied the effect of velocity slip on the linear stability of pipe flow. Our results show that the leading eigenvalue increases with streamwise slip length l_x but remains negative, i.e. streamwise slip renders the flow less stable but does not cause linear instability, similar to the effect of isotropic slip length on the flow (Průša 2009). Interestingly, our results suggest that the leading eigenvalue is independent of Re, or equivalently, the slowest decay rate of disturbances scales as Re^{-1} (note that time is scaled by Re^{-1} in our formulation). It should be pointed out that this scaling holds at sufficiently high Reynolds numbers ($\gtrsim 10^4$). For relatively low Reynolds numbers (100 and 1000 in our study), there is a very slight deviation from the scaling for $l_x \lesssim 0.1$ and the deviation is substantial at larger l_x . The Re^{-1} -scaling of the decay rate is the same as what was observed for the mode $(\alpha, n) = (0, 1)$ of the classic pipe flow by Meseguer & Trefethen (2003). In addition, our results show that the streamwise wavenumber at which the eigenvalue maximizes is not $\alpha = 0$ but also scales as Re^{-1} . However, if l_x is very large (≥ 1.0 , and note that in applications the slip length is generally much smaller), the eigenvalue maximizes at $\alpha = 0$ at relatively low Reynolds numbers (100 and 1000 in our study) and this scaling also only holds at high Reynolds numbers ($\gtrsim 10^4$).

This destabilizing effect appears to be opposite to the stabilizing effect of streamwise slip reported for channel flow (the stabilizing effect is mainly observed for 2-D perturbations) (Lauga & Cossu 2005; Min & Kim 2005; Chai & Song 2019). Here, we only provide a possible explanation for the least stable/most unstable perturbation (referred to as the leading perturbation for simplicity) of the two flows. We speculate that the different flow structures of the leading perturbations of the two flows are responsible. For pipe flow, we proved that the u_r component of the leading perturbation for $\alpha = 0$ modes vanishes. The globally leading perturbation has very small α ($\sim 1/Re$), which indicates that u_r should be nearly vanishing. Therefore, the production rate of kinetic energy

$$-\int_{V} u_{r} u_{x} \frac{\mathrm{d}U_{x}}{\mathrm{d}r} \,\mathrm{d}V \tag{7.1}$$

should be also very small and the decay rate of disturbances should be dominated by the dissipation rate. Intuitively, velocity slip reduces the dissipation rate due to the reduced wall friction, therefore, the decay rate of the least stable perturbation decreases, i.e. the flow appears to be destabilized. In contrast, for channel flow, the leading perturbations are 2-D Tollmien–Schlichting waves for small slip length, which have a substantial wall-normal velocity component (comparable to the streamwise component). Therefore, the kinetic energy production is significant and even dominant in the variation of the kinetic energy. Streamwise slip reduces the base shear (dU_x/dr) and therefore subdues the production. If the reduction in the production rate outweighs the decrease in the energy dissipation rate due to the reduced wall friction, the flow will be stabilized. This is probably why the stabilizing effect of the streamwise slip on channel flow was observed.

On the contrary, azimuthal slip, given sufficiently large slip length, causes linear instability, similar to the findings of Chai & Song (2019) for channel flow. Our results show that azimuthal slip destabilizes helical waves with wavelengths considerably larger than the pipe diameter, whereas it does not affect the stability of waves with much shorter wavelengths and in the long wavelength limit, i.e. $\alpha \to 0$. The critical Reynolds

number decreases sharply as l_{θ} increases and gradually levels off at around a few hundred as $l_{\theta} \gtrsim 0.3$ and at approximately 260 as $l_{\theta} \to \infty$. A similar destabilizing effect was reported for channel flow (Chai & Song 2019). Azimuthal slip serves as an example for a perturbation to the linear operator associated with the linearized Navier–Stokes equations with the no-slip boundary condition that destabilizes the originally stable system.

