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CO-CHAIRS

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1. Introduction

Research in atomic and molecular collision processes and spectral line broadening has been very active since our last report (Schultz & Stancil 2007, Allard & Peach 2007). Given the large volume of the published literature and the limited space available, we have attempted to identify work most relevant to astrophysics. Since our report is not comprehensive, additional publications can be found in the databases at the web addresses listed in the final section. Elastic and inelastic collisions among electrons, atoms, ions, and molecules are included and reactive processes are also considered, but except for charge exchange, they receive only sparse coverage.

Numerous meetings on collision processes and line broadening have been held throughout the report period. Important international meetings that provide additional sources of data through their proceedings are: the XXIV *International Conference on Photonic, Electronic, and Atomic Collisions* (ICPEAC) (Fainstein *et al.* 2006), XXV ICPEAC (Becker *et al.* 2007), the NASA *Laboratory Astrophysics Workshop* (Weck *et al.* 2006), the 18th *International Conference on Spectral Line Shapes* (ICSLS) (Oks & Pindzola 2006) and the VIth *Serbian Conference on Spectral Line Shapes in Astrophysics* (SC-SLSA) (Popović & Dimitrijević 2007). The 19th ICSLS has just taken place in June 2008.

2. Electron collisions with atoms, ions, molecules, and molecular ions

Collisions of electrons with atoms, ions, molecules, and molecular ions are the major excitation mechanism for a wide range of astrophysical environments. In addition, electron collisions play an important role in ionization and recombination, contribute to cooling and heating of the gas, and may contribute to molecular fragmentation and formation. In the following sections we summarize recent work on electron collisions with astrophysically relevant species, including elastic scattering, excitation, dissociation, ionization, recombination, and electron detachment from negative ions.

2.1. *Electron-atom scattering*

New work on elastic scattering from neutral atoms is limited to Xe (Linert *et al.* 2007), Cs (Zatsarinny & Bartschat 2008), In (Rabasović *et al.* 2008) and Au (Maslov *et al.* 2008). The excitation of atomic oxygen has been investigated (Wang & Zhou 2006, Barklem 2007), while new work on ionization has been carried out for Mg (Bolognesi *et al.* 2008).

2.2. *Electron-ion scattering*

For atomic ions, new work has primarily focused on excitation and includes: C⁺ (Wilson *et al.* 2005), N⁺ (Hudson & Bell 2005), O⁺ (Tayal 2007), Al¹²⁺ (Aggarwal *et al.* 2005,

Aggarwal *et al.* 2008), S^{4+} (Hudson & Bell 2006), Ar^{16+} (Aggarwal & Keenan 2005b), Ca^+ (Meléndez *et al.* 2007), Kr^{6+} (Ishikawa & Vilkas 2008), Fe^+ (Ramsbottom *et al.* 2007), Fe^{3+} (McLaughlin *et al.* 2006), Fe^{4+} (Ballance *et al.* 2007), Fe^{6+} (Witthoef & Badnell 2008), Fe^{9+} (Aggarwal & Keenan 2005a), Fe^{11+} (Storey *et al.* 2005), Fe^{15+} (Aggarwal & Keenan 2006), Fe^{17+} (Witthoef *et al.* 2006), Fe^{19+} (Witthoef *et al.* 2007), Fe^{22+} (Chidichimo *et al.* 2005), Fe^{25+} (Aggarwal *et al.* 2008), Ni^{3+} (Meléndez & Bautista 2005), and for Si, Cl, and Ar isonuclear sequences (Colgan *et al.* 2008).

New elastic data exists for He^+ and Li^{2+} (Bhatia 2008). Ionization and detachment have been studied for H^- (Jung 2008); C^{2+} , N^{3+} , and O^{4+} (Fogle *et al.* 2008); Ne^{4+} and Au^{47+} (Pindzola *et al.* 2008); and Si, Cl, and Ar isonuclear sequences (Colgan *et al.* 2008).

Another important process is recombination for which a number of new works including radiative and dielectronic recombination have appeared: C^{2+} , N^{3+} , and O^{4+} (Fogle *et al.* 2005); Si^{3+} (Schmidt *et al.* 2007); Mg^{2+} (Fu *et al.* 2008); and Fe^{16+} (Chen 2008).

