

Vector Solution for the Intersection of Two Circles of Equal Altitude

Andrés Ruiz González

(*Navigational Algorithms. San Sebastián.*)

(Email: aruiz4@euskalnet.net)

A direct method for obtaining the two possible positions derived from two sights using vector analysis instead of spherical trigonometry is presented. The geometry of the circle of equal altitude and of the two body fixes is analyzed, and the vector equation for simultaneous sights is constructed. The running fix problem is also treated.

KEY WORDS

1. Circle of Equal Altitude. 2. Celestial Navigation. 3. Sight Reduction. 4. Vector Analysis.

1. INTRODUCTION. Most of the known methods to solve the two body problem, [10], [13–17], [19], [24–26], use the navigational triangle [1] and spherical trigonometry. Recursive methods using linear approximations as in [18], or physics analogies [12] are smart approaches. Sight reduction general procedures [2–5], can be used to obtain the fix, but for two sights the estimated position is needed. In reference [11] an analytical solution using plane geometry is used for both solutions.

On the other hand, vector calculus is a powerful tool widely used for engineering and physics, [9]. In navigation it is used in naval kinematics, for current calculations, (set & drift) and the great circle sailing can be formulated using the vector analysis, [20–23]; in [27] a vector approach of celestial navigation is made. In this paper a solution for the two body problem in celestial navigation using this technique is presented. The following sub paragraphs present some concepts about the celestial circle of position, *CoP*.

1.1. *Variables and symbols.* Let us denote the dot product by \bullet and the cross product by Λ , and a vector by \vec{v} or by \mathbf{v} . The variables used are:

	Variable	Intervals
GHA	Greenwich Hour Angle	$0 \leq \text{GHA} \leq 360^\circ$ (W to E)
Dec	Declination	-90° (S) $\leq \text{Dec} \leq +90^\circ$ (N)
Ho	Observed altitude	$0 \leq H \leq 90^\circ$
B	Latitude	-90° (S) $\leq B \leq +90^\circ$ (N)
L	Longitude	-180° (W) $\leq L \leq +180^\circ$ (E)

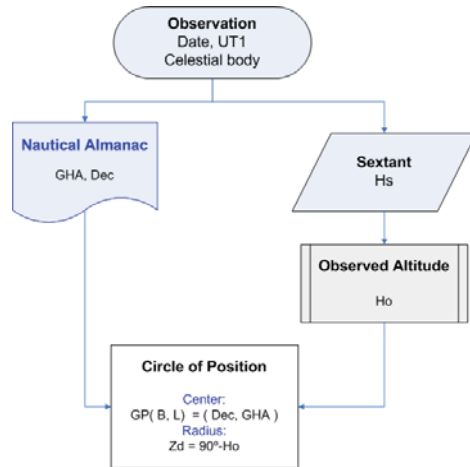


Figure 1. Circle of Equal Altitude parameters.

1.2. *The Circle of Equal Altitude.* A celestial object is far enough away from the observer that the incoming light rays are nearly parallel to each other. Thus, there is a point on the surface of Earth where the object is directly overhead at a given time; this point is called the *geographical position, GP, (or of a star; substellar point)*. Using the spherical model of the Earth, there is a circle on her surface centred about the object’s geographical position where the angle between the horizon and the celestial object, called the altitude, is constant at a given instant. This circumference forms a celestial line of position, a small circle, known as a circle of equal altitude. The great circle distance from this pole to the circle is the zenith distance of the body, *Zd*.

At the time of observation the observer of the celestial object must be located somewhere along that circle. The geographical position of a celestial body is calculated from an ephemeris or obtained from the *Nautical Almanac*, and the altitude is measured with a sextant. The observed altitude is the sextant altitude corrected for index error and dip, for refraction and if appropriate corrected for parallax and semi-diameter. This process is summarized in Figure 1.

Geometrically a Circle of Equal Altitude is generated by the intersection of a circular cone having its vertex in the centre of the Earth, half angle $\alpha = 90^\circ - Ho$, and with the vector from O to GP as its axis, with the unit sphere. The distance from the centre of the Earth to the plane containing the CoP is $\sin(Ho)$, as shown in Figure 2.

