

Bremsstrahlung soft X-ray emission from clusters heated by a Gaussian laser beam

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Abstract

A theoretical model of soft X-ray emission from laser irradiated clusters is developed. An intense short pulse laser of Gaussian radial and temporal profiles impinged on a clustered gas jet, heats the cluster electrons, leading to Bremsstrahlung emission of X-rays. As the clusters expand under hydrodynamic pressure, plasma frequency of the cluster electrons ω_{pe} decreases. When plasma frequency of a cluster approaches plasma resonance $\omega_{pe} = \sqrt{3}\omega$ (where ω is the laser frequency), the electrons are resonantly heated by the laser and a rapid rise in X-ray emission occurs. After a while, when cluster expansion detunes the plasma resonance, X-ray emission falls off.

Keywords: Atomic cluster; Bremsstrahlung; Laser; X-ray emission

1. INTRODUCTION

The interaction of intense short pulse laser with atomic clusters (Krainov & Smirnov, 2002; Ditmire *et al.*, 1998; 1996; 1997; 1999; Milchberg *et al.*, 2001; Shokri *et al.*, 2004; Mishra *et al.*, 2011; Kanapathipillai *et al.*, 2004, Kumar *et al.*, 2013; Mulser *et al.*, 2005) has become an important area of research with wide ranging applications from harmonic generation (Tiwari & Tripathi, 2006; Shim *et al.*, 2007; Huillier & Balcou, 1993; Para *et al.*, 2000) and X-ray generation (Shao *et al.*, 1996; Liu *et al.*, 2009; Fukuda *et al.*, 2004; Jha *et al.*, 2008; Kumarappan *et al.*, 2001) to energetic neutron production (Ditmire *et al.*, 1995; Hegelich *et al.*, 2006). The clusters (Hagen & Obert, 1972) are the Vander-walls bonded assemblies of about 10^2 – 10^7 atoms that are formed during rapid cooling of supersonic gas flowing out of a nozzle. The cluster can be considered as small spherical balls of nanometer dimension, whose radius is much smaller than laser wavelength. The main characteristic of a gaseous cluster is the near solid density inside the cluster that results in an enhanced absorption of laser energy via collisional processes and consequently in an increase of X-ray conversion efficiency. This results in efficient generation of incoherent short wavelength light for application such as extreme

ultraviolet (EUV) lithography, EUV and X-ray microscopy, and X-ray tomography.

The X-ray emission spectrum (Dorchies *et al.*, 2008; Mocek *et al.*, 2002; Deiss *et al.*, 2006; Smirnov & Becker, 2006; Kumar & Tripathi, 2009; Issac *et al.*, 2003; Sailja *et al.*, 2005) is obtained from clusters irradiated by a femto-second laser pulse. One can see a continuous Bremsstrahlung emission, and three different spectral structures identified as the K-shell emission, corresponding to transitions from the L, M, N-shells down to the K-shell. The nano-plasma model proposed first by Ditmire *et al.* (1996) and improved by Milchberg *et al.* (2001) shows, that the laser absorption is enhanced during the cluster expansion when the electron density approaches thrice the critical density $\omega = \omega_{pe}/\sqrt{3}$. Ditmire *et al.* (1995) have reported that the enhanced absorption of laser by the clusters results in the production of high ion charge states via collisional ionization giving strong X-ray emission from the hot plasma. Dorchies *et al.* (2008) have reported the experimental evidence of sub-picosecond X-ray bursts produced by laser-cluster interaction in the multi-KeV range. The theoretical understanding of X-ray emission from clustered gases in these studies is primarily based on uniform laser intensity model. Kumar and Tripathi (2009) have recently developed a theory of Rayleigh scattering of an intense laser, with Gaussian temporal and radial profiles, from expanding clusters.

In this paper, we develop a theoretical model of cluster heating and X-ray emission by intense short pulse Gaussian

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laser beam. The physics of the problem is as follows. The laser quickly ionizes all the atoms of the clusters converting them into plasma balls and also heats the electrons inside them. The heating is faster on the laser axis and slower as one moves radially away. The heated clusters expand on sub-picosecond time scale and their electron density decreases. As the plasma frequency approaches plasma resonance, $\omega_{pe} = \sqrt{3}\omega$, electron temperature rises sharply. The free electrons get accelerated and emit radiation in the X-ray regime, when they pass through the vicinity of positive ions (Fig. 1). The sharp rise in X-ray emission rate is first expected from clusters in the axial region. However, when the cluster electron plasma frequency falls below the plasma resonance, the emission rate should slow down due to reduction in the rate of Bremsstrahlung encounters. In Section 2, we calculate laser induced cluster heating and expansion. In Section 3, we study the Bremsstrahlung X-ray emission. In Section 4, we discuss the results.