2.3. Electron-molecule scattering

For molecules, new elastic scattering references have appeared as follows: C_3 (Munjál & Baluga 2006); CH_4 (Cho *et al.* 2008); H_2CO (Kaur & Baluja 2005); C_6H_6 , C_6F_6 , C_6H_{12} , C_6H_{14} , C_6F_{14} , C_8H_{16} , C_8H_{18} , and C_8F_{18} (Shi *et al.* 2008); SF_4 (Szmytkowski *et al.* 2005); SO_2Cl_2 (Szmytkowski *et al.* 2006); SiN_2 , $SiCO$, and $CSiO$ (Fujimoto *et al.* 2007); pyrazine (Winstead 2007); propane (Bettega, da Costa, & Lima 2008); methanol and ethanol (Khakoo *et al.* 2008); propene and cyclopropane (Makochekanwa *et al.* 2008); and NeF (Kaur *et al.* 2008).

For excitation, new references include: H_2 (da Costa *et al.* 2005, Kato *et al.* 2008); C_3 (Munjál & Baluga 2006); CH_2 (Allan 2007); CH_4 (Čurik *et al.* 2008); CO and C_2H_4 (da Costa *et al.* 2007); CO_2 (Rescigno *et al.* 2007); CF_4 (Irrera & Gianturco 2005); N_2 (Tashiro & Morokuma 2007, Khakoo *et al.* 2007), ethene (Allan, Winstead, & McKoy 2008); and NeF (Kaur *et al.* 2008).

New work for dissociative processes are for dissociative electron attachment to H_2O (Haxton *et al.* 2007), C_2H_2 (Chourou & Orel 2008, May *et al.* 2008), and C_4H_2 (May *et al.* 2008). Research on molecular ionization has been limited to H_2 and D_2 (Martín 2007) and H_2O (Kaiser *et al.* 2007).

2.4. Electron-molecular ion scattering

References on dissociative processes have appeared for: H_2^+ (Motapon *et al.* 2008), He_2^+ (Buhr *et al.* 2008), NeH^+ and NeD^+ (Ngassam *et al.* 2008), H_3^+ (Kokoouline & Greene 2005), D_2H^+ (Zhaunerchyk *et al.* 2008b), HCO^+ and DCO^+ (Douguet *et al.* 2008), O_3^+ (Zhaunerchyk *et al.* 2008a), and $HCNH^+$ (Ngassam *et al.* 2005).

For excitation, a study has been conducted for H_3^+ (Faure *et al.* 2006). New results for ionization include H_2^+ (Pindzola *et al.* 2005), while detachment has been investigated for Si_2^- (Lindahl *et al.* 2008). Finally, vibrational excitation due to electron impact has been studied for NeH^+ and NeD^+ (Ngassam *et al.* 2008).

3. Heavy particle collisions

3.1. Ion-atom and atom-atom collisions

Charge exchange has seen a substantial amount of activity over the report period as it plays an important role in a variety of environments. Studies for collisions on H include: H^+ (Bradley *et al.* 2005, Dubois *et al.* 2005, Zeng *et al.* 2008), He^{2+} (Havener *et al.* 2005), Be^{4+} (Minami *et al.* 2006), C^{6+} (Liu *et al.* 2005), N^{2+} and O^{2+} (Barragán *et al.*

2006a, Barragán *et al.* 2006c), O^{8+} (Perez & Olson 2005), F^{2+} (Dutta *et al.* 2005), Ne^{10+} (Errea *et al.* 2005a, Barragán *et al.* 2006b), Si^{3+} (Wang *et al.* 2006, Bruhns *et al.* 2008), S^{16+} (Janowicz *et al.* 2005), Cl^{7+} (Zhao *et al.* 2007), and Ar^{18+} (Errea *et al.* 2005a). A database for cross sections for all carbon ions colliding with hydrogen has been constructed by Suno & Kato (2006).