1.3. *Vector equation for the Circle of Equal Altitude.* In Figure 3, let OP be the observer’s position at the time of sight, and GP the geographical position of the celestial body at the same instant. The dot product of the vectors defined by the centre of the Earth and these points is the cosine of the angle between them, which is the zenith distance of the observed body. Then, the vector equation of the circle of equal altitude is:

$$\vec{OP} \cdot \vec{GP} = \cos(90^\circ - Ho) \tag{1}$$

It is possible to write the azimuth in vector form [8].



Figure 2. Earth's normal section to the plane of the Circle of Position (CoP).

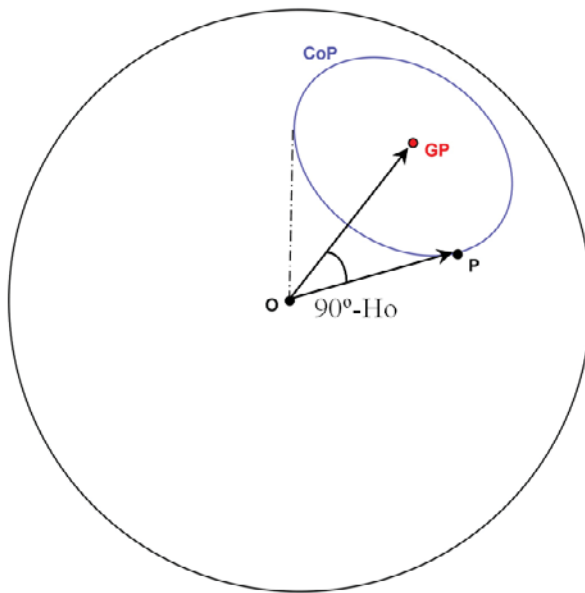


Figure 3. Circle of Equal Altitude and vectors.

1.4. *Coordinate Systems.* Using a right-handed orthonormal basis $\{\vec{i}, \vec{j}, \vec{k}\}$, as in Figure 4, the Cartesian system of coordinates is defined where the origin **O**, is the centre of the Earth with Axes:

- Z: from O to the North Pole.
- X: from O to the Greenwich meridian, included in the Earth's equatorial plane.
- Y: defined by $\vec{j} = \vec{k} \wedge \vec{i}$

Since angles are used, not distances, the hypothesis that the Earth is a sphere of unit radius is valid.

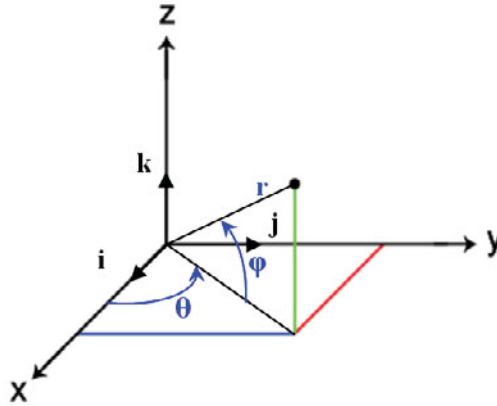


Figure 4. Right-handed orthonormal basis, and coordinates.

The relationship between the equatorial coordinates (Dec, GHA), and geographical coordinates (B, L), with the spherical ones, (φ, θ), arises from Figure 4:

$$\begin{aligned} r &= 1 & r &= 1 \\ \text{Dec} &= \varphi & B &= \varphi \\ \text{GHA} &= 360^\circ - \theta & L &= \theta \end{aligned}$$

According to this formulation the unit vector in Cartesian coordinates, (x,y,z), from the centre of the Earth to the geographical position of any body is:

$$\vec{GP} = \cos \text{Dec} \cos \text{GHA} \cdot \vec{i} - \cos \text{Dec} \sin \text{GHA} \cdot \vec{j} + \sin \text{Dec} \cdot \vec{k} \tag{2}$$