2. CLUSTER HEATING AND EXPANSION

Consider a gas jet target embedded with clusters of initial radius r_{c0} and density n_c . Each cluster has free electron density n_{e0} inside it. A Gaussian laser beam is launched into it with electric field

$$\vec{E}_L = \hat{x}A_L(z, t) \exp(-i(\omega t - kz)), \quad (1)$$

where $A_L^2 = A_{L0}^2 \exp(-r^2/r_0^2) \exp(-(t - z/c)^2/\tau_L^2)$ and magnetic field $\vec{B} = (\vec{k} \times \vec{E}/\omega)$. We ignore the effects of diffraction and nonlinear refraction. The laser quickly ionizes all the atoms of the clusters converting them into plasma balls with overlapping electron and ion spheres. The electron sphere oscillates in the \hat{x} direction due to the laser field. At any instant, the center of electron cloud is shifted by a distance Δ with respect to the center of the ion sphere. The equation of motion for the cluster electrons is

$$m \frac{d^2 \Delta}{dt^2} = -\frac{\omega_{pe}^2 m \Delta}{3} - eE_L - m\nu \frac{d\Delta}{dt}, \quad (2)$$

where $\omega_{pe} = (4\pi n_e e^2/m)^{1/2}$ is the cluster electron plasma frequency and $-e$, m , ν , and n_e are the electronic charge, mass, electron-ion collision frequency, and modified electron density (due to cluster expansion). The oscillatory velocity of the cluster electrons may be written as

$$\vec{v}_{osc} = \frac{d\vec{\Delta}}{dt} = \frac{e\omega \vec{E}_L}{mi(\omega^2 - \omega_{pe}^2/3 + i\nu\omega)}. \quad (3)$$

The laser heats the electrons at an average rate

$$\frac{d}{dt} [3T_e/2] = -(e/2)\vec{E}_L^* \cdot \vec{v}_{osc},$$

where the energy loss rate via collisions has been ignored which, is valid when pulse duration is shorter than the energy relaxation time. Following Kumar and Tripathi (2009), we take $\nu = \nu_0 (T_e/T_0)^{-3/2}$ and define $\xi = T_e/T_0$, $\tau = \nu_0 t'$, $t' = (t - z/c)$, and write the temperature equation as,

$$\frac{d}{d\tau} [\xi] = \frac{2}{5} \frac{a^2 \exp(-\tau^2/\nu_0^2 \tau_L^2)}{\left[\frac{1}{b_1 \xi^{3/2}} + \xi^{3/2} \left(1 - \frac{d}{3\eta^3}\right)^2 \right]}. \quad (4)$$

where $a^2 = a_0^2 \exp(-r^2/r_0^2)$, $a_0^2 = e^2 A_{L0}^2 / 3m\omega^2 T_0$, $b_1 = \omega^2/\nu_0^2$, $d = \omega_{pe0}^2/\omega^2$, $\eta = r_c/r_{c0}$, r_c is the modified cluster radius (due to expansion) and T_0 is the normalizing temperature. If one takes the expansion of cluster to be adiabatic, one must add a cooling term to Eq. (4). For adiabatic expansion in the absence of heating, $T_e V^{\gamma-1} = \text{const.}$, where $V = 4\pi r_c^3/3$ is the volume of the cluster and γ is the ratio of specific heats, one may write,

$$\frac{dT_e}{dt} = -\frac{T_e(\gamma-1)dV}{V dt} = -\frac{3T_e(\gamma-1)dr_c}{r_c dt}. \quad (5)$$