Regarding the stability of the classic pipe flow to streamwise independent perturbations, using an energy analysis, Joseph & Hung (1971) concluded the absolute and global stability of the flow, i.e. the flow is asymptotically (as $t \to \infty$) stable to such perturbations with arbitrary amplitude. Here, for the linear case and from a mathematical point of view, we rigorously proved that the eigenvalues of streamwise independent modes ($\alpha = 0$) are real and negative, for arbitrary slip length and arbitrary Reynolds number. In addition, the eigenvalue of the $\alpha = 0$ modes is proved to be strictly independent of Reynolds number in our formulation, in agreement with the numerical calculation by Meseguer & Trefethen (2003). We derived analytical solutions to the eigenvalue and eigenvector for $\alpha = 0$ modes and verified our theory by numerical calculations. We also proved that, the eigenvalues of $\alpha = 0$ modes consist of two groups: one group is associated with disturbances with $\Phi \equiv 0$, i.e. $u_r \equiv 0$, and the other is associated with disturbances with $\Phi \not\equiv 0$, i.e. $u_r \not\equiv 0$ (see figures 13 and 15). The two groups distribute alternately. For the streamwise slip case, the latter group remains constant while the former group changes with l_x . It is the other way round for the azimuthal slip case. Interestingly, for the streamwise slip case, the leading eigenvalue belongs to the $\Phi \equiv 0$ group and does not switch group as l_x changes, whereas for the azimuthal slip case, it switches from the $\Phi \equiv 0$ group to the $\Phi \not\equiv 0$ group as l_{θ} crosses 1.0 from below (see figure 15). When both l_x and l_θ are non-zero, l_x dominates the leading eigenvalue if $l_{\theta} < 1$. If $l_{\theta} > 1$, the leading eigenvalue first remains constant and can only start to increase at a threshold as l_x increases (see figure 16). Such analytical solutions might inspire asymptotic analysis in the limit of small streamwise wavenumber.

Non-modal analysis shows that streamwise slip greatly reduces the transient growth, whereas azimuthal slip significantly increases the transient growth for disturbances with very small streamwise wavenumbers but nearly does not affect that for disturbances with larger streamwise wavenumbers, aside from the linear instability caused by the slip. Both streamwise slip and azimuthal slip give the Re^2 -scaling of the maximum transient growth, the same as in the no-slip case (Schmid & Henningson 1994; Meseguer & Trefethen 2003). Similar effects were observed for channel flow (Chai & Song 2019).

Linear instability caused by anisotropic slip at low Reynolds numbers is of interest for small flow systems, such as microfluidics, in which the Reynolds number is usually low but the non-dimensional slip length can be significantly large using advanced surface texturing techniques. The instability can be exploited to enhance mixing or heat transfer in applications involving small flow systems. Larger non-modal growth caused by azimuthal slip can potentially cause earlier subcritical transition to turbulence. In addition, introducing modal instability into originally subcritical flows may also help to better understand the transition mechanism in such flows.

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Author ORCIDs.

Baofang Song http://orcid.org/0000-0003-4469-8781.

Appendix A. Numerics

First, we briefly explain the implementation of the boundary condition (2.12) and (2.13) in the eigenvalue problem for the linear system (2.2).

The eigenvalue equation reads

$$-L^{-1}Mq = \lambda q, \tag{A1}$$

where q is the unknown vector composed of $\hat{\Phi}$ and $\hat{\Omega}$, see (2.5). Boundary conditions (2.12) and (2.13) couple $\hat{\Phi}$ and $\hat{\Omega}$, unlike in the no-slip case. In our Fourier-spectral-Chebyshev-collocation discretization, q is a $2N \times 1$ vector and the operator $-L^{-1}M$ is discretized as a $2N \times 2N$ matrix, where N is the number of collocation grid point on the radius. Adopting the Chebyshev differentiation matrix of Trefethen (2000) which is of spectral accuracy, the matrix is dense and the radial differentiation of q at a single grid point is calculated using the value of q at all collocation points on the radius. Given that $\hat{\Phi}$ is known at r=1, i.e. $\hat{\Phi}=0$ at r=1, the size of the system can be reduced by one. The boundary conditions (2.12) and (2.13) give two linear algebraic equations about $\hat{\Phi}$ and $\hat{\Omega}$ at the collocation points, from which we can further eliminate two unknowns. By doing so, the system size is reduced by three, i.e. to $(2N-3) \times (2N-3)$, from which the eigenvalue problem can be solved with the boundary conditions being taken accounted for.

Second, we show the convergence test of our numerical calculation. We consider the case of Re = 3000, $\alpha = 0.5$ and n = 1, as presented in figures 1(b) and 6(b). We change the number of Chebyshev points and check the convergence of the mean mode, wall mode and centre mode separately. Tables 1-3 show the resolutions and the eigenvalue of an arbitrarily selected mean mode and the rightmost eigenvalues corresponding to the wall mode and centre mode. It can be seen that grid numbers N = 32, 64 and 128 give very close values of the eigenvalues, which differ only after approximately seven digits after the decimal point (the relative difference is $O(10^{-11})$), for all the slip length of 0.005, 0.05, 0.5 and the case of $l_{\theta} = \infty$. The convergence test shows that, for the calculation of the rightmost eigenvalues and for calculating the mean mode at Re = 3000, 32 points are sufficient. Solely for calculating the rightmost eigenvalue, we used 32 points for $Re = 100, 1000 \text{ and } 10000, \text{ and } 64 \text{ points for } Re = 10^5 \text{ and } Re = 10^6.$ We checked the convergence by doubling the number of grid points and found these numbers sufficient. It should be noted that, although 32 points are sufficient for calculating both of the rightmost eigenvalue and the spectrum at Re = 3000, more grid points may be needed for accurately calculating the spectrum than for calculating the rightmost eigenvalue at higher Reynolds numbers (Trefethen 2000; Meseguer & Trefethen 2003).