And the unit vector in Cartesian coordinates from the centre of the Earth to any point on the surface of the Earth is:

$$\vec{OP} = \cos B \cos L \cdot \vec{i} + \cos B \sin L \cdot \vec{j} + \sin B \cdot \vec{k} \tag{3}$$

2. VECTOR EQUATION FOR THE INTERSECTION OF TWO SIMULTANEOUS CIRCLES OF POSITION. In the general case, two CoP intersect at two points: I_1 and I_2 , (Figure 5). The coordinates of these two crossings are the solution to the problem. Using the vector notation in section 1.3 there are three unknown variables, the Cartesian coordinates of OP ; so three equations are needed:

$$\begin{aligned} \vec{OP} \cdot \vec{GP}_1 &= \sin Ho_1 \\ \vec{OP} \cdot \vec{GP}_2 &= \sin Ho_2 \\ \vec{OP} \cdot \vec{OP} &= 1 \end{aligned} \tag{4}$$

The last equation takes account of the fact that the observer is on the surface of a unit sphere: $x^2 + y^2 + z^2 = 1$. To solve the system some methods have been published [4], [11].

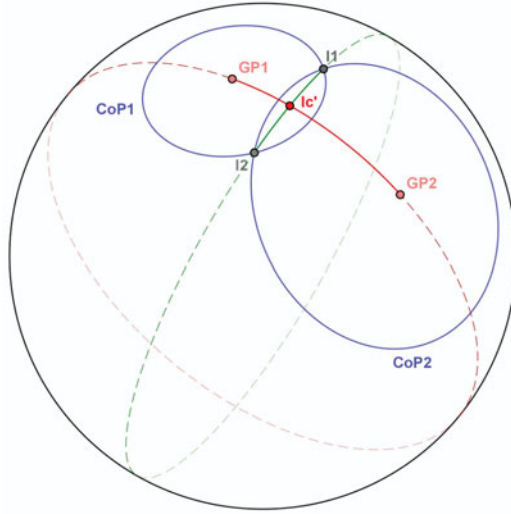


Figure 5. Crossing of two CoP.

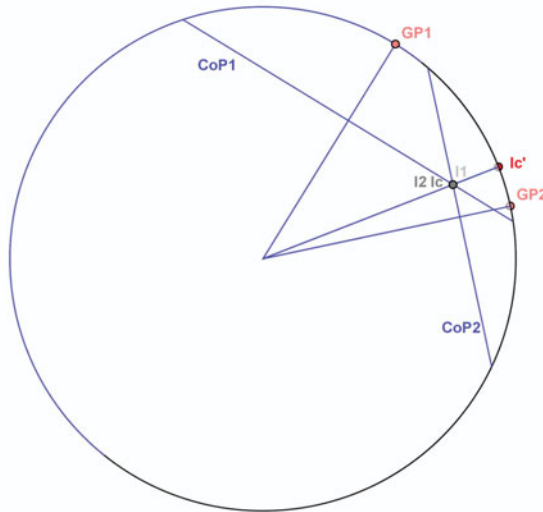
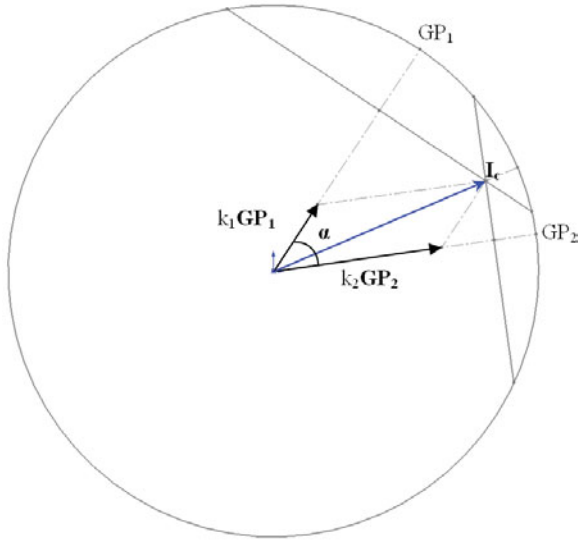


Figure 6. Section by the plane containing the two geographical positions and the centre of the Earth.