Neutral helium is also an important target for which studies have been carried out for the incident ions: He^+ (Bradley *et al.* 2005), C^{4+} (Hoshino *et al.* 2007), F^{7+} (Zouros *et al.* 2008), and $Ne^{(2-6)+}$ (Hasan 2005).

Charge exchange due to proton impact on Ca (Dutta *et al.* 2006, Pandey & Dubey 2007) and Mg (Pandey & Dubey 2007), alpha particles on Na (Lee 2006), and N^{7+} on atomic oxygen (Perez & Olson 2005) have been studied, while radiative charge transfer in Ne^{2+} collisions with He (Zhao *et al.* 2006) has been investigated.

Elastic scattering due to proton impact on He, Ne, and Ar (Ovchinnikov *et al.* 2006) have been studied. Inelastic processes involving hyperfine changing collisions have been investigated for: $H + H$ (Zygelman 2005); $(n - n')$ -mixing in $H^*(n) + H(1s)$ collisions (Mihajlov *et al.* 2005); fine structure transitions for $O + H$ and $C + H$ (Abrahamsson & Krems 2007), $C^+ + H$ and $Si^+ + H$ (Barinovs *et al.* 2005) and depolarization collisions of excited atoms and ions with hydrogen (Derouich *et al.* 2005b, Derouich *et al.* 2005a, Derouich 2007, Derouich & Barklem 2007, Sahal-Bréchet *et al.* 2007). Electronic transitions have been studied for $O + H$ (Krems *et al.* 2006) and H impact on S^{3+} , Ar^{13+} , and Fe^{13+} (Burgess & Tully 2005).

Excitation, charge transfer, and ionization due to He^{2+} collisions with H including Debye screening in a dense plasma has been studied by Liu *et al.* (2008a), Liu *et al.* (2008b). Electron detachment in He collisions with H^- (Huang *et al.* 2005, Ogurtsov *et al.* 2006) and C^- (Huang *et al.* 2005) has also been investigated. Rate coefficients have been computed for the formation of HD (Dickinson 2005); CH^+ (Barinovs & van Hemert 2005); and SO, SO^+ , and S_2 (Andreazza & Marinho 2005) by radiative association.

3.2. Ion-, atom-, and molecule-molecule collisions

In photo-ionized environments, multiply charged ions may coexist with neutral molecules. Examples include x-ray ionized regions and solar wind interactions with cometary gas. In these environments charge transfer plays an important role. Recent studies of ion-molecule charge transfer include He^{2+} (Dubois *et al.* 2005, Kusakabe *et al.* 2006), C^{4+} (Zarour *et al.* 2005), O^+ (Kimura *et al.* 2006), O^{5+} and Ar^{5+} (Dubois *et al.* 2005), and F^{7+} (Zouros *et al.* 2008) with H_2 ; H^+ (Lindsay *et al.* 2005a, Wells *et al.* 2005, Kumar *et al.* 2006, Lin *et al.* 2007) and He^{2+} (Kusakabe *et al.* 2006) with CO; He^{2+} (Abu-Haija *et al.* 2005b, Kusakabe *et al.* 2006), $Ar^{5,4+}$ (Abu-Haija *et al.* 2005a), and Kr^{8+} (Kaneyasu *et al.* 2005) with N_2 ; H^+ and O^+ (Luna *et al.* 2005), He^{2+} (Abu-Haija *et al.* 2005b, Kusakabe *et al.* 2006), and $Ar^{5,4+}$ (Abu-Haija *et al.* 2005a) with O_2 ; H^+ (Kimura *et al.* 2006), He^{2+} (Bodewits, *et al.* 2005, Seredyuk *et al.* 2005c), O^{6+} (Seredyuk *et al.* 2005b, Bodewits & Hoekstra 2007), and a range of ions (Otranto & Olson 2008) with water; H^+ (Lindsay *et al.* 2005a), He^{2+} (Abu-Haija *et al.* 2005b, Kusakabe *et al.* 2006), and O^{6+} (Seredyuk *et al.* 2005b) with CO_2 ; H^+ with NH_2 (Suno *et al.* 2006); He^{2+} (Abu-Haija *et al.* 2005b, Kusakabe *et al.* 2006) with NH_3 ; H^+ (Lindsay *et al.* 2005b), He^{2+} (Seredyuk *et al.* 2005a), and C^{4+} and O^{6+} (Seredyuk *et al.* 2005b) with CH_4 ; H^+ (Suzuki *et al.* 2005) and He^{2+} (Seredyuk *et al.* 2005a) with C_2H_4 ; and He^{2+} (Seredyuk *et al.* 2005a) with C_2H_6 . Other charge exchange studies include $H_2^+ + H$ (Errea *et al.* 2005b); $N_2^+ + N_2$ and $O_2^+ + O_2$ (Tong & Nanbu 2007); and $O^+ + O_2$ (Martinez *et al.*