Appendix B. The dependence of the leading eigenvalue on the azimuthal wavenumber n for $\alpha = 0$

Our numerical calculations in §§ 3 and 4 showed that n=1 modes are the least stable/most unstable modes. Here we show the dependence of max λ of $\alpha=0$ modes on the azimuthal wavenumber n and prove that n=1 is indeed the least stable azimuthal mode. For this purpose, we only need to prove that the minimum non-zero roots of (5.45) and (5.50) all increase with n.

Note that the root of (5.45) is independent of l_x . Because the zeros of $J_n(z)$ and $J_{n+1}(z)$ distribute alternately, it can be easily seen that, if $l_\theta \le 1$, the minimum positive root of (5.45) is located between the minimum positive zeros of $J_n(z)$ and $J_{n+1}(z)$. Therefore, the

$Re = 3000, \alpha = 0.5, n = 1, l_x = 0.005 \text{ and } l_\theta = 0$				
N	mean mode	wall mode		
32	-622.180970763082-1006.803438258427i	-106.901701568167-600.000309196652i		
64	-622.180970617437-1006.803438460110i	-106.901701568129-600.000309195248i		
128	-622.180970549502-1006.803438403115i	-106.901701574121-600.000309193566i		
N	centre mode			
32	-87.144443682906-1299.140360156589i			
64	-87.144443682943-1299.140360156482i			
128	-87.144443681739-1299.140360155874i			
	$Re = 3000, \alpha = 0.5, n = 1, l_x = 0.05 \text{ and } l_\theta = 0$			
N	mean mode	wall mode		
32	-530.922788817816-953.518793751342i	-68.522846170133-646.557599422173i		
64	-530.922788815748-953.518793737632i	-68.522846166147-646.557599425783i		
128	-530.922788808102-953.518793754673i	-68.522846166147-646.557599425783i		
N	centre mode			
32	-80.309678415026-1203.366029301252i			
64	-80.309678415070-1203.366029301237i			
128	-80.309678413737-1203.366029300673i			
$Re = 3000, \alpha = 0.5, n = 1, l_x = 0.5 \text{ and } l_\theta = 0$				
N	mean mode	wall mode		
32	-582.2041335894776-833.8681859837275i	-32.3669400581618-546.6565442837639i		
64	-582.2041335893674-833.8681859850334i	-32.3669400570418-546.6565442835596i		
128	-582.2041336044291-833.8681859835294i	-32.3669400643051-546.6565442852506i		
N	centre mode			
32	-36.8382964038017-757.7599724649968i			
64	-36.8382964035955-757.7599724652722i			
128	-36.8382963977050-757.7599724664044i			

Table 1. The convergence of the eigenvalue corresponding to the mean mode (arbitrarily selected) and the rightmost wall mode and centre mode as the radial grid number N. The streamwise slip cases of $l_x = 0.005$, 0.05 and 0.5 for Re = 3000, $\alpha = 0.5$, n = 1 and $l_\theta = 0$ are listed.

minimum positive root of (5.45) increases with n because the positive zeros of $J_n(z)$ and $J_{n+1}(z)$ all increase with n, i.e. λ_1 decreases as n increases if $l_{\theta} \leq 1$.

If $l_{\theta} > 1$, we need to prove that the minimum positive root of

$$g_n(z) = (1 - l_\theta)J_n(z) + l_\theta z J_{n-1}(z) = 0, \quad n \geqslant 2$$
 (B1)

is smaller than that of $g_{n+1}(z) = 0$, i.e. (5.45). We already showed in (5.54*a*,*b*) that $F(z, l_{\theta}) > 0$ at sufficiently small *z*, i.e. $g_n(z) > 0$ for sufficiently small *z*. Denoting the minimum positive root of (5.45) as z_0 , we only need to show that $g_n(z_0) < 0$. Using the property of Bessel functions, namely

$$J_{n+1}(z) + J_{n-1}(z) = \frac{2n}{z} J_n(z)$$
 (B2)

and (5.45), $g_n(z)$ can be rewritten as

$$g_n(z_0) = (1 - l_\theta)J_n(z_0) + l_\theta z_0 \left(\frac{2n}{z_0} - \frac{l_\theta}{l_\theta - 1}z_0\right)J_n(z_0).$$
 (B3)

