

# The Neutral Islands during the Late Epoch of Reionization

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**Abstract.** The large-scale structure of the ionization field during the epoch of reionization (EoR) can be modeled by the excursion set theory. While the growth of ionized regions during the early stage are described by the “bubble model”, the shrinking process of neutral regions after the percolation of the ionized region calls for an “island model”. An excursion set based analytical model and a semi-numerical code (`islandFAST`) have been developed. The ionizing background and the bubbles inside the islands are also included in the treatment. With two kinds of absorbers of ionizing photons, i.e. the large-scale under-dense neutral islands and the small-scale over-dense clumps, the ionizing background are self-consistently evolved in the model.

**Keywords.** cosmology: theory, intergalactic medium, large-scale structure of universe

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## 1. Introduction

Analytical modeling and approximated fast simulation are essential to explore the large parameter space and complicated history of reionization. As the large scale ionization field is closely correlated with the underlying density fluctuations (Battaglia *et al.* 2013), the excursion set theory (Bond *et al.* 1991; Lacey & Cole 1993) which give an intuitive description of the structure formation process are therefore also applied to the modeling of reionization process. In the “bubble model” (Furlanetto *et al.* 2004) the criterion of a region being ionized is translated to a density barrier, and the mass distribution of large ionized regions (“bubbles”) is obtained by computing the first up-crossing distribution of random trajectories with respect to this bubble barrier. Semi-numerical simulations of the reionization process were also developed, which is much faster than radiative transfer simulations (e.g. Zahn *et al.* 2007; Mesinger & Furlanetto 2007; Mesinger *et al.* 2011).

In the bubble model, the ionized regions during the early stage of EoR are assumed to be isolated and spherical. After the percolation of the HII regions, however, these basic assumptions are no longer valid, and the non-local source of ionizing photons must be accounted. In order to give a better description of the late stage of EoR, we proposed to consider the evolution of large-scale neutral regions (“islands”) which become isolated after percolation of ionized regions (Xu *et al.* 2014). The island model accounts for the ionizing background that is inevitable after percolation, and identifies neutral islands in the excursion set framework with first down-crossings of the island density barrier. Based on the island model, a semi-numerical code, `islandFAST`, is developed to make fast simulation of the evolution of the three-dimensional ionization field during the late EoR (Xu *et al.* 2017). It predicts the evolution of the ionization field as well as the ionizing background during the late stage of EoR. Here we present the basic idea and main results of the analytical and semi-numerical models.

## 2. The Island Model

In the island model, a region remains neutral if the available number of ionizing photons, either produced inside or outside, is not enough to ionize all hydrogen atoms in the region. The condition for island formation consists of two terms:

$$\xi f_{\text{coll}}(\delta_M; M, z) + \frac{\Omega_m}{\Omega_b} \frac{N_{\text{back}} m_{\text{H}}}{M X_{\text{H}} (1 + \bar{n}_{\text{rec}})} < 1, \quad (2.1)$$

where the first term on the L.H.S. accounts for the ionizing photons produced by collapsed objects inside, and the second term accounts for the contribution by the ionizing background around it.  $\xi$  is the ionizing efficiency parameter,  $f_{\text{coll}}$  is the collapsed fraction, and  $N_{\text{back}}$  is the number of consumed background ionizing photons. Assuming Gaussian density fluctuations, this condition can be written as a constraint on the density contrast of the region, i.e. the *island barrier*. Assuming spherical islands, and that the consumption rate of background ionizing photons is proportional to their surface area, we derive the basic island barrier shown with green curves in the left panel of Fig. 1. Different from the bubble barrier, the island barrier bends down at small scales because of the ionizing background. One can derive the mass distribution and the volume fraction of neutral regions by solving for the first-*down*-crossing distribution of random trajectories with respect to the island barrier.