2006), while Cornelius (2006) has developed a scaling relation for total cross sections with H_2 .

For applications to x-ray emission from comets and planetary atmospheres, charge exchange has been considered for molecular targets including $\text{C}^{(3-6)+}$, $\text{N}^{(4-7)+}$, and $\text{O}^{(5-7)+}$ with methane (Djurić *et al.* 2008); highly charged L-shell Fe ions with various neutrals (Wargelin *et al.* 2005, Beiersdorfer *et al.* 2008); and for a range of ions and molecules (Otranto *et al.* 2006).

The internal level populations of molecular rovibrational states are primarily controlled through collisional excitation by atom and molecule impact. Investigations have been carried out for excitation of H_2 by H (Wrathmall & Flower 2006, Wrathmall *et al.* 2007), He (Mack *et al.* 2006), H_2 (Lee *et al.* 2006), and H^- (Giri & Sathyamurthy 2006); HF by He (Reese *et al.* 2005); CO by H (Yang *et al.* 2006a, Shepler *et al.* 2007), He (Yang *et al.* 2006a), and H_2 (Wernli *et al.* 2006, Yang *et al.* 2006b); CO^+ by H (Andersson *et al.* 2008); CN by C_2H_2 (Olkhov & Smith 2007); CS by He (Lique *et al.* 2006b, Lique & Spielfiedel 2007); PN by He (Tobola *et al.* 2007); SiO by H (Palov *et al.* 2006) and He (Dayou & Balança 2006); SiS by He (Vincent *et al.* 2007); SO by He (Lique *et al.* 2006a, Lique *et al.* 2006c); H_2O by He (Yang & Stancil 2007) and H_2 (Dubernet *et al.* 2006), NH_3 (Machin & Roueff 2005, Yang & Stancil 2008), NH_2D (Machin & Roueff 2006), and ND_2H by He (Machin & Roueff 2007); and HC_2N by He and H_2 (Wernli *et al.* 2007).

Other investigations include: collisional dissociation of highly excited H_2 by He Ohlinger *et al.* (2007), fragmentation of CO by slow C^{6+} and Ar^{11+} (Wells *et al.* 2008), and dissociation and fragmentation of N_2 by slow Xe^{q+} ions, ($q=15-21$), (Zhu *et al.* 2005). Mutual neutralization in H_2^+ collisions with H^- has been studied by Liu *et al.* (2006) and the formation of HeH_2^+ via radiative association by Mrugala & Kraemer (2005).

4. Reactive scattering and chemistry

Due to space limitations, we cannot review the many advances in reactive scattering and chemical processes relevant to astrophysics. One noteworthy and relevant study involves a quasi-classical trajectory investigation of $\text{H} + \text{CH}_4 \rightarrow \text{H}_2 + \text{CH}_3$ (Xie *et al.* 2006), and updates to the UMIST Astrochemistry database which gives rate coefficient fits for 4572 reactions, has been completed recently (Woodall *et al.* 2007).

5. Stark broadening

Knowledge of line widths and shifts for atomic transitions is very important for the interpretation of stellar spectra and also for circumstellar conditions and galactic H II regions.

The *Critical Review of Selected Data on Experimental Stark Widths and Shifts for Spectral Lines of Neutral and Ionized Atoms* for the period 2001-2007, see Lesage (2008), contains tables where measured values are listed and compared with semi-classical calculations. A book entitled *Stark Broadening of Hydrogen and Hydrogenlike Spectral Lines in Plasmas* has been published by Oks (2006) and contains many useful references.