For $\gamma = 5/3$. The rate of temperature decrease due to the cluster expansion is

$$\frac{dT_e}{dt} = -2 \frac{T_e dr_c}{r_c dt}. \quad (6)$$

With this cooling rate Eq. (4) can be written as,

$$\frac{d}{d\tau} [\xi] = \frac{2}{5} \frac{a^2 \exp(-\tau^2/\nu_0^2 \tau_L^2)}{\left[\frac{1}{b_1 \xi^{3/2}} + \xi^{3/2} \left(1 - \frac{d}{3\eta^3}\right)^2 \right]} - 2 \frac{\xi}{\eta} \frac{d}{d\tau} (\eta). \quad (7)$$

As a cluster gets heated, its radius r_c expands with sound velocity $(dr_c/dt) = (T_e/m_i)^{1/2}$,

where m_i is the mass of the ion. Defining $v_{s0} = (T_0/m_i)^{1/2}$, $\eta = r_c/r_{c0}$, $g = v_{s0}/r_{c0}\nu_0$, we may write

$$\frac{d}{d\tau} (\eta) = (\xi)^{1/2} g. \quad (8)$$

As the radius increases, plasma frequency decreases, keeping $\omega_{pe}^2 r_c^3 = \text{constant}$, i.e., $\omega_{pe}^2 r_c^3 = \omega_{pe0}^2 r_{c0}^3$, where ω_{pe0} is the value of ω_{pe} at $t = 0$. Thus, $\omega_{pe}^2 = \omega_{pe0}^2 / \eta^3$. Eqs. (7)–(8) are coupled in ξ and η and can be solved numerically. One may note (cf. Kumar & Tripathi, 2009) that the heating rate rises sharply when the electron density in the cluster approaches three times that of the critical density, i.e., $\omega_{pe} = \sqrt{3}\omega$. The timing of resonance would depend on the radial location of the cluster besides its size and ion mass. In Figures 2 and 3, we have plotted the variation of electron temperature and cluster radius with time for clusters located at $r/r_0 = 0.1, 1$ when the parameters are: $a_0 = 1$, $b = 10^5$, $d = 5$, $g = 0.3$ and $\nu_0^2 \tau_L^2 = 1$. One may note that initially the cluster radius expands slowly. As the

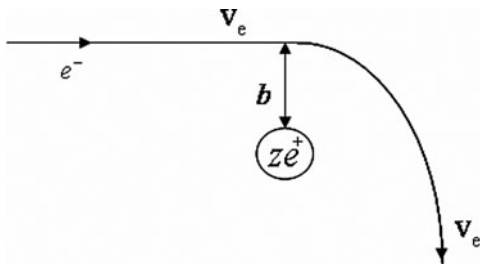


Fig. 1. Schematic of the Bremsstrahlung process.

plasma resonance is reached it rises sharply. The clusters near the axis expand more quickly and plasma resonance occurs earlier in time. For the off axis clusters, the laser field is weak and clusters expand slowly, taking longer time to reach the resonance.

3. BREMSSTRAHLUNG X-RAY EMISSION

The free electrons of a cluster when pass through the vicinity of ions with initial velocity v_e and impact parameter b get accelerated and emit radiation in the X-ray regime. The energy radiated by the electron, using the Larmor's formula (Jackson, 1962; Liu & Tripathi, 1994; Landau & Lifshitz, 1975) is

$$U = \frac{2e^2 a_c^2 \tau_c}{3c^3} = \frac{4Ze^6}{3m^2 c^3 b^3 v_e}, \tag{9}$$

where $\tau_c = 2b/v_e$ is the collision time of the electron with the ion, $a_c = Ze^2/mb^2$ is the maximum acceleration suffered by the electrons, Z is the ion charge state, and c is the speed of light. In the Bremsstrahlung process (Fig. 1) the acceleration is the function of time. For the frequency spectrum of the emitted radiation, it is necessary to Fourier analysis of the acceleration

$$a_c(\omega) = \frac{1}{\pi} \int_{-\infty}^{\infty} a_c(t) e^{-i\omega t} dt,$$

$$a_c(t) = \int_0^{\infty} a_c(\omega) e^{i\omega t} d\omega.$$

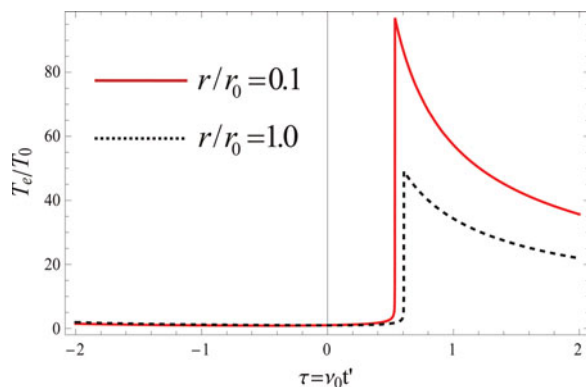


Fig. 2. (Color online) Normalized temperature of cluster electrons (T_e/T_0) as a function of normalized time ($\tau = \nu_0 t'$) for $a_0 = 1.0$ at radial locations.