$Re = 3000, \alpha = 0.5, n = 1, l_{\theta} = 0.005 \text{ and } l_{x} = 0$				
N	mean mode	wall mode		
32	-878.143428857252-1000.225260064803i	-147.212095816072-565.187041064729i		
64	-878.143429638119-1000.225258436485i	-147.212095816866-565.187041057138i		
128	-878.143429646590-1000.225258451813i	-147.212095817673-565.187041057973i		
N	centre mode			
32	-88.016026669448-1311.995868654545i			
64	-88.016026669520-1311.995868654469i			
128	-88.016026668333-1311.995868653794i			
$Re = 3000, \alpha = 0.5, n = 1, l_{\theta} = 0.05 \text{ and } l_{x} = 0$				
N	mean mode	wall mode		
32	-818.018362966476-992.813175584590i	-30.701258513947-434.347551571943i		
64	-818.018361688704-992.813172416195i	-30.701258511491-434.347551570607i		
128	-818.018361708603-992.813172372158i	-30.701258513706-434.347551572422i		
N	centre mode			
32	-88.015855528115-1312.001178576785i			
64	-88.015855528186-1312.001178576701i			
128	-88.015855526922-1312.001178575825i			
$Re = 3000, \alpha = 0.5, n = 1, l_{\theta} = 0.5 \text{ and } l_{x} = 0$				
N	mean mode	wall mode		
32	-794.161410360041-1006.311883474201i	33.866513836957-391.619593748286i		
64	-794.161408826090-1006.311880389063i	33.866513839707-391.619593747032i		
128	-794.161408715927-1006.311880293383i	33.866513840046-391.619593749138i		
N	centre mode			
32	-88.013747237500-1312.003625793072i			
64	-88.013747237549-1312.003625792987i			
128	-88.013747236338-1312.003625792171i			

Table 2. The convergence of the eigenvalue corresponding to the mean mode (arbitrarily selected) and the rightmost wall mode and centre mode as the radial grid number N. The azimuthal slip cases of $l_{\theta} = 0.005$, 0.05 and 0.5 for Re = 3000, $\alpha = 0.5$, n = 1 and $l_x = 0$ are listed.

$Re = 3000, \alpha = 0.5, n = 1, l_{\theta} = \infty \text{ and } l_{x} = 0$				
N	mean mode	wall mode		
32	-637.771735030362-1019.911762553870i	45.558397965700-383.081240915272i		
64	-637.771735090255-1019.911762536828i	45.558397968779-383.081240913500i		
128	-637.771735157437-1019.911762419706i	45.558397965363-383.081240919212i		
N	centre mode			
32	-88.013312066798-1312.003902287461i			
64	-88.013312066887-1312.003902287400i			
128	-88.013312039270-1312.003902226469i			

Table 3. The convergence of the eigenvalue corresponding to the mean mode (arbitrarily selected) and the rightmost wall mode and centre mode as the radial grid number N. The azimuthal slip case of $l_{\theta} = \infty$ for Re = 3000, $\alpha = 0.5$, n = 1 and $l_x = 0$ is listed.

It is easily seen that $J_n(z_0) > 0$, therefore, we need to show that

$$(1 - l_{\theta}) + l_{\theta} z_0 \left(\frac{2n}{z_0} - \frac{l_{\theta}}{l_{\theta} - 1} z_0 \right) < 0$$
 (B4)

or, equivalently,

$$z_0^2 > (1 - l_\theta^{-1}) \left(2n - 1 + l_\theta^{-1} \right),$$
 (B5)

in order to prove that $g_n(z_0) < 0$, given $l_\theta > 1$. Noticing that $(1 - l_\theta^{-1})(2n - 1 + l_\theta^{-1}) < 2n - 1$ if $l_\theta > 1$, we can prove $g_n(z_0) < 0$ if we can show that $z_0^2 > 2n - 1$.