5.1. Developments in line broadening theory

Poquérousse & Alexiou (2006) have extended standard semi-classical impact theory for hydrogen-like ions to include penetrating collisions. For transitions in highly excited hydrogen-like atoms and ions Stambulchik & Maron (2008) have produced a simple analytical method for calculating line shapes and Gigosos *et al.* (2007) have developed

an exact expression for the impact broadening operator of hydrogen. The asymmetry of hydrogen lines has been studied by Demura *et al.* (2008a) and Demura *et al.* (2008b).

Stambulchik & Maron (2008) have studied broadening of lines subject to external electric and magnetic fields and Godbert-Mouret *et al.* (2006) have developed new code to calculate their lineshapes. Dubau *et al.* (2007) have carried out quantum mechanical calculations of electron impact broadening for XUV lines in plasmas.

5.2. Isolated lines

For isolated lines, Stark broadening is dominated by collisions with plasma electrons. Broadening parameters have been determined theoretically for:

Ar II 476.5 nm, 480.6 nm and Kr II 469.4 nm lines (Dimitrijević & Csillag 2006), Cd I 33 singlets and 37 triplets (Simić *et al.* 2005a), 26 Ne V multiplets (Hamdi *et al.* 2007) and 15 Si VI multiplets (Hamdi *et al.* 2008).

Also for: 4 N IV, 3 O V, 1 F VI and 4 Ne IV (Elabidi *et al.* 2008a), 2 N II, O III, F IV and Ne V (Ivković *et al.* 2005), O II (Mahmoudi *et al.* (2005)), F II (Srećković *et al.* 2005b), 1 F III (Simić *et al.* 2005b), 6 Ar I (Dimitrijević *et al.* 2007a), 9 Cr I (Dimitrijević *et al.* 2005), 3 Mn II (Popović *et al.* 2008), 3 Te I (Dimitrijević *et al.* 2007b, Dimitrijević *et al.* 2008), F III (Simić *et al.* 2005b), Ne VII, Ne VIII and Si XI (Elabidi *et al.* 2008b), S II, S III and S IV (Milovanović N. & Dimitrijević 2007), Cu III, Zn III and Se III (Simić *et al.* 2006), Ga II (N'Dollo & Donga-Passi 2006), Sn II (Colón & Alonso-Medina 2006) and Pb III (Alonso-Medina *et al.* 2008) lines.

Broadening parameters have been obtained experimentally for:

C I 247.8561 nm (Djeniže *et al.* 2006b), Mg II 448.1 nm (Djeniže *et al.* 2005b), Fe I 381.58 nm (Bengoechea *et al.* 2006) and Fe I 538.34 nm (Bengoechea *et al.* 2005) lines.

Also for: 2 N I (Bartecka *et al.* 2005), 3 O II (Bukvić *et al.* 2005), 23 O III (Srećković *et al.* 2005a), 10 N II and 8 O III (Ivković *et al.* 2005), F II (Srećković *et al.* 2005b), 4 Ne I (Dzierżega *et al.* 2006), 13 Ne I (Jovićević *et al.* 2005), 7 Ar I (Milosavljević *et al.* 2006), 6 Ar II (Iglesias *et al.* 2006), 6 Mn I (Srećković *et al.* 2007), 11 Mn II and 3 Mn III (Djeniže *et al.* 2006e), 17 Ni II (Mayo *et al.* 2008), 35 Kr II (del Val *et al.* 2008), 2 Ag I and 2 Au I (Djeniže *et al.* 2006c), 2 Ag I, 11 Ag II and 3 Ag III (Djeniže *et al.* 2005a), 26 Au II (Ortiz & Mayo 2005), 43 Sn I and 27 Sn II (Alonso-Medina & Colón 2008), 12 Sn I and 16 Sn II (Djeniže *et al.* 2006d), 16 In III (Djeniže *et al.* 2006a), 10 Pb III lines (Alonso-Medina & Colón 2007), 31 Pb II (Colón & Alonso-Medina 2006), 38 Xe III (Peláez *et al.* 2006a) and shifts of 110 Xe and 42 Xe III (Ćirišan *et al.* 2006) lines.