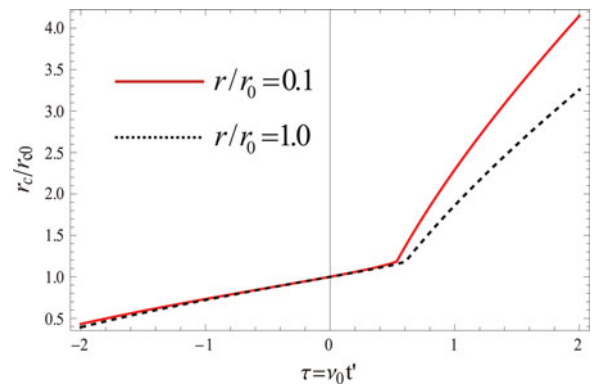


Fig. 3. (Color online) Normalized cluster radius (r_c/r_{c0}) as a function of normalized time ($\tau = \nu_0 t'$) for the different values of radial locations at normalized axial laser amplitude $a_0 = 1.0$.

Using the Parseval's formula

$$\int_{-\infty}^{\infty} [a_c(t)]^2 dt = \pi \int_0^{\infty} |a_c(\omega)|^2 d\omega.$$

The energy emitted as the electron passes by the ion is

$$U(\omega) = \int_{-\infty}^{\infty} P(t) dt = \frac{2e^2}{3c^3} \int_{-\infty}^{\infty} [a_c(t)]^2 dt = \frac{2\pi e^2}{3c^3} \int_0^{\infty} |a_c(\omega)|^2 d\omega. \tag{10}$$

The two components of acceleration are;

$$a_{cx}(t) = \frac{Ze^2}{m} \frac{v_e t}{(b^2 - (v_e t)^2)^{3/2}},$$

$$a_{cz}(t) = \frac{Ze^2}{m} \frac{b}{(b^2 - (v_e t)^2)^{3/2}}.$$

Their Fourier Transforms are:

$$a_{cx}(\omega) = \frac{1}{\pi} \int_{-\infty}^{\infty} a_{cx}(t) e^{-i\omega t} dt = \frac{Ze^2}{\pi m} \frac{2\omega}{v_e^2} iK_0(\omega b/v_e),$$

$$a_{cz}(\omega) = \frac{1}{\pi} \int_{-\infty}^{\infty} a_{cz}(t) e^{-i\omega t} dt = \frac{Ze^2}{\pi m} \frac{2\omega}{v_e^2} K_1(\omega b/v_e),$$

where $K_0(\omega b/v_e)$ and $K_1(\omega b/v_e)$ are the modified Bessel functions of argument $(\omega b/v_e)$. Using these expressions in Eq. (10) we get

$$U(\omega) \simeq \left. \begin{aligned} &\frac{8Z^2 e^6}{3\pi m^2 c^3} \frac{1}{b^2 v_e^2} && \text{for } \frac{\omega b}{v_e} \ll 1 \\ &\frac{8Z^2 e^6}{3\pi m^2 c^3} \frac{1}{b^2 v_e^2} e^{-(\omega b/v_e)} && \text{for } \frac{\omega b}{v_e} \gg 1 \end{aligned} \right\}. \tag{11}$$

The number of electron ion collisions per unit time with impact parameter ranging from b to $b + db$ is $n_i v_e 2\pi b db$.