In § 5.3 we proved that the minimum positive root of (5.45) decreases as l_{θ} increases, therefore, z_0 is minimized at $l_{\theta} = +\infty$. To prove that $z_0^2 > 2n-1$ for any $l_{\theta} > 1$, we only need to show that $z_0^2 > 2n-1$ holds for $l_{\theta} = +\infty$, with which (5.45) reduces to $J_{n+1}(z) - zl_{\theta}J_n(z) = 0$. That $F(z, l_{\theta}) > 0$ at sufficiently small z indicates that $J_{n+1}(z) - zJ_n(z) < 0$ at sufficiently small z given $l_{\theta} = +\infty$. Therefore, $J_{n+1}(z_0) - z_0J_n(z_0) = 0$ requires that $J_{n+1}(z) - zJ_n(z) < 0$, i.e. $J_{n+1}(z) < zJ_n(z)$ for any $z < z_0$. In fact, we can show that $J_{n+1}(z) < zJ_n(z)$ if $0 < z^2 < 2n-1$ such that z_0 has to satisfy $z_0^2 > 2n-1$. Using the series form of Bessel function, $J_{n+1}(z) < zJ_n(z)$ means

$$\sum_{k=0}^{+\infty} \frac{(-1)^2}{k!} \frac{1}{(n+k+1)!} \left(\frac{z}{2}\right)^{2k+n+1} < z \sum_{k=0}^{+\infty} \frac{(-1)^2}{k!} \frac{1}{(n+k)!} \left(\frac{z}{2}\right)^{2k+n}.$$
 (B6)

Because of the absolute convergence of the two infinite series in (B6), we only need to show that for any positive even number k,

$$\frac{1}{k!} \frac{1}{(n+k+1)!} \left(\frac{z}{2}\right)^{2k+n+1} - \frac{1}{(k+1)!} \frac{1}{(n+k+2)!} \left(\frac{z}{2}\right)^{2k+2+n+1} < \frac{z}{k!} \frac{1}{(n+k)!} \left(\frac{z}{2}\right)^{2k+n} - \frac{z}{(k+1)!} \frac{1}{(n+k+1)!} \left(\frac{z}{2}\right)^{2k+2+n}, \tag{B7}$$

if $0 < z^2 < 2n - 1$. Rearranging (B7), we have

$$z^2 < 4(k+1)(n+k+2),$$
 (B8)

which obviously holds because $z^2 < 2n - 1 < 4n + 4k + 8 \le 4(k+1)(n+k+2)$. To sum up, we have proved that $J_{n+1}(z) < zJ_n(z)$ if $z^2 < 2n - 1$, therefore, z_0 , which satisfies $J_{n+1}(z_0) - z_0J_n(z_0) = 0$, must satisfy $z_0^2 > 2n - 1$. Consequently, (B5) and $g_n(z_0) < 0$ hold, and thus the minimum positive root of (B1) is smaller than z_0 , which indicates that the minimum positive root of (5.45) increases with n, i.e. λ_1 decreases with n.

Next, we prove that the minimum positive root of (5.50) also increases with n. The equation

$$h_n(z) = (1 + nl_x)z^n J_n(z) - l_x z^{n+1} J_{n+1}(z) = 0$$
(B9)

shares the non-zero roots with (5.50), therefore, we only need to prove the same statement for (B9). Denoting the minimum positive zero of $h_{n-1}(z)$ as z_0 , we next show that $h_n(z)$ monotonically increases in $[0, z_0]$ such that there is no positive root of (B9) in $[0, z_0]$, i.e. the minimum positive root of (B9) increases with n. Using the property of

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Bessel functions, namely $(z^{n+1}J_{n+1}(z))' = z^{n+1}J_n(z)$, where "denotes the derivative with respect to z, we take the derivative of $h_n(z)$ with respect to z and obtain

$$h'_n(z) = (1 + nl_x)z^n J_{n-1}(z) - l_x z^{n+1} J_n(z) = z \left(h_{n-1}(z) + l_x z^{n-1} J_{n-1}(z) \right).$$
 (B10)

It is easy to see that $h_{n-1}(z)$ is positive at sufficiently small z (the derivation is similar to that of $F(z, l_{\theta})$ being positive at sufficiently small z, see (5.54a,b)), consequently, $h_{n-1}(z) > 0$ in $(0, z_0)$. As z_0 is smaller than the minimum positive zero of $J_{n-1}(z)$, we have $J_{n-1}(z) > 0$ in $(0, z_0)$. Therefore, $h'_n(z) > 0$ in $(0, z_0)$, i.e. $h_n(z)$ monotonically increases and there is no positive root of (B9) in $(0, z_0]$. In other words, the minimum positive root of $h_n(z) = 0$ is always larger than that of $h_{n-1}(z) = 0$, i.e. the minimum positive root of (5.50) increases with n, and therefore λ_2 decreases with n. Now, we reach the conclusion that $\max \lambda = \max\{\lambda_1, \lambda_2\}$ decreases with n for $\alpha = 0$ modes because λ_1 and λ_2 both decrease with n.

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