Djurović *et al.* (2006) have presented a review of experimental work on Stark broadening of 80 singly ionized xenon lines and Purić *et al.* (2008) have used published Stark widths for spectral lines originating from 3s-3p transition arrays of multiply charged ions, to establish trends from which Stark widths are predicted for Mg VII, Mg IX, Mg X, Na VII, Na VIII, Al VIII, Al IX, Si XI, Ti XI, Cr XIII, Cr XIV, Fe XV, Fe XVI, Fe XXIII and Ni XVIII.

Stehlé *et al.* (2005) have examined current Stark broadening theory as a basis for diagnostics of low-temperature plasmas and Mahmoudi *et al.* (2008) have provided new expressions for diagonal multiplet factors of complex configurations, required for studies of isolated lines. Zmerli *et al.* (2008) have proposed an improved interpolation method for widths as a function of temperature.

5.3. Transitions in hydrogenic and helium-like systems

New quantum mechanical calculations of the broadening of Ly β , Ly γ and Ly δ have been carried out by de Kertanguy *et al.* (2005). Transitions in the Balmer series have been

studied by Gigosos & González (2006), and by Stambulchik *et al.* (2007) for $n \leq 15$. Broadening of high- n transitions are considered by Lisgo *et al.* (2006) and benchmarked against electron density measurements. New experiments for $H\beta$ for the wide range of plasma parameters have been carried out by Djurović *et al.* (2005) and Griem *et al.* (2005) compare $H\alpha$ profiles measured at high electron densities with theoretical results.

Broadening of the radio recombination lines of hydrogen has been studied theoretically by Watson (2006) and Gavrilenko & Oks (2007) and lines of hydrogen-like and helium-like ions of C, Si and Ar have been examined by Stambulchik & Maron (2006) who find that ion dynamics is very important.

New theoretical calculations of broadening have been reported for He I 667.8 nm and 587.6 nm lines (Ben Chaouacha *et al.* (2007)) and He I 728.1, 706.5, 504.8, 492.2 and 471.3 nm lines in a dense plasma (Omar *et al.* (2006)). Peláez *et al.* (2006b) have carried out experiments for He I 318.8 nm and 402.6186 nm lines.

6. Broadening by neutral atoms and molecules

The analysis of experimental molecular spectra in order to extract line shape parameters is often very difficult. Line shapes can be affected by collisional narrowing and the dependence of collisional broadening and shifting on molecular speed. When these effects are sufficiently important, fitting Voigt profiles to experimental spectra produces systematic errors in the parameters retrieved.

A collection of papers concerning the status of the molecular spectroscopic database, HITRAN 2000, has been published by Rothman *et al.* (2003) and this has recently been updated for the current version HITRAN 2004 by Rothman *et al.* (2005).

6.1. Broadening of atomic lines

Some theoretical work has been published in the period 2005-2008 and the transitions with the perturbing atoms or molecules are listed below.

Li; 2s-3s transition broadened by Ar, Kr and Xe (Rosenberry *et al.* 2007).

Li; wings of the resonance line broadened by He and H_2 (Allard *et al.* 2005).

Na and K; wings of the resonance lines broadened by He (Zhu *et al.* 2006).

Li, Na and K; impact widths for the resonance lines broadened by He (Mullamphy *et al.* 2007).

Rb and Cs; resonance line profiles, including far line wings, broadened by He and H_2 (Allard & Spiegelman 2006).

Fe II; 24188 lines broadened by collisions with H (Barklem & Aspelund-Johansson 2005).

Sr: $5s^2^1S_0 \rightarrow 5s5p^3P_1$ and $5s5p^3P_{0,1,2} \rightarrow 5s6s^3S_1$ transitions broadened by the rare gases (Holtgrave & Wolf 2005).

6.2. Broadening and shift of molecular lines

Much new data have been published since the last report was prepared. The molecules are listed below with their perturbing atomic or molecular species and are labelled by 'E' and 'T' to indicate experimental work and theoretical analysis respectively.