Hence, the power radiated per electron is

$$P(\omega) = \int_{b_{\min}}^{b_{\max}} U(\omega)v_e n_i 2\pi b db \simeq \frac{16Z^2 n_i e^6}{3m^2 c^3 v_e} \ln \left[\frac{b_{\max}}{b_{\min}} \right], \quad (12)$$

where $b_{\max} \approx (v_e / \omega)$ and $b_{\min} = (\hbar / mv_e)$. The contribution to the frequency comes ω only from those electrons that have velocity $v_e > (2\hbar \omega/m)^{1/2}$. If we consider the electron velocity distribution function to be Maxwellian,

$$f_e = n_e \left(\frac{m}{2\pi T_e} \right)^{3/2} \exp \left(-\frac{mv_e^2}{2T_e} \right),$$

The total power radiated per unit frequency interval per unit volume is

$$P_B(\omega) = \int_{(2\hbar\omega/m)^{1/2}}^{\infty} P f_e 4\pi v_e^2 dv_e \simeq \int_{(2\hbar\omega/m)^{1/2}}^{\infty} \frac{64\pi Z^2 n_i e^6}{3m^2 c^3} \times v_e n_e \left(\frac{m}{2\pi T_e} \right)^{3/2} \exp(-mv_e^2/2T_e) dv_e. \quad (13)$$

After solving the above equation we get

$$P_B(\omega) = \frac{32Z^2 n_i n_e e^6}{3m^2 c^3 v_{th}} \left(\frac{T_e}{T_0} \right)^{-1/2} \ln \left(\frac{b_{\max}}{b_{\min}} \right) \exp \left(-\frac{\hbar\omega/T_0}{T_e/T_0} \right), \quad (14)$$

where $v_{th} = (2T_0/m)^{1/2}$ is the thermal velocity of electrons and n_e is the density of electrons. The total power radiated per unit frequency interval due to the clusters in plasma channel of length L is

$$P_B^T(\omega) = \int_0^{\infty} \int_0^L \frac{4\pi r_c^3 n_c}{3} P_B 2\pi r dr dz, \quad (15)$$

For deuterium cluster $Z = 1$, $n_e = n_i$, Eq. (15) gives

$$P_B^T(\omega) = \frac{16Z^2 \omega_{pe0}^3 L r_0^2 n_c r_c^3 c \omega_{pe0}^3 m}{9\pi^{3/2} c^3 v_{th} n_{e0} c r_0} \int_0^{\infty} \left(\frac{T_e}{T_0} \right)^{-1/2} \times \left(\frac{\omega_{pe}}{\omega_{pe0}} \right)^6 \left(\frac{r}{r_0} \right) \exp \left(-\frac{\hbar\omega/T_0}{T_e/T_0} \right) dr. \quad (16)$$

We may normalize Eq. (16) by the factor $(m\omega_{pe0}^3/n_{e0}c)$, then we get

$$\frac{P_B^T}{\omega_{pe0}^3 m/n_{e0}c} = \frac{16Z^2 \omega_{pe0}^3 L r_0^2 n_c r_c^3 c}{9\pi^{3/2} c^3 v_{th}} I, \quad (17)$$

where

$$I = \frac{1}{r_0} \int_0^{\infty} \left(\frac{T_e}{T_0} \right)^{-1/2} \left(\frac{\omega_{pe}}{\omega_{pe0}} \right)^6 \exp \left(-\frac{\hbar\omega/T_0}{T_e/T_0} \right) \left(\frac{r}{r_0} \right) dr.$$

The integrand in I has explicit dependence on r/r_0 . Implicit

dependence on r/r_0 comes through cluster electron temperature and cluster electron plasma frequency. We have solved Eq. (17) numerically for the following parameters: $a_0 = 1$, $r_{c0} = 10^{-6}$ cm, $n_c = 10^{13}/\text{cm}^3$, $r_0 = 10^{-2}$ cm, $v_{th} = 10^7$ cm/sec, $\omega_{pe0} = 5 \times 10^{15}$ rad/sec, $L = 0.01$ cm, $c = 3 \times 10^{10}$ cm/sec, $\hbar \omega/T_0 = 10$. Figure 4 shows the variation of the normalized total X-ray emission by the clusters $(P_B^T / (m\omega_{pe0}^3 / n_{e0} c))$ as a function of normalized time $(\tau = v_0 t')$. It shows rapid rise in X-ray emission at $\tau \sim 0.6$. The emission attains a maximum at $\tau = 0.8$ and then falls off. Figure 5 shows the spectrum of Bremsstrahlung X-ray radiation. The spectrum falls off exponentially at frequency $\hbar\omega \gg T_e$. Initially, X-ray emission rises slowly but rises resonantly when $\omega = \omega_{pe}/\sqrt{3}$. After a while, when cluster expansion detunes the plasma resonance, X-ray emission falls off. At any given instant of time the total X-ray emission comes from all the clusters at the different radial locations. The plasma resonance first occurs in clusters at laser axis. However, their number is small; hence enhancement in X-ray emission rate is small. As the plasma resonance layer $(\omega = \omega_{pe}/\sqrt{3})$ moves to high value of (r/r_0) , the number of resonantly heated electrons goes up and the net value of Bremsstrahlung X-ray emission rate is enhanced.