Another approach arises considering Figures 5 & 6. The great circle passing through the two geographical positions is perpendicular to the one defined by the points I_1 and I_2 , and intersect in a point I'_c on the surface of the sphere. Each CoP is contained in a plane, (section 1.2), the intersection of these planes is the line defined by the points I_1 and I_2 . The line OI'_c and the line I_1I_2 have a point in common: I_c . The vector from the centre of the Earth to each GP in rectangular coordinates is:

$$\vec{GP}_j = (\cos Dec_j \cos GHA_j - \cos Dec_j \sin GHA_j \sin Dec_j) \tag{5}$$



$$\begin{aligned}
 k_1 \cos \alpha &= \sin Ho_2 - k_2 \\
 k_2 \cos \alpha &= \sin Ho_1 - k_1 \\
 \text{then:} \\
 k_1 &= \frac{\sin Ho_1 - \sin Ho_2 \cos \alpha}{\sin^2 \alpha} \\
 k_2 &= \frac{\sin Ho_2 - \sin Ho_1 \cos \alpha}{\sin^2 \alpha}
 \end{aligned}$$

Figure 7. Vector to the point I_c .

for the two crossings $j=1,2$. This vector is perpendicular to the plane that contains the CoP.

The vector from the origin to the point I_c can be written like a linear combination of the GP vectors:

$$\vec{OI_c} = k_1 \vec{GP_1} + k_2 \vec{GP_2} \tag{6}$$

Where k_1 and k_2 are the scale factors derived from Figure 7, and alpha is the angle between the two GP vectors:

$$\vec{GP_1} \cdot \vec{GP_2} = \cos \alpha \tag{7}$$

The direction of the line connecting the two intersection points: I_1 and I_2 , is defined by the unit vector

$$\vec{r} = \frac{\vec{GP_1} \wedge \vec{GP_2}}{\|\vec{GP_1} \wedge \vec{GP_2}\|} \tag{8}$$

perpendicular to the great circle passing through the two geographical positions. The vectors from I_c to the points solution of the problem are:

$$\vec{I_c I_1} = L_1 \vec{r} \text{ and } \vec{I_c I_2} = L_2 \vec{r}.$$

From Figure 8, L_1 and L_2 are distances from the middle point of the $I_1 I_2$ line to each intersection point.

$$L_j = I_c I_j \tag{9}$$

$$OI_c^2 + I_c I_j^2 = OI_j^2 = 1 \tag{10}$$

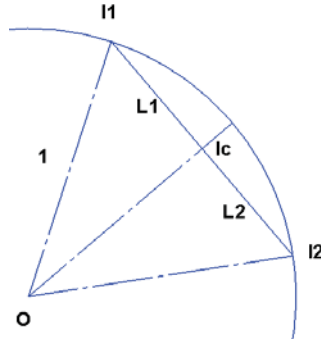


Figure 8. Section by the plane containing the line I1I2 and O.

where $j = 1, 2$.

OI_c is obtained using the dot product:

$$\vec{OI}_c \cdot \vec{OI}_c = OI_c^2 \tag{11}$$

$$L_j = \pm \sqrt{1 - \vec{OI}_c \cdot \vec{OI}_c} \tag{12}$$

The vector from the origin to each intersection is:

$$\vec{OI}_j = \vec{OI}_c + \vec{IcI}_j \tag{13}$$

And finally, the vector equation for the solution points is:

$$\vec{OI}_j = k_1 \vec{GP}_1 + k_2 \vec{GP}_2 + L_j \frac{\vec{GP}_1 \wedge \vec{GP}_2}{\|\vec{GP}_1 \wedge \vec{GP}_2\|} \tag{14}$$

for $j = 1, 2$.