H_2 - H_2 collision-induced absorption (T) (Orton *et al.* 2007); in binary mixtures H_2 - N_2 and H_2 -CO (E) (Abu-Kharma *et al.* 2006).

H_2 lines broadened and shifted by He (T) (Ma *et al.* 2007).

HDO lines broadened and shifted by N_2 (E) (Bach *et al.* 2005).

HCN lines broadened and shifted by HCN and air (E) (Devi *et al.* 2005); N_2 (E) (Smith *et al.* 2008); H_2 , N_2 , O_2 , CH_3CN and rare gases (E) (Rohart *et al.* 2007).

HC_3N lines broadened by H_2 , He and N_2 (E) (Colmont *et al.* 2007b).

H₂CO lines broadened by H₂CO, N₂ and O₂ (E) (Staak *et al.* 2005).

HCO⁺ lines broadened by He and Ar (T) (Buffa 2007).

HNO₃ lines broadened by air (E) (Cazzoli *et al.* 2005).

HO₂ lines broadened by air (E) (Ibrahim *et al.* 2007); H₂O and N₂ (E) (Kanno *et al.* 2005).

HI lines broadened by HI (E) (Bulanin *et al.* 2005, Hartmann *et al.* 2005); He (E) (Flaud *et al.* 2006).

HI, HBr lines broadened by rare gases (E) (Domanskaya *et al.* 2007).

H₂O lines broadened by H₂O (T) (Tolchenov & Tennyson 2005, Ptashnik *et al.* 2005, Antony & Gamache 2007, Antony *et al.* 2007); air (E) (Liu *et al.* 2007, Seta *et al.* 2008); H₂O and air (E+T) (Toth 2005, Jenouvrier *et al.* 2007, Ibrahim *et al.* 2008); N₂ (E+T) (Aldener *et al.* 2005, Bandyopadhyay *et al.* 2007, Tran *et al.* 2007, Bykov *et al.* 2008); N₂ and air (E+T) (Aldener *et al.* 2005, Bandyopadhyay *et al.* 2007, Tran *et al.* 2007, Bykov *et al.* 2008, Hodges *et al.* 2008); N₂ and O₂ (E) (Golubiatnikov *et al.* 2008, Hoshina *et al.* 2008); H₂O, N₂ and O₂ (E) (Cazzoli *et al.* 2007, Koshelev *et al.* 2007, Cazzoli *et al.* 2008); H₂O and Ar (E) (Li *et al.* 2008); H₂O, H₂, N₂, O₂, CO₂ and rare gases (E) (Golubiatnikov 2005); H₂O, H₂, N₂, O₂, CO₂, He and air (E) (Brown *et al.* 2005).

CH₄ lines broadened by air (E) (Predoi-Cross *et al.* 2006); CH₄ and air (E) (Predoi-Cross *et al.* 2007a); CH₄, air, He and H₂ (E) (Lucchesini & Gozzini 2007); CH₄ (E) (Lepère 2006, Wishnow *et al.* 2007); H₂ and He (T) (Tran *et al.* 2006); N₂ (E) (Mondelain *et al.* 2007, Martin & Lepère 2008); CH₄ and N₂ (E) (Menard-Bourcin *et al.* 2007); N₂ and O₂ (E) (Mondelain *et al.* 2005, Lepère *et al.* 2005); N₂, O₂ and air (T) (Antony *et al.* 2008). C₂H₂ self-broadened lines (E+T) (Lepère *et al.* 2007, Nguyen *et al.* 2008, Lyulin *et al.* 2008); broadened by He (T) (Thibault 2005, Nguyen *et al.* 2006); H₂, D₂, N₂, air and rare gases (E) (Arteaga *et al.* 2007); CO₂ (E+T) (Martin *et al.* 2006).

C₂H₄ lines broadened by N₂ (E) (Blanquet *et al.* 2005).

C₃H₂ lines self broadened (E) (Achkasova *et al.* 2006).

CH₃Br lines broadened by N₂ (E+T) (Jacquemart *et al.* 2007, Tran *et al.* 2008); CH₃Br and N₂ (E) (Jacquemart & Tran 2008).