4. DISCUSSION

The laser heated clusters emit X-rays via Bremsstrahlung process. When a Gaussian laser beam interacts with clusters, laser quickly ionizes all the atoms of cluster and electrons are set free in the cluster. These free electrons of a cluster when pass through the vicinity of ions get accelerated and emit radiation in the X-ray regime. The electrons in the cluster acquire energy and oscillate in the laser field. Due to electron-ion collision the cluster gets heated and begins to expand under hydrodynamic pressure. The heating rate of the cluster electrons is maximum on the axis and decreases away from the axis. As the clusters expand under

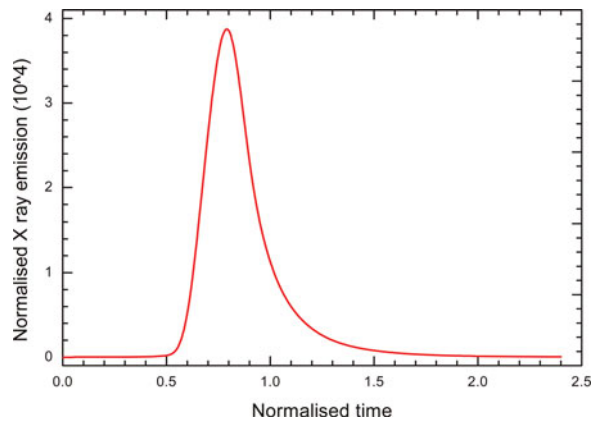


Fig. 4. (Color online) Normalized total X-ray emission by all clusters $(P_B^T / (m\omega_{pe0}^3 / n_{e0} c))$ as a function of normalized time $(\tau = v_0 t')$ for $a_0 = 1.0$ and $\hbar \omega/T_0 = 10$.

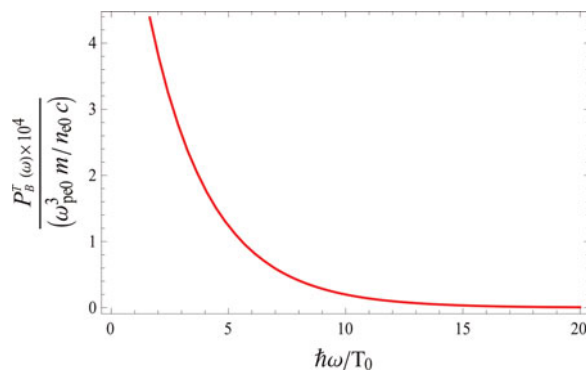


Fig. 5. (Color online) Bremsstrahlung soft X-ray spectrum of the emitted radiation.

hydrodynamic pressure, plasma frequency of the cluster electrons ω_{pe} decreases. As a consequence, clusters in the axial region expand more rapidly and approach plasma resonance $\omega_{pe} = \sqrt{3}\omega$ quickly, whereas the clusters at the larger r/r_0 attain the resonance at the later times. At any given instant of time the total X-ray emission comes to all the clusters at the different radial locations. As the plasma resonance layer move to high values of r/r_0 , the number of resonantly heated electrons goes up and the net values of X-ray emission rate is enhanced. When the plasma frequency of cluster electrons falls down below the plasma resonance, the emission rate slows down due to reduction in the rate of Bremsstrahlung encounters. The X-ray emission peak has a finite temporal width, of the order of the characteristic cluster expansion time $(r_{c0}/v_{s0})(n_e/n_{cr})^{1/3}$. The X-ray emission peak thus turns out to be after order of picosecond, which is comparable to the time width of the peak of X-ray emission, reported by Ditmire *et al.* for the Argon and Xenon clusters. For the high cluster densities and longer plasma length absorption may become important.

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