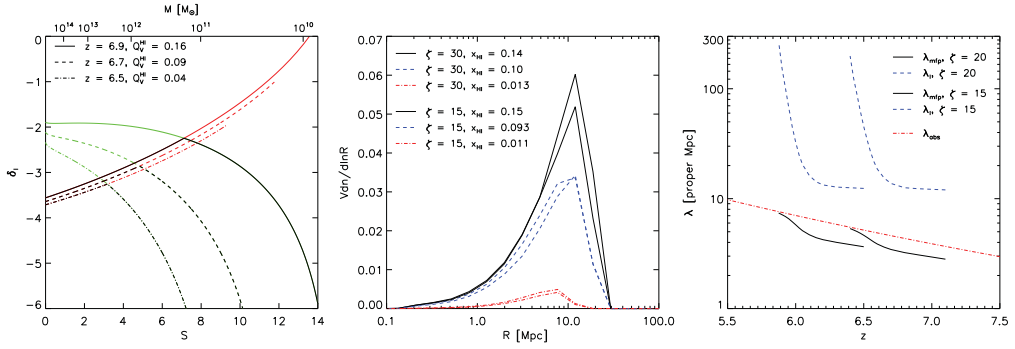
Note that there is a “bubbles-in-island” effect, as there might be self-ionized regions inside a host island. The distribution of bubbles inside islands are obtained by computing the conditional probability distribution of the trajectories. The average bubbles-in-island fraction can be derived by integrating over all possible bubble sizes for a given island, and this fraction is limited by a percolation threshold  $p_c$  in order to define the bona fide islands, in the sense that an island is not percolated through by bubbles inside. This percolation criterion of  $q_{\text{B}} < p_c$  acts as an additional barrier for finding islands (plotted with the red curves in the left panel of Fig. 1), which combines with the basic island barrier to define the host islands. The percolation threshold also demarcates the scope of application of both bubble model and island model. The bubble model can be reliably applied before bubble filling factor becomes larger than  $p_c$ , while the island model should be used below the redshift at which the island filling factor equals to  $p_c$ .

With the combined island barrier, the island model predicts the size distribution and evolution of islands after the islands become isolated. As the redshift decreases, small islands disappear rapidly while the large ones shrinks. Eventually, all these under-dense islands are ionized, only compact neutral regions such as galaxies or minihalos remain.

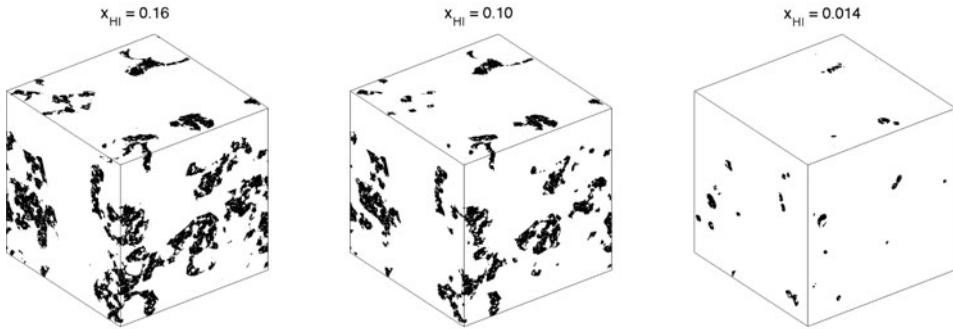
## 3. The IslandFAST Model

The semi-numerical simulation `islandFAST` is developed from the `21cmFAST` (Mesinger *et al.* 2011), but extended to treat the late EoR by the island model (see Xu *et al.* (2017) for details). Compared with the `21cmFAST`, the major differences of the `islandFAST` are the two-step filtering algorithm in generating the ionization field in order to take bubbles-in-island effect into account, and the self-consistent treatment for the ionizing background accounting for various absorption systems. The mean free path of ionizing photons and the corresponding intensity of the ionizing background is obtained by an iterative procedure. Adaptive redshift steps are applied in order to achieve convergence in the ionization field and the intensity of the ionizing background.

We find that the late reionization proceeds very fast, and the bubbles-in-island effect is quite obvious, showing complex morphology of the ionization field (Fig. 2). In the central panel of Fig. 1 we show the comoving size distributions of neutral islands for



**Figure 1.** *Left panel:* The basic island barriers (green curves), the percolation threshold induced barriers (red curves), and the effective barriers (black curves). The solid, dashed and dot-dashed curves are for  $z = 6.9, 6.7$  and  $6.5$  respectively. *Central panel:* The size distribution of neutral islands in the *islandFAST* simulation, with SAM neutral fraction threshold  $f_{\text{HI}}^c = 0.5$ . The thick and thin lines are for  $\zeta = 30$  and  $15$  respectively. The solid, dashed, and dot-dashed curves correspond to mean neutral fractions  $x_{\text{HI}}$  as indicated in the legend. *Right panel:* The evolution of the MFP of ionizing photons (*solid lines*). The dashed lines show the MFP due to neutral islands, while the dot-dashed line is the MFP due to small absorbers only. The thick lines are from the simulation with  $\zeta = 20$ , and the thin lines are for  $\zeta = 15$ .



**Figure 2.** The 3-dimensional visualization of the ionization fields from the *islandFAST* with a box size of  $200 h^{-1} \text{Mpc}$ ,  $512^3$  resolution, and  $\zeta = 20$ . The neutral islands are shown as black patches, and the ionized regions are left white. The three boxes have the mean neutral fractions of  $0.16, 0.10$ , and  $0.014$ , from left to right respectively.

various global neutral fraction. We find a characteristic scale of  $\sim 10$  Mpc for the neutral islands at each redshift. This scale decreases very slowly as the islands being ionized. Judging from the simulation box, this is perhaps because the neutral islands have very complex shape, and using the spherical average method (SAM) to identify islands tend to divide a large island into pieces. As the reionization proceeds, the larger islands gradually break into smaller ones that would be identified as similar sizes by the SAM. Also, the large islands shrink gradually, and partially compensates for the disappearance of the smaller islands.

While generating the ionization field, the *islandFAST* simultaneously predicts the level of the ionizing background. We find that a higher ionizing efficiency would lead to earlier growth and higher intensity of the ionizing background. The intensity of the ionizing background is boosted by an order of magnitude if we neglect small absorbers, resulting in a much faster completion of reionization. Similar trends are seen in the evolution of the mean free path (MFP) of the ionizing photons as shown in the right panel of Fig. 1. The MFP due to the under-dense islands alone,  $\lambda_{\text{I}}$ , and that limited by

over-dense absorbers,  $\lambda_{\text{abs}}$ , are also shown. We find that the MFP are mainly limited by the small-scale absorbers, and the shading effect of the large-scale islands only reduce the MFP moderately during the late EoR. Therefore, the level of the ionizing background is primarily regulated by small-scale absorbers, and they significantly delay and prolong the reionization process.

#### 4. Summary

We have developed the island model—an analytical model of reionization using the excursion set approach, and the corresponding semi-numerical simulation, `islandFAST`, for the evolution of isolated neutral regions after the percolation of ionized bubbles, complementary to the bubble model for the early reionization epoch. As the case of bubble model, although the excursion set model uses spherical window function for computing the mean density at different scales, the model can and does produce complicated reionization topology. Our model incorporated the effect of ionizing background photons and accounted for the bubbles-in-island effect. A percolation threshold is introduced to demarcate the scope of application for the island model and the bubble model, and also to define the bona fide islands. Given a model for the small-scale absorbers, the `islandFAST` generates the ionization field after percolation, and simultaneously predicts the intensity of the ionizing background.

The late reionization proceeds quite fast. As the large islands shrink with time, the small ones are swamped by the ionizing photons. If characterized by the SAM with a high neutral fraction threshold, the islands exhibits a relatively robust characteristic scale of  $\sim 10$  Mpc throughout the late EoR. We find that the large-scale islands dominate the morphology of the ionization field, while the small scale absorbers dominate the opacity of the IGM, and play a major rule in determining the level of the ionizing background and the speed of reionization process.

With the advent of next generation EoR experiment, reionization models will be put to rigorous tests, and the late state is of particular interest as it might be observationally more accessible. We expect these models would be helpful for interpreting such observations.

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