Transforming the Cartesian coordinates, $\vec{OI}_j = (x_j \ y_j \ z_j)$, in to spherical ones, the latitude and the longitude of the two points of intersection will be obtained.

$$\begin{aligned} B_j &= ATAN2(z_j, \sqrt{x_j^2 + y_j^2}) \\ L_j &= ATAN2(y_j, x_j) \end{aligned} \tag{15}$$

For the North Pole (0,0,1), and the South Pole (0,0,-1), the solution is undetermined. The algorithm is shown in a flow chart in the Appendix.

2.1. *The Fix.* Only one of the two points of intersection obtained is the true position of the observer. The dead reckoning position determines the fix.

2.2. *Singular case.* There is a theoretical case very improbable in navigation: if the two CoPs are tangents, there is only one solution, but the vector equation gives the correct result. The two points of intersection I₁ and I₂ in Figures 5 and 6, the middle point Ic and its projection on the surface of the Earth Ic' are the same. The line I₁I₂ degenerates to a point, as well as the two crossings, and L_j = 0.

$$I_1 \equiv I_2 \equiv Ic \equiv I'_c \quad IcI_j = 0 \quad OI_c = OI_j = 1 \text{ and } L_j = 0$$

Then the vector equation for the position is:

$$\vec{\mathbf{O}}\mathbf{I}_j = k_1 \vec{\mathbf{G}}\mathbf{P}_1 + k_2 \vec{\mathbf{G}}\mathbf{P}_2 \quad \text{for } j: 1 \equiv 2$$

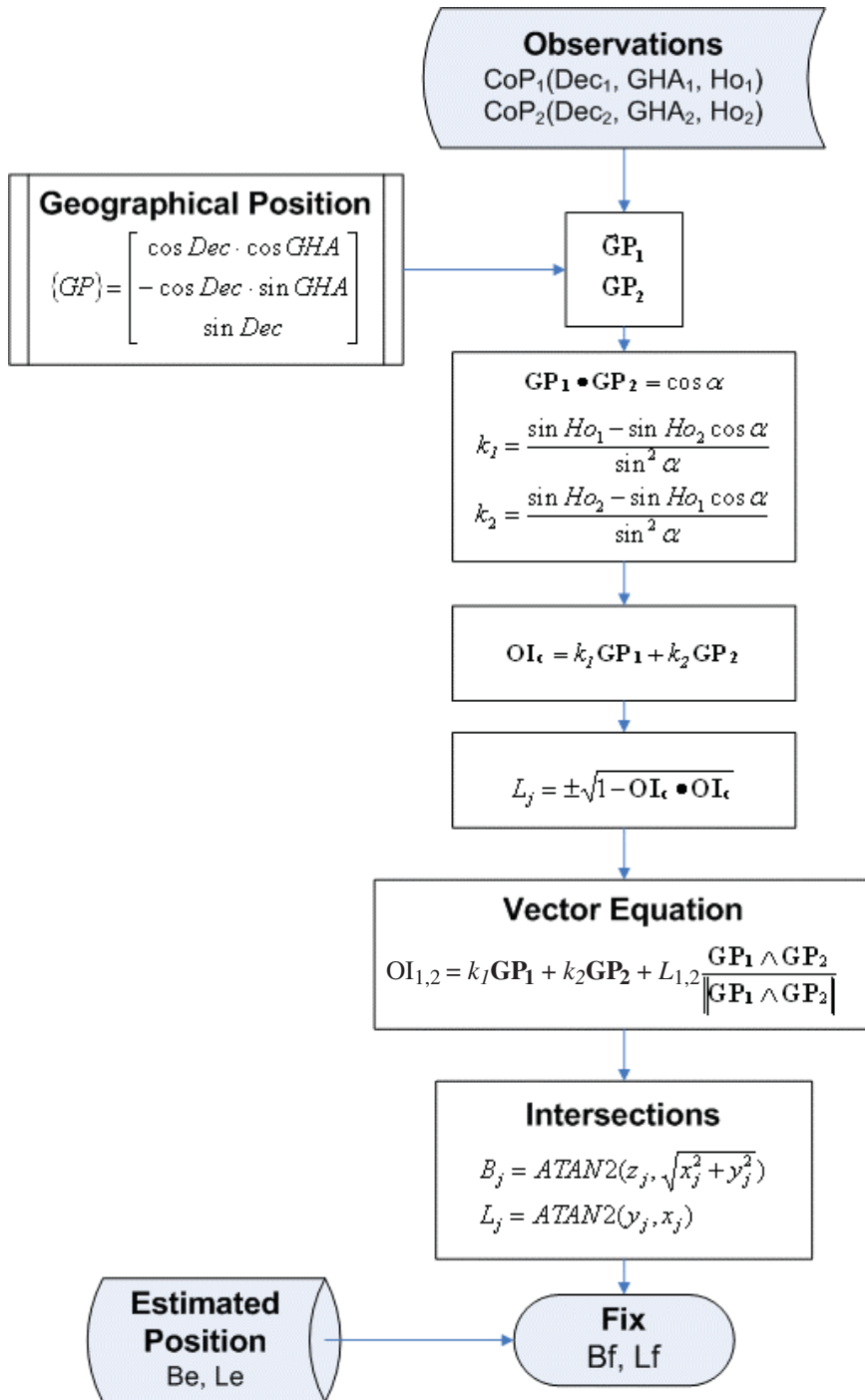
2.3. *Correction for the motion of the observer.* When the two sights are not taken at the same time, it is necessary to move the first CoP to the time of the second one, or both to a common instant. Many methods have been presented, but the correct way to move a CoP is to advance or retire the GP due to the motion of the observer [6], [7].

Using the technique described in reference [6], the correction is a function of the estimated position of the observer, (Be,Le), and his motion: course and speed, (C,S). And since what we are looking for is the true position, an iterative process is required in order to reach the solution for the running fix. The algorithm is described in the Appendix.

3. **CONCLUSION.** A vector solution in rectangular coordinates for the two body fix problem has been presented. The development uses only vector algebra and coordinate transformation between geographical and Cartesian coordinates avoiding the use of the spherical trigonometry. The schema for deriving the equation is clear and intuitive. The equation is a compact function of the geographical positions, and the scale factors are also functions of the observed altitudes.

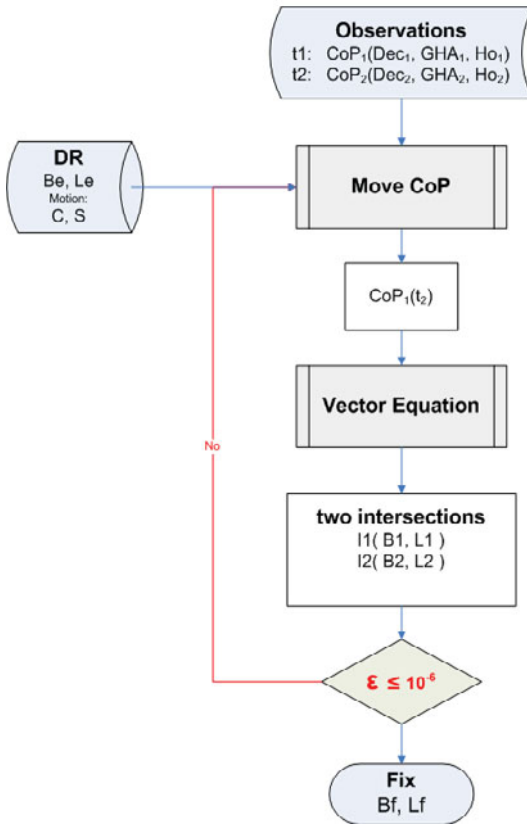
APPENDIX

A1. Fix Algorithm.



A2. Running Fix Algorithm.

Vector Solution for the intersection of two Circles of Equal Altitude. Running Fix



An implementation of these two algorithms is available at the author's web site: *CelestialFix.exe* running under Windows. Other implementations are very easy in a calculator, electronic spreadsheet, or math dedicated software with vector analysis functions performed.

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