CH₃F lines broadened by H₂ (E) (Lerot *et al.* 2006b); N₂, O₂ (E) (Lerot *et al.* 2006a); CH₃F (Lerot *et al.* 2005).

CH₃CN broadened by CH₃CN and N₂ (E) (Rinsland *et al.* 2008).

CO lines broadened by He (E) (Thibault *et al.* 2007); CO₂ (E) (Sung & Varanasi 2005); Ar (E+T) (Wehr *et al.* 2006a, Wehr *et al.* 2006b); N₂, O₂, CO₂ and rare gases (Colmont *et al.* 2007a).

CO₂ lines broadened by air (E) (Predoi-Cross *et al.* 2007c, Toth *et al.* 2007); self-broadened lines (E) (Hikida *et al.* 2005, Le Barbu *et al.* 2006, Predoi-Cross *et al.* 2007b), (T) (Toth *et al.* 2006); air and CO₂ (E) (Devi *et al.* 2007a, Toth *et al.* 2008, Joly *et al.* 2008, Devi *et al.* 2007a); air and Ar (E) (Li *et al.* 2008); N₂ and O₂ (E) (Hikida & Yamada 2006).

Cs₂ lines broadened by N₂ (E+T) (Misago *et al.* 2006); O₂ (E+T) (Misago *et al.* 2007); Ar and air (E+T) (Misago *et al.* 2008).

N₂ self-broadened lines (E) (El-Kader & Moustafa 2005, Hashimoto & Kanamori 2006); N₂-H₂ collision-induced absorption (E) (Boissoles *et al.* 2005).

NH₃ lines broadened by NH₃ (E) (Leary *et al.* 2008); N₂, O₂ and air (Dhib *et al.* 2007).

N₂O lines broadened by air (E) (Grossel *et al.* 2008).

O₂ self-broadened lines (E) (Tretyakov *et al.* 2007, Predoi-Cross *et al.* 2008a); lines broadened by O₂ and N₂ (Tretyakov *et al.* 2005); N₂ (E+T) (Predoi-Cross *et al.* 2008b).

O₃ lines broadened by O₃ (E) (Yamada & Amano 2005); N₂ and O₂ (E+T) (Rohart *et al.* 2008).

OCS lines broadened by OCS (E) (Matton *et al.* 2006); N₂ and O₂ (E) (Koshelev *et al.* 2006).

PH₃ lines broadened by N₂ (E+T) (Bouanich *et al.* 2005, Bouanich & Blanquet 2007).

SO₂ self-broadened lines (E+T) (Zéninari *et al.* 2007, Henningsen *et al.* 2008).

7. Databases

A database for atomic and molecular processes is maintained at the Oak Ridge National Laboratory Controlled Fusion Atomic Data Center (CFADC) at the address <cfadc.phy.ornl.gov> and a useful on-line database of rovibrational collisional excitation data, BASECOL, can be found at <basecol.obspm.fr>.

Some collisional data are also available on the Leiden Atomic and Molecular Database <www.strw.leidenuniv.nl/~moldata> and the UMIST Astrochemistry database is at <www.udfa.net>.

A 'virtual observatory' for astronomers can be found at <cdsarc.u-strasbg.fr> and the latest version of the database High resolution Transmission, HITRAN 2004, is at <www.hitran.com>.

The current version of the database Gestion et Etude des Informations Spectroscopiques Atmosphériques (GEISA-03) is at <ara.lmd.polytechnique.fr> and the Spherical Top data System (STDS) has address <icb.u-bourgogne.fr/OMR/SMA/SHTDS/STDS>.

The National Institute for Standards and Technology (NIST) maintains a database at <www.physics.nist.gov/PhysRefData>, which contains the Bibliography on Atomic Line Shapes and Shifts up to 2008 and the database at the Observatoire de Paris, <amrel.obspm.fr/balss> contains a Bibliography up to 2007.

The Vienna Atomic Line Database (VALD) can be found at <ams.astro.univie.ac.at/~vald> and the Belgrade database at <www.aob.bg.ac.yu/BELDATA